

Astronomy 202b: Cosmology

Tue./Thu., 9:30-11:00AM, Spring 2012

Syllabus

Course Instructors

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Course Requirements

Problem sets (due every second week): 30% of grade

Numerical project: 20% Final Exam: 50%

Course Texts

Required:

★ Loeb, A. & Furlanetto, S. 2012, The First Galaxies
(Princeton: Princeton Univ. Press)

★ Mo, H., Van den Bosch, F. & White, S.D.M. 2010, Galaxy Formation and Evolution
(Cambridge: Cambridge Univ. Press)

★ Padmanabhan, T. 1993, Structure Formation in the Universe
(Cambridge: Cambridge Univ. Press)

Recommended:

★ Schneider, P. 2006, Extragalactic Astronomy and Cosmology
(Berlin: Springer)

★ Mukhanov, V. 2005, Physical Foundations of Cosmology
(Cambridge: Cambridge Univ. Press)

Course Outline

1. Introduction	1/24, 1/26
(a) The Big Picture	
(b) Composition of the current universe	
(c) Scales in Astrophysics	
(d) Standard observables	
2. The Big Bang: Birth of our Universe	1/31, 2/2, 2/7
(a) The cosmological principle: isotropy and homogeneity	
(b) The Hubble expansion: the redshift-distance relation	
(c) Relics from the Big-Bang: the microwave background, synthesis of light elements	
(d) Geometry of the Universe: the expansion factor, lookback time	
(e) Cosmological parameters: the Hubble constant, the mean mass density	
3. Thermal History of the Universe	2/9, 2/14, 2/16
(a) Early times: Planck era, inflation, baryogenesis, electroweak and QCD phase transitions, neutrino decoupling	
(b) Nucleosynthesis	
(c) Ionization history of the universe	
(d) The Microwave background	
4. Structure Formation I: Linear Regime ...	2/21, 2/23, 2/28, 3/1, 3/6, 3/8, 3/20
(a) Friedmann equations, fluid equations	
(b) Linear theory, Jeans' instabilities	
(c) Cosmological density fields: stochastic fields, Gaussian fields, ergodic hypothesis, power spectra, correlation functions	
(d) Anisotropies in the Microwave background	
(e) Lyman- α forest	
(f) Reionization and 21cm cosmology	
5. Structure Formation II: Non-Linear Regime ...	3/22, 3/27, 3/29, 4/3, 4/5
(a) Spherical collapse, Press-Schechter, mass function	
(b) Dark Matter: evidence, detection, searches	

(c) Dark Matter halos: mass distribution and virial properties	
(d) Galaxy clusters: structure, evolution, cosmological probes	
(e) Galaxy formation: classification, mergers, starburst, quiescent accretion, physical models	
(f) Black holes: AGN, quasars, coevolution with galaxies	
6. Numerical Cosmology	4/10
(a) N-body methods	
(b) Hydrodynamical methods and state-of-the-art results	
Numerical projects	4/12, 4/17, 4/19