

## Infrared Emission from UAP

Figure 8 on page 22 of the attached paper implies that the surface temperature of the Space Shuttle reaches a value of  $T_s \sim 1200\text{K}$  for a Mach number,  $M \sim 25$ , and an airplane wing reaches  $T_s \sim 500\text{K}$  at  $M \sim 3$ .

The resulting infrared luminosity of the surface of a UAP of radius  $r$ , moving at  $M \sim 3$ , would be:

$$L_{\text{IR}} \sim \sigma T_s^4 \times \pi r^2 = 10^{11} \frac{\text{erg}}{\text{s}} \times \left(\frac{r}{1\text{m}}\right)^2 \quad \text{for } T_s \sim 500\text{K}$$

Where  $\sigma = 5.7 \times 10^{-5} \text{ erg/cm}^2 \cdot \text{s} \cdot \text{deg}^4$  is the Stefan-Boltzmann constant.

For a telescope of diameter,  $d$ , at a distance,  $D$ , this gives a photon count rate of,

$$\frac{dN_{\text{IR}}}{dt} = \left(\frac{L_{\text{IR}}}{3kT_s}\right) \times \left(\frac{\frac{\pi}{4}d^2}{4\pi D^2}\right) = 3 \times 10^{14} \frac{\text{photons}}{\text{s}} \left(\frac{r}{1\text{m}}\right)^2 \left(\frac{d}{1\text{m}}\right)^2 \left(\frac{D}{10\text{km}}\right)^{-2}$$

at a blackbody wavelength of,

$$\lambda_{\text{IR}} \sim 3\mu\text{m} \left(\frac{T_s}{500\text{K}}\right)^{-1} \quad \text{for the peak of } B_\lambda.$$

This flux is larger by an order of magnitude than the thermal emission rate off surfaces at room temperature and can easily be detected with IR sensors.