



DO FILAMENTS CROSS
"CORE" BOUNDARIES?

Alyssa Goodman & Hope Chen,
Harvard-Smithsonian Center for Astrophysics

+

Stella Offner, UMASS
Jaime Pineda, ETH Zurich & MPE

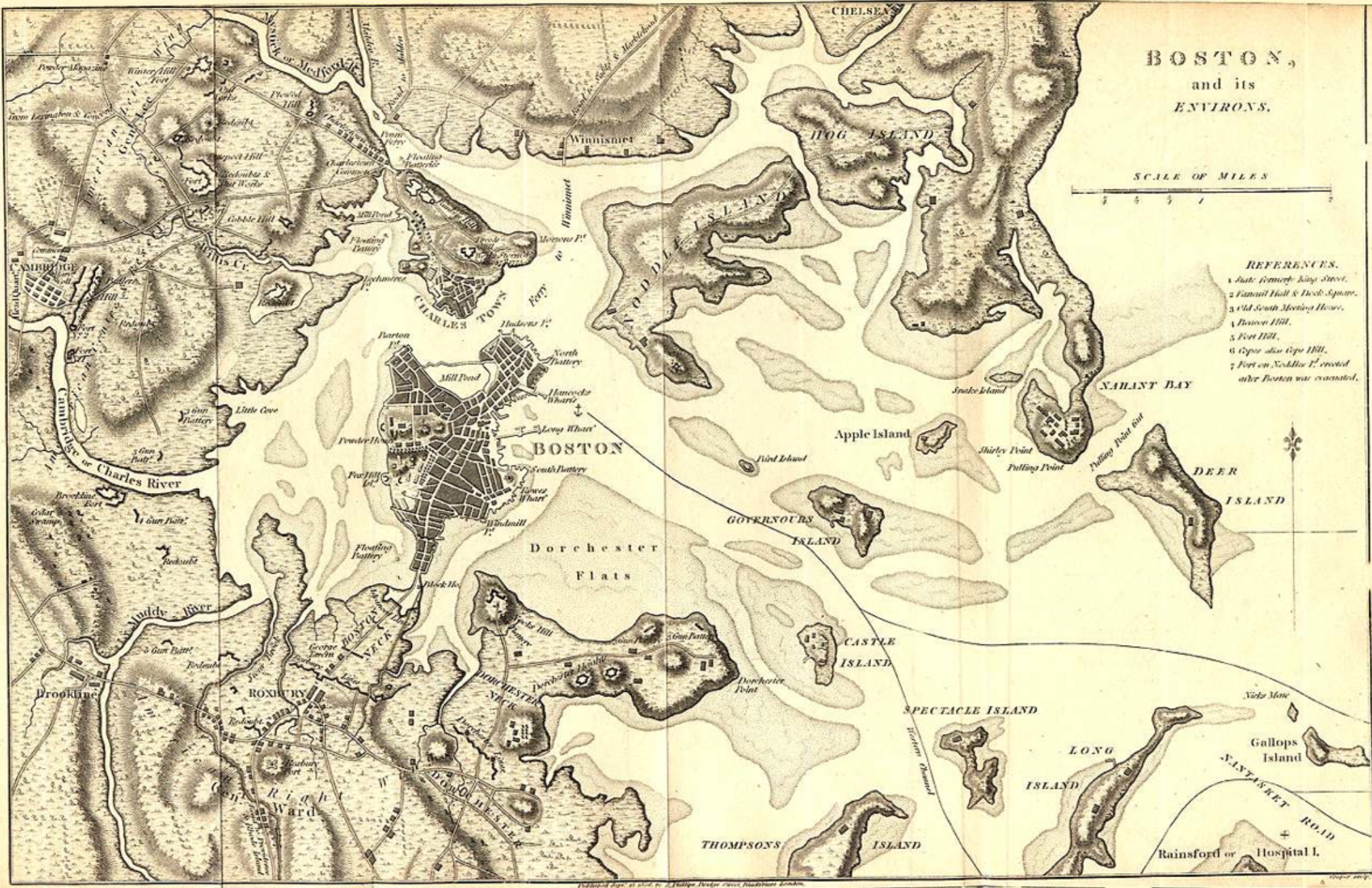
BOSTON, and its ENVIRONS.

SCALE OF MILES



REFERENCES.

- 1 State formerly King Street.
- 2 Faneuil Hall & Dock Square.
- 3 Old South Meeting House.
- 4 Beacon Hill.
- 5 Fort Hill.
- 6 Cape also Cape Hill.
- 7 Fort on No. 11th St. erected after Boston was evacuated.

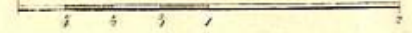


Published Sept. 23. 1806 by J. Phillips Deeds & Co. Newburyport London.

Boston 1806

BOSTON, and its ENVIRONS.

SCALE OF MILES



REFERENCES.

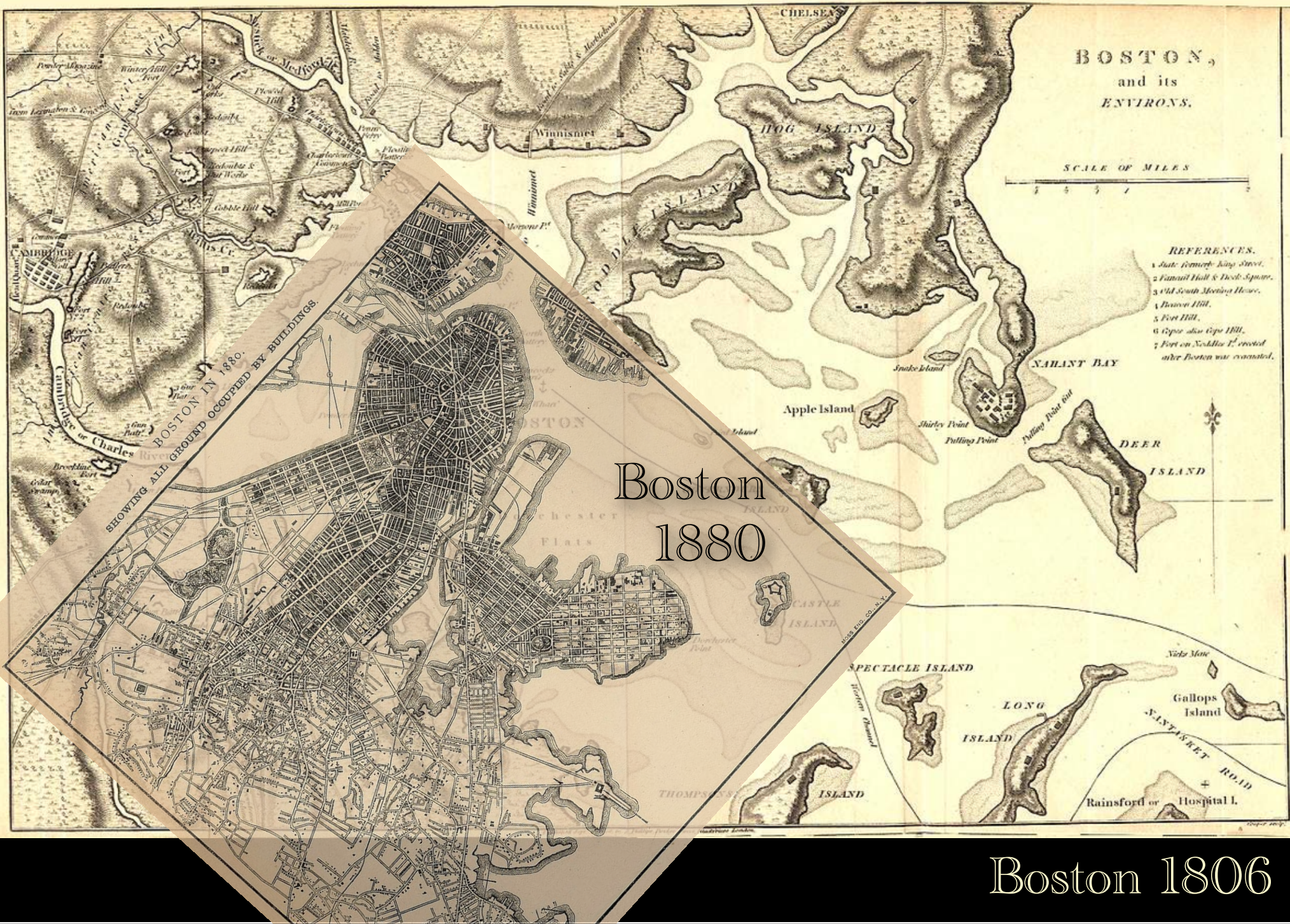
- 1 State formerly King Street.
- 2 Faneuil Hall & Dock Square.
- 3 Old South Meeting House.
- 4 Beacon Hill.
- 5 Fort Hill.
- 6 Cape also Cape Hill.
- 7 Fort on Noddin's I. erected after Boston was evacuated.

Boston 1880

SHOWING ALL GROUND OCCUPIED BY BUILDINGS.

BOSTON IN 1880.

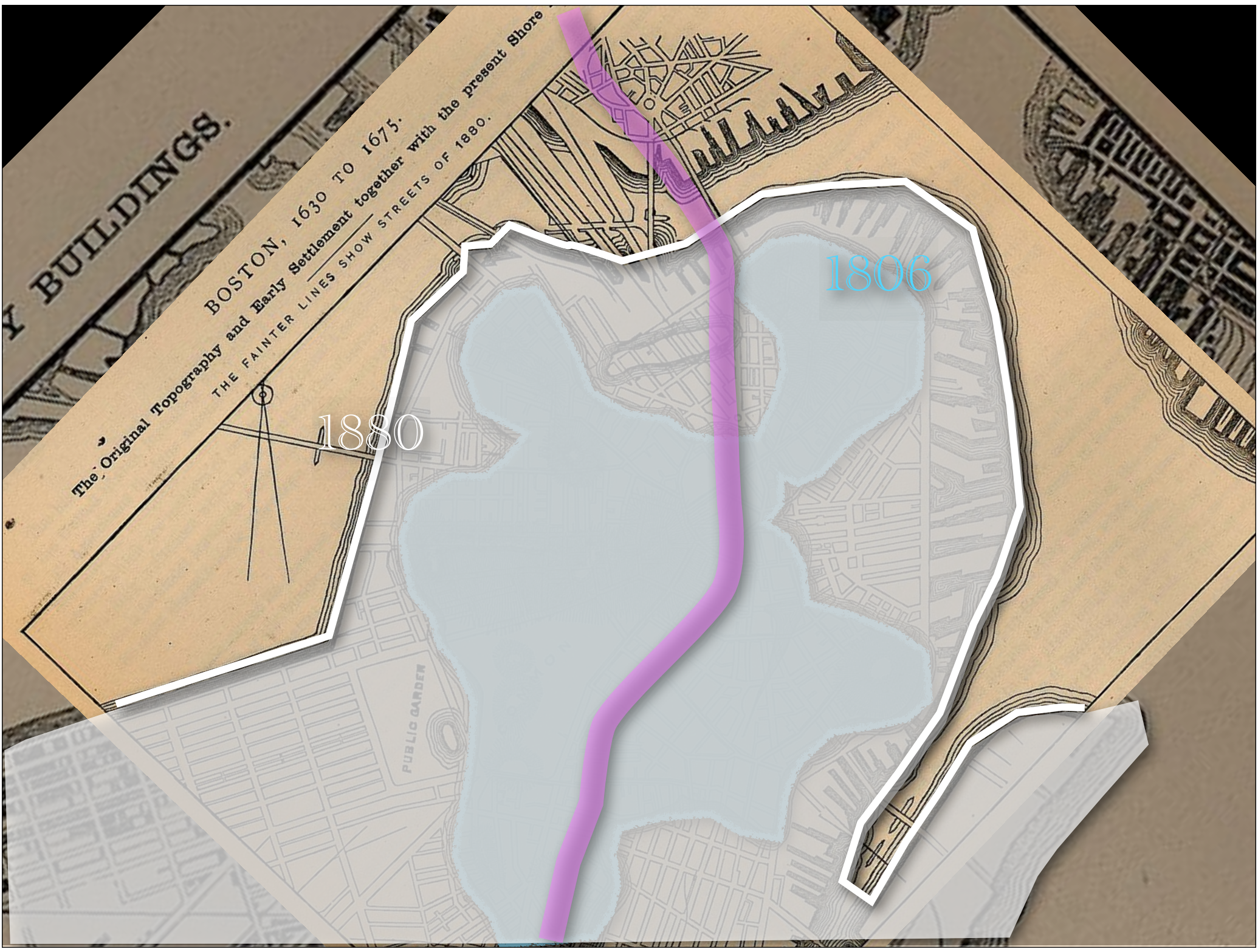
Boston 1806





Y BUILDINGS.

Boston
1880



THE ORIGINAL TOPOGRAPHY AND EARLY SETTLEMENT TOGETHER WITH THE PRESENT SHORE
THE FAINTER LINES SHOW STREETS OF 1880.

BOSTON, 1630 TO 1675.

1880

1806

PUBLIC GARDEN

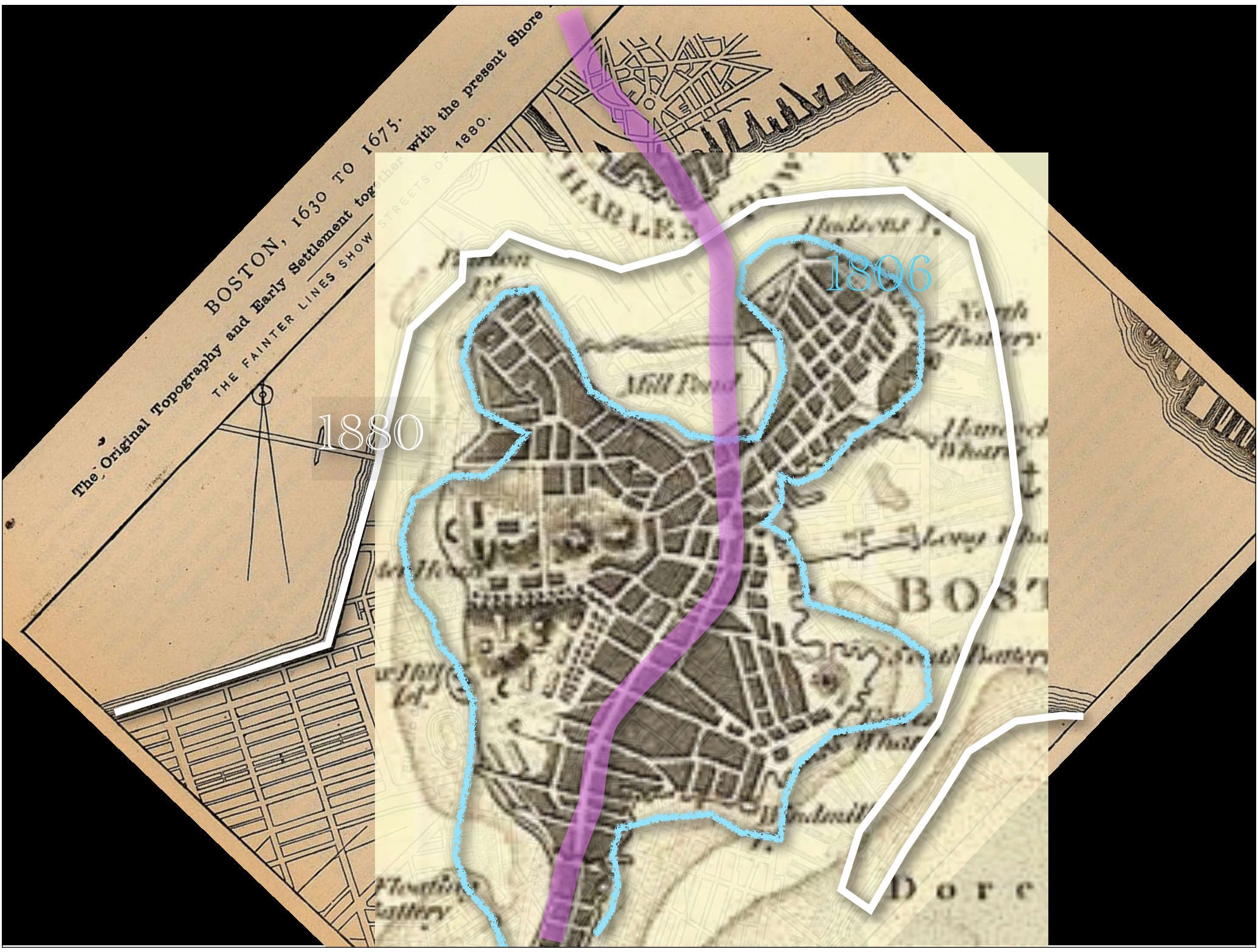
BOSTON, 1630 TO 1675.

The Original Topography and Early Settlement together with the present Shore
THE FAINTER LINES SHOW

with the present Shore
1880.

1880

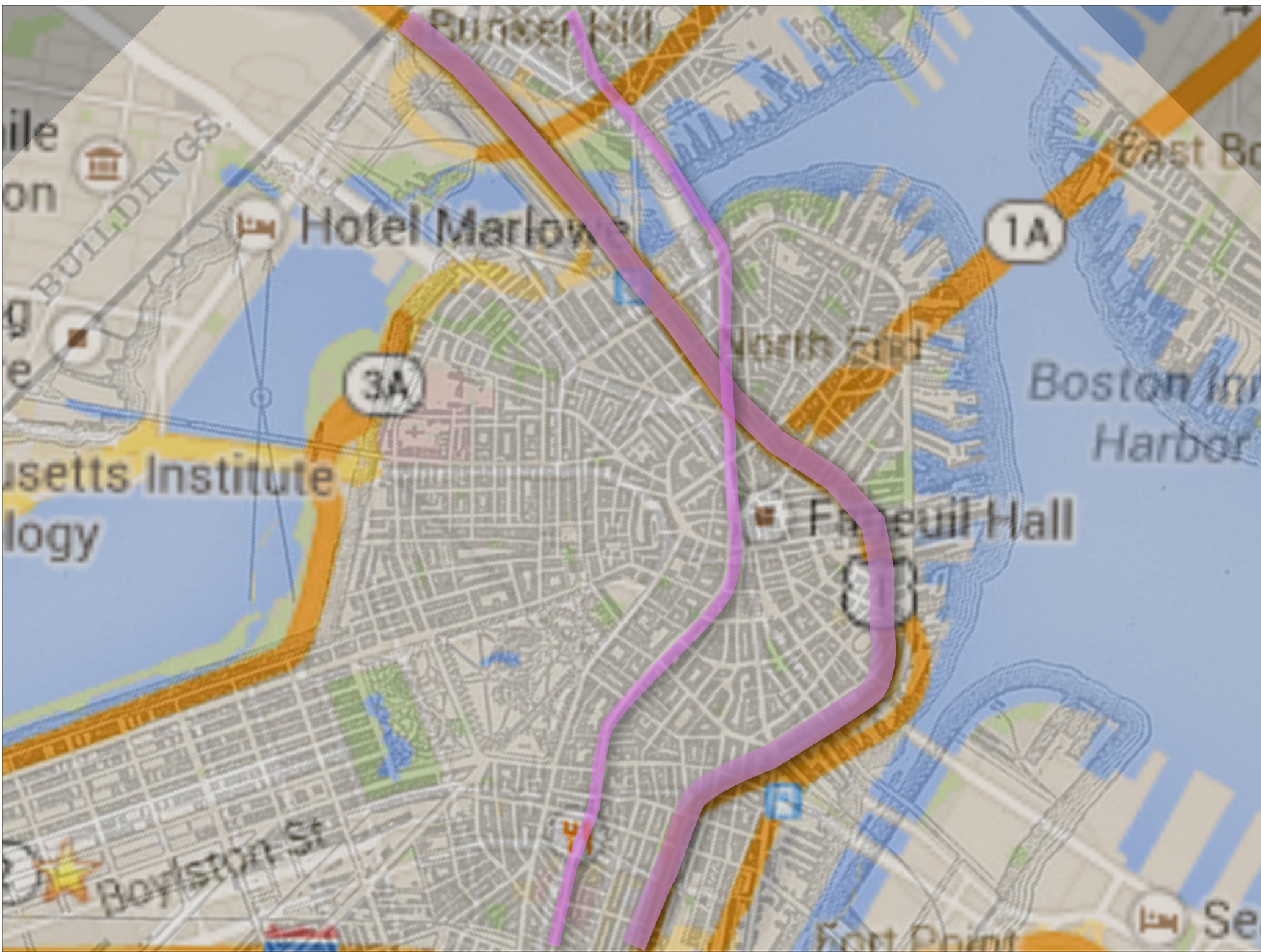
1806

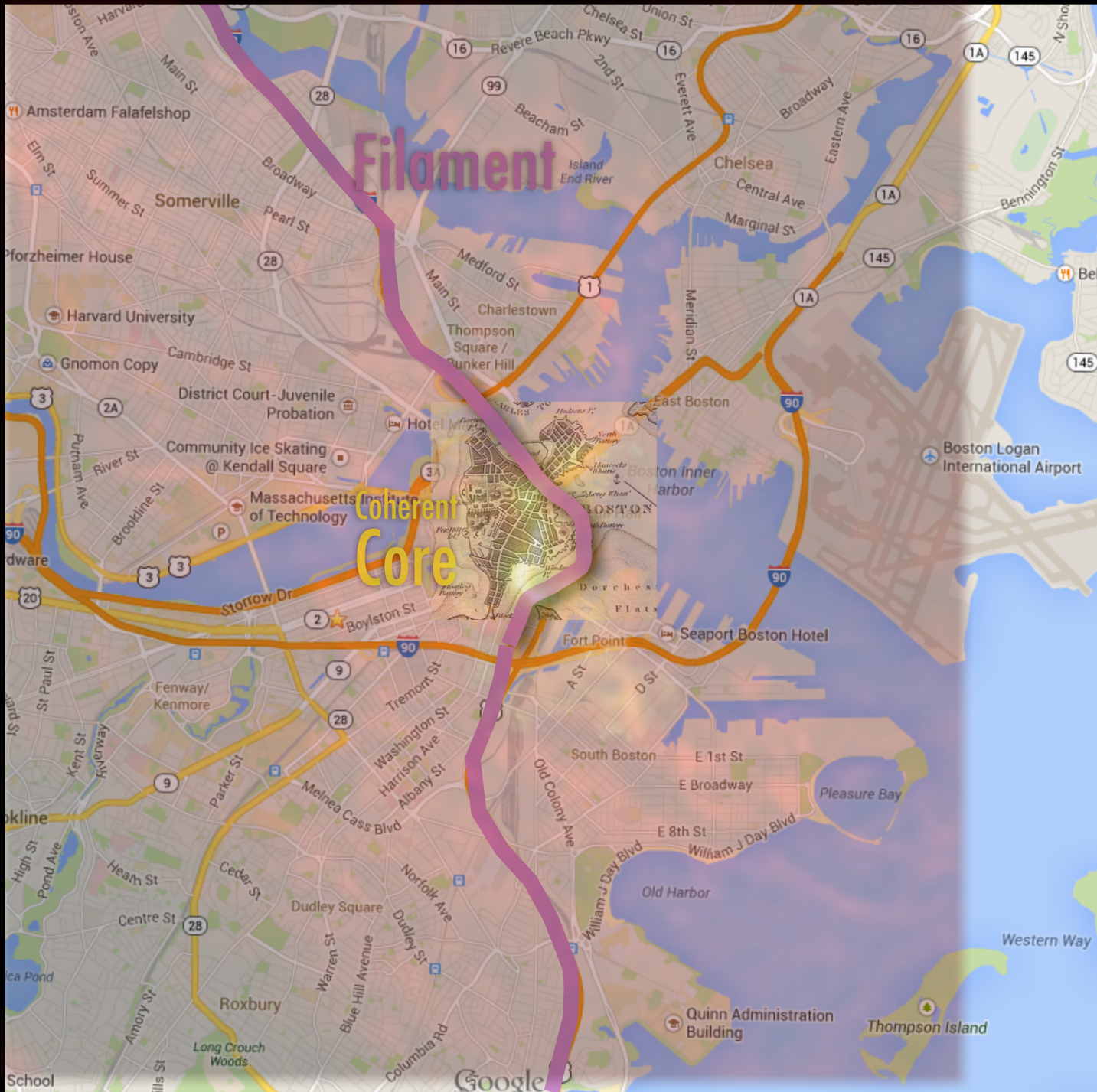


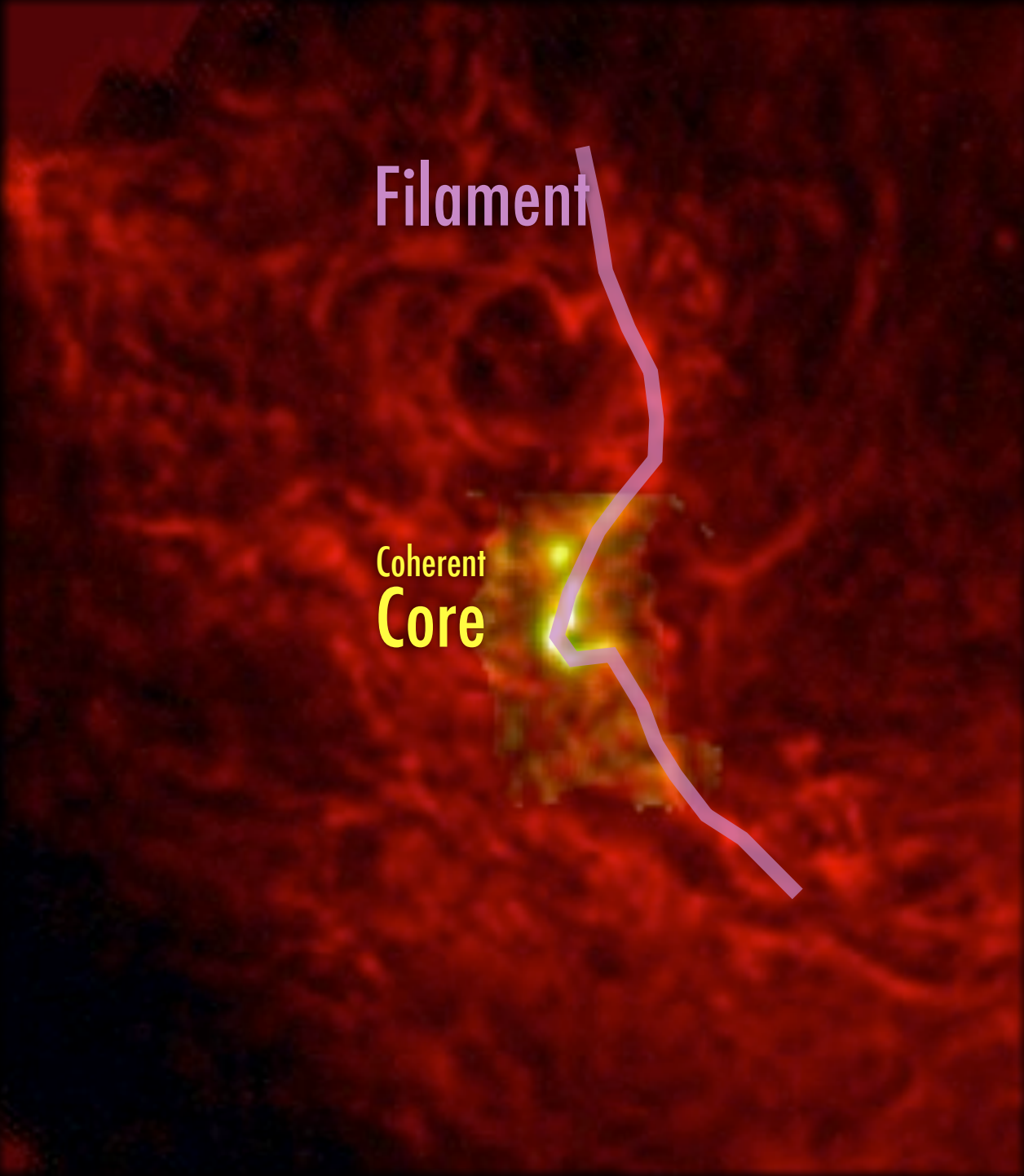


Y BUILDINGS.





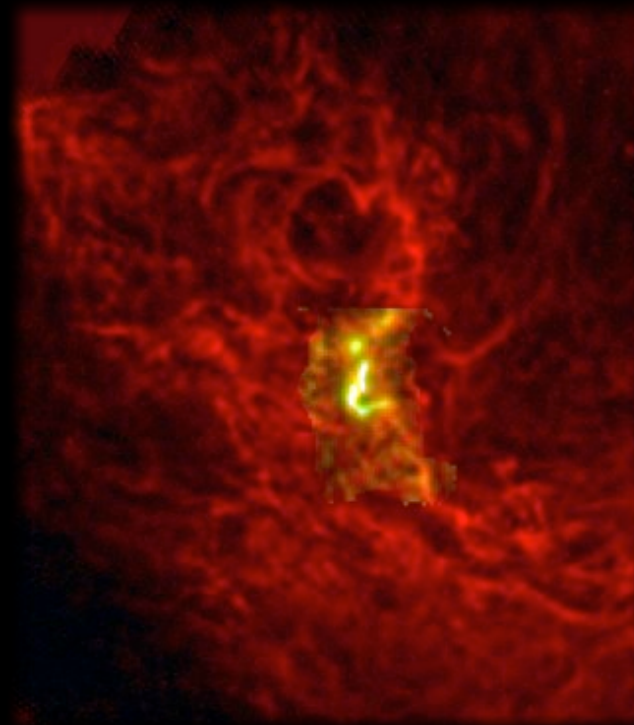




Filament

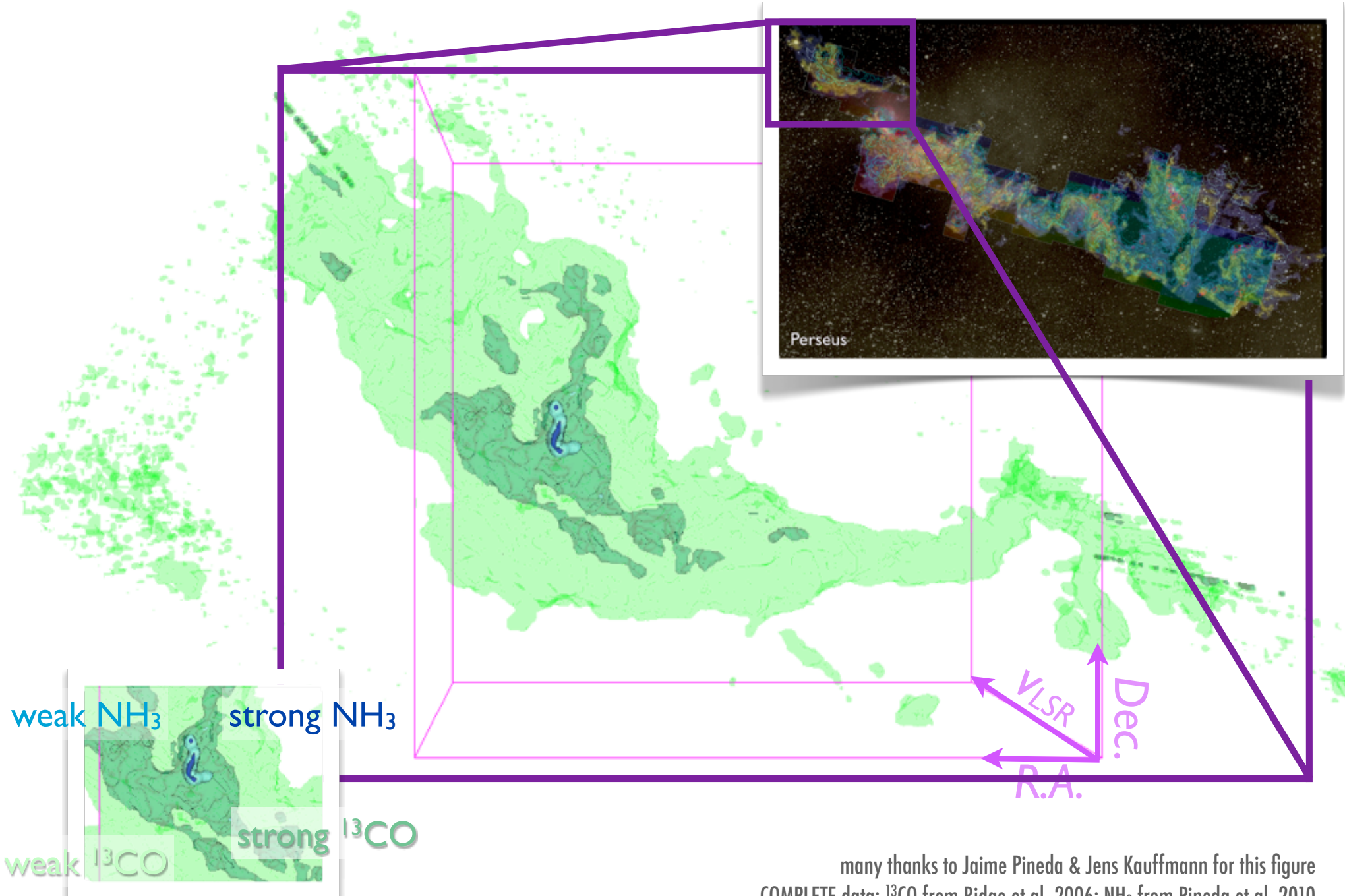
Coherent
Core

COHERENT CORES ISLANDS OF CALM IN TURBULENT SEAS(?)



30-year story: Myers & Benson 1983, Goodman et al. 1998, Pineda et al. 2010, 2011, 2014

THE B5 REGION, IN PERSEUS



many thanks to Jaime Pineda & Jens Kauffmann for this figure
COMPLETE data: ^{13}CO from Ridge et al. 2006; NH_3 from Pineda et al. 2010

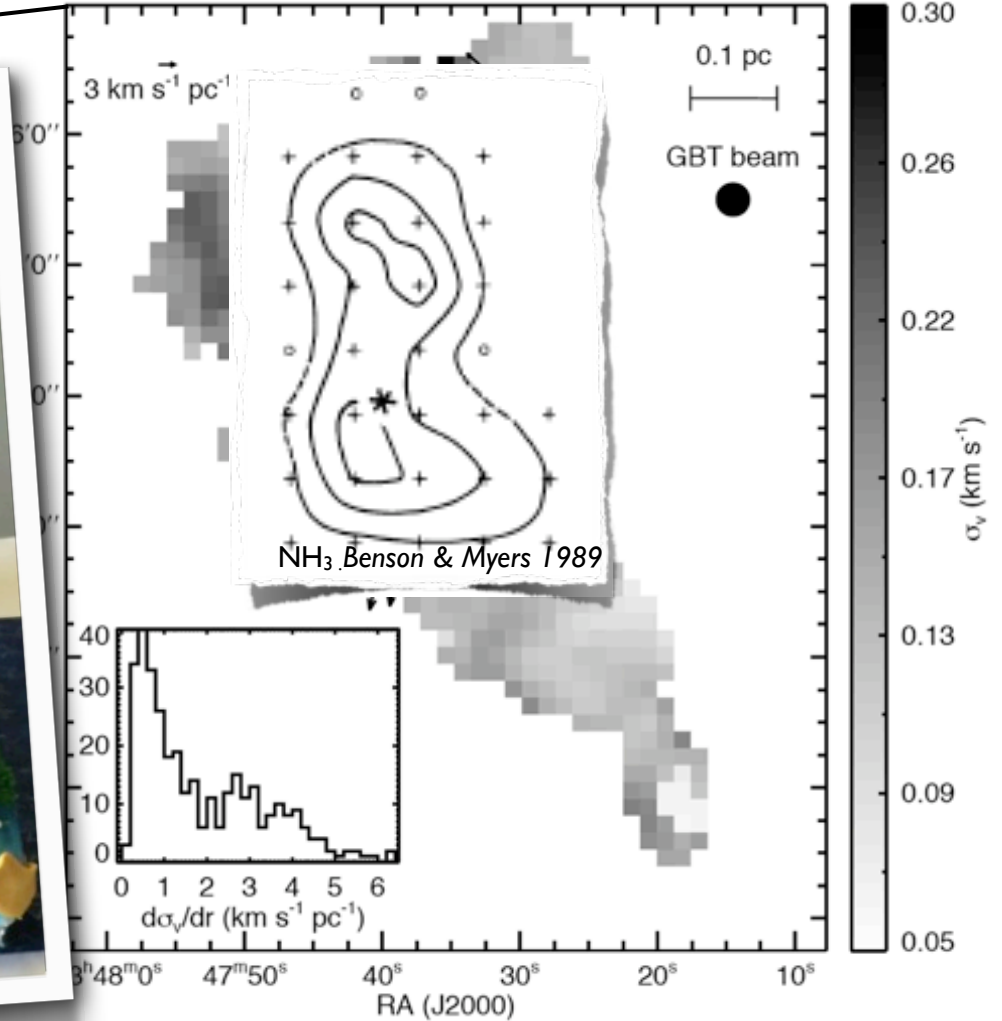
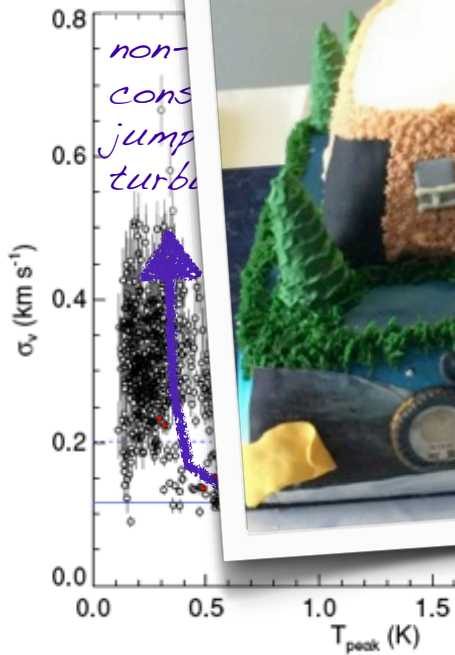
STRONG EVIDENCE FOR "VELOCITY COHERENCE" IN DENSE CORES

greyscale shows NH_3 velocity dispersion, arrows show gradient in dispersion

weak NH_3 strong NH_3

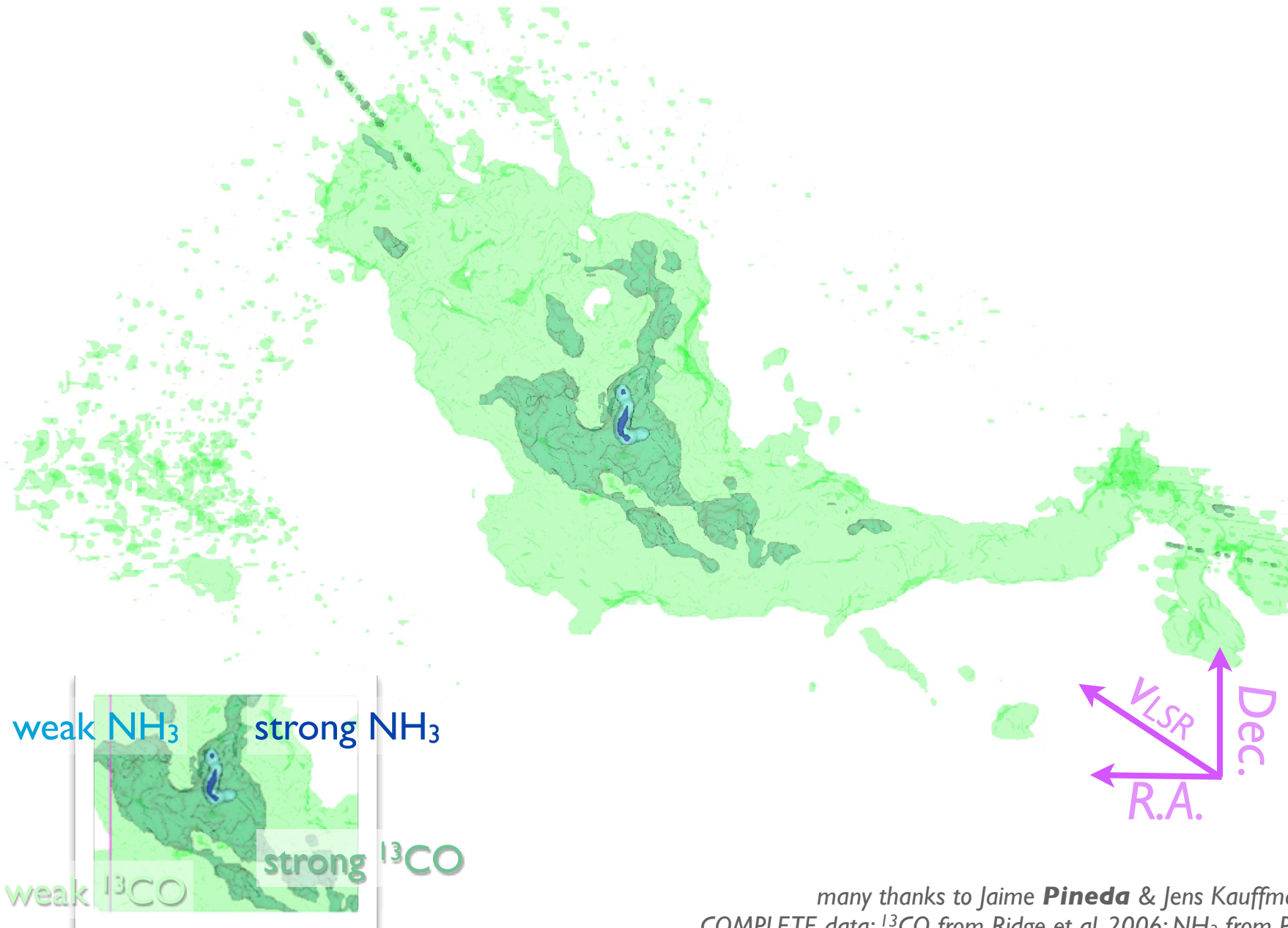
weak NH_3

weak NH_3



GBT NH_3 observations of the B5 core (Pineda et al. 2010)

POSITION-VELOCITY STRUCTURE OF THE B5 REGION IN PERSEUS



many thanks to Jaime **Pineda** & Jens Kauffmann for this figure
COMPLETE data: ¹³CO from Ridge et al. 2006; NH₃ from Pineda et al. 2010

BUT THEN... VLA (JAIME) FOUND SUB-STRUCTURE

THE ASTROPHYSICAL JOURNAL LETTERS, 739:L2 (5pp), 2011 September 20

PINEDA ET AL.

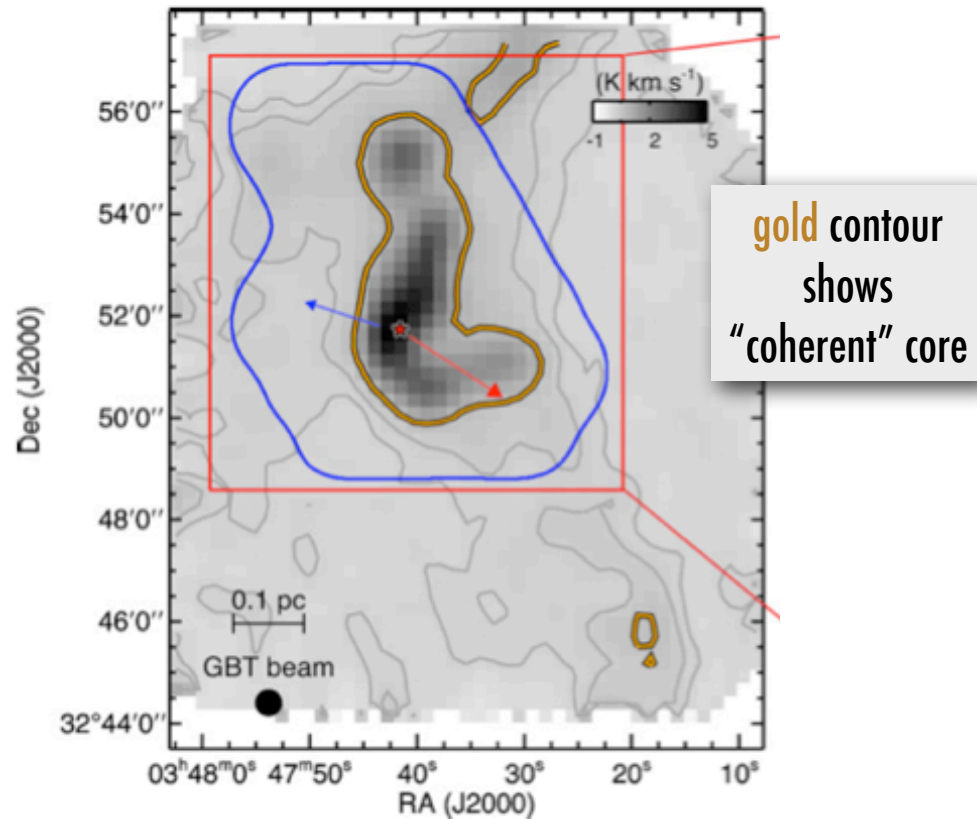
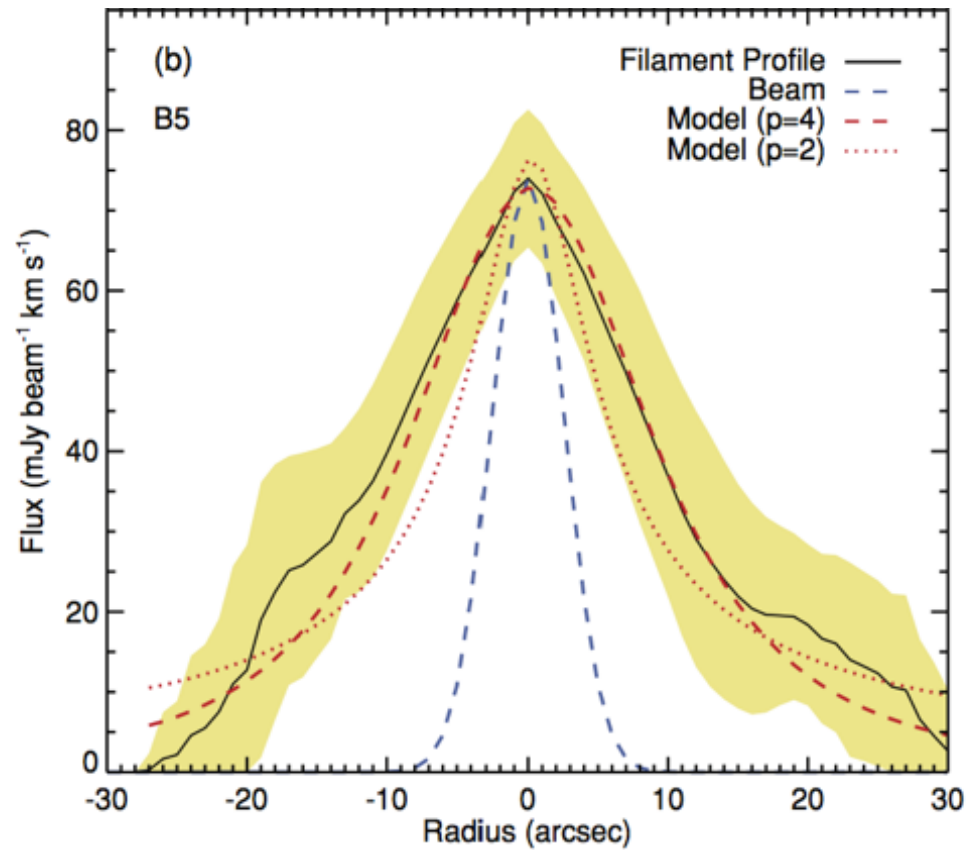


Figure 1. Left panel: integrated intensity map of B5 in NH_3 (1,1) obtained with GBT. Gray contours show the 0.15 and 0.3 K km s^{-1} level in NH_3 (1,1) integrated intensity. The orange contours show the region in the GBT data where the non-thermal velocity dispersion is subsonic. The young star, B5-IRS1, is shown by the star in both panels. The outflow direction is shown by the arrows. The blue contour shows the area observed with the EVLA and the red box shows the area shown in the right panel. Right panel: integrated intensity map of B5 in NH_3 (1,1) obtained combining the EVLA and GBT data. Black contour shows the 50 $\text{mJy beam}^{-1} \text{ km s}^{-1}$ level in NH_3 (1,1) integrated intensity. The yellow box shows the region used in Figure 4. The northern starless condensation is shown by the dashed circle.

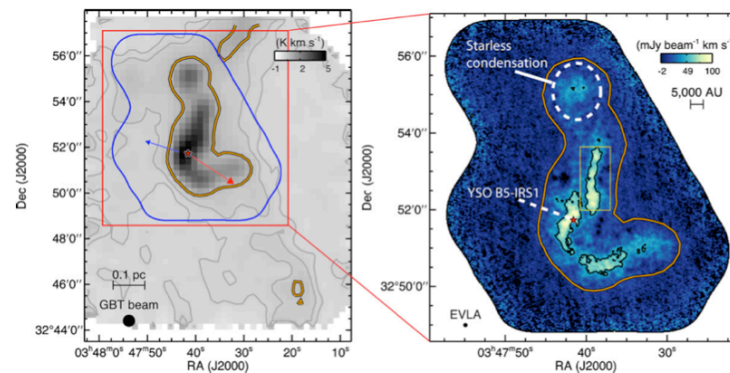
BUT MAYBE IT'S DIFFERENT?



isothermal,
hydrostatic filaments,
not turbulent ones?

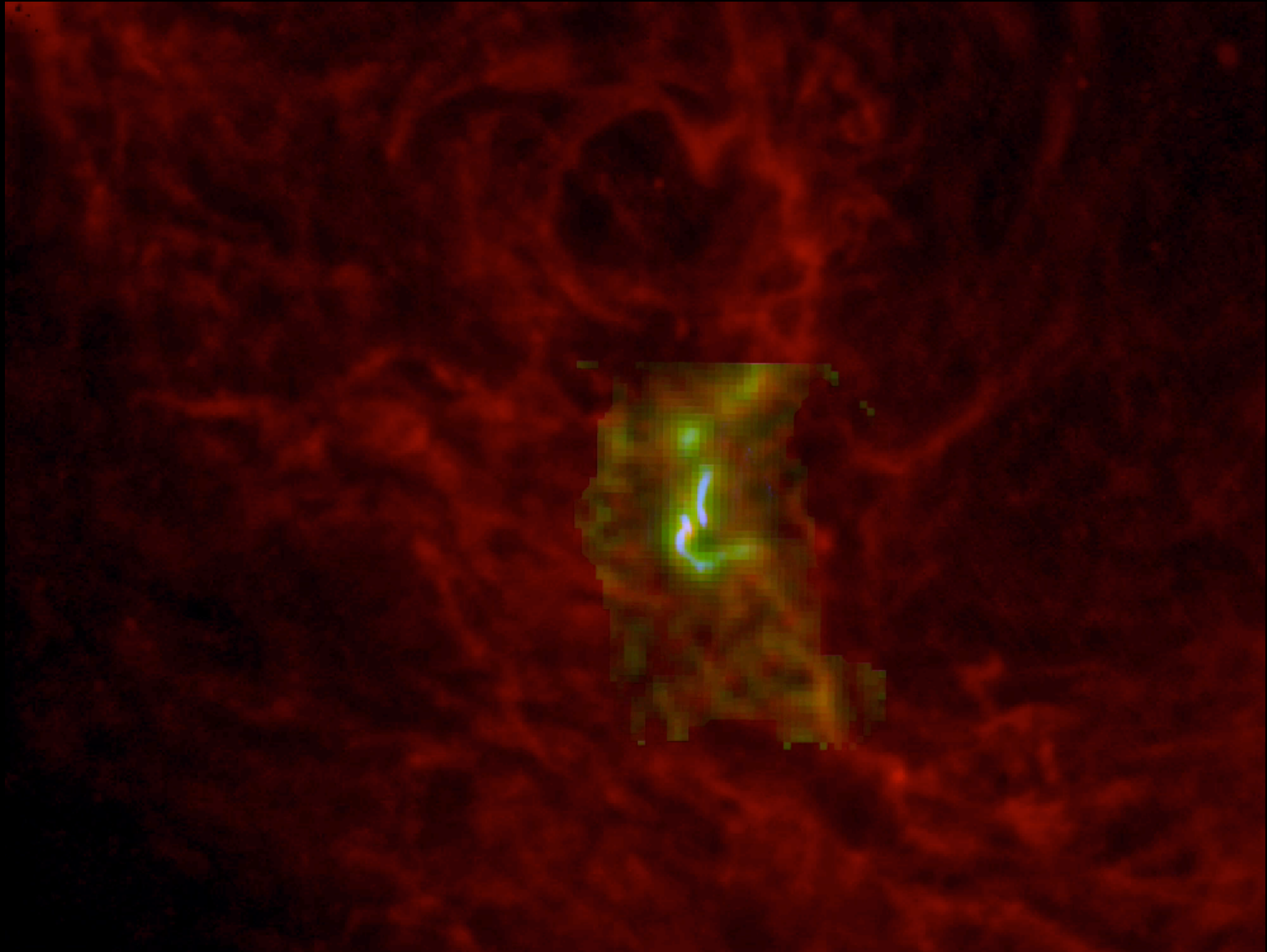
THE ASTROPHYSICAL JOURNAL LETTERS, 739:L2 (5pp), 2011 September 20

PINEDA ET AL.



BUT WHAT IF FILAMENTS CONTINUE ACROSS "CORE" BOUNDARIES?!

blue =VLA ammonia (high-density gas); green=GBT ammonia (lower-res high-density gas); red=Herschel 250 micron continuum (dust)





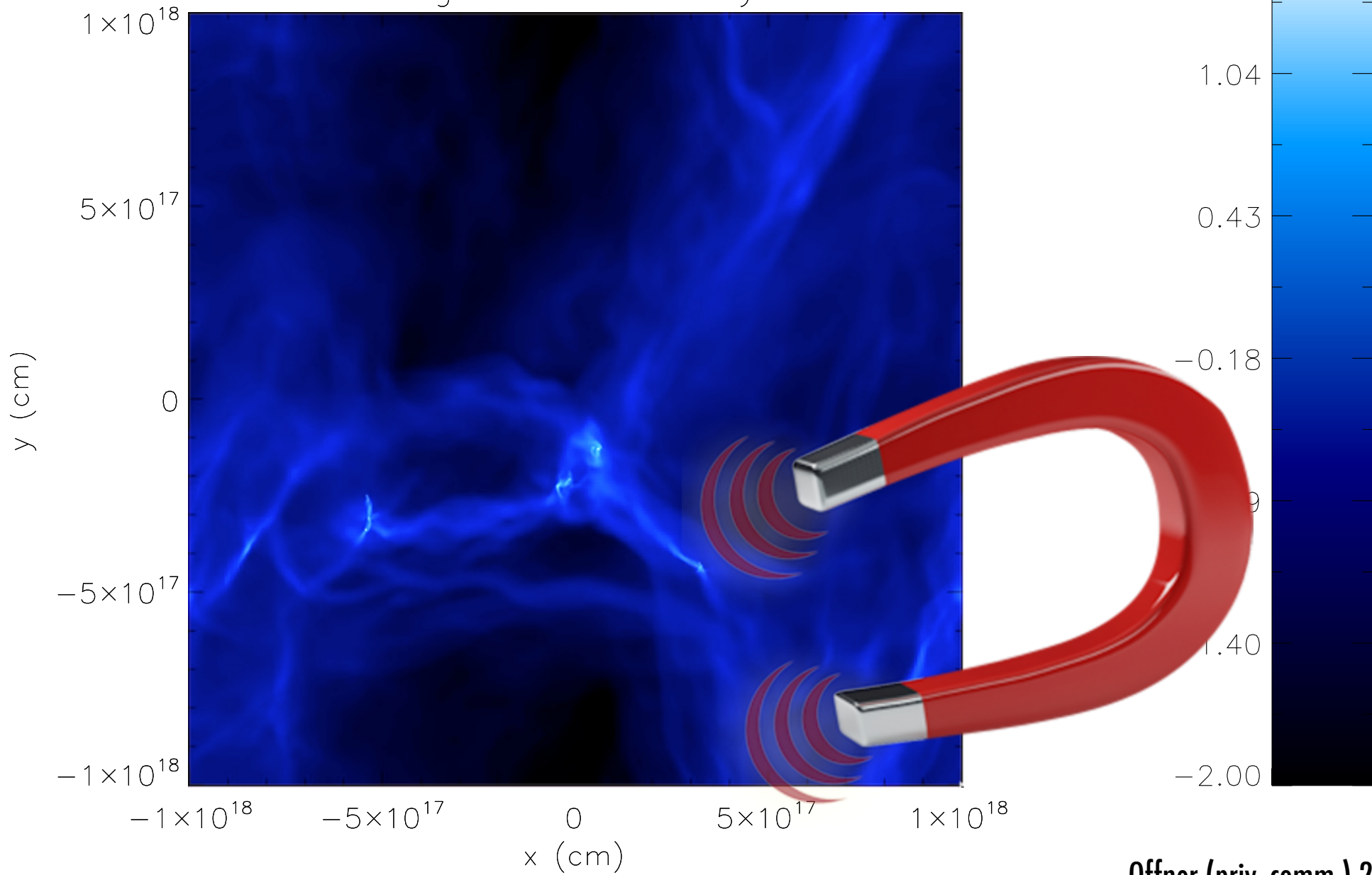
1998



2008

B5-ISH SIMULATION (NO MAGNETIC FIELD)

Log Column Density

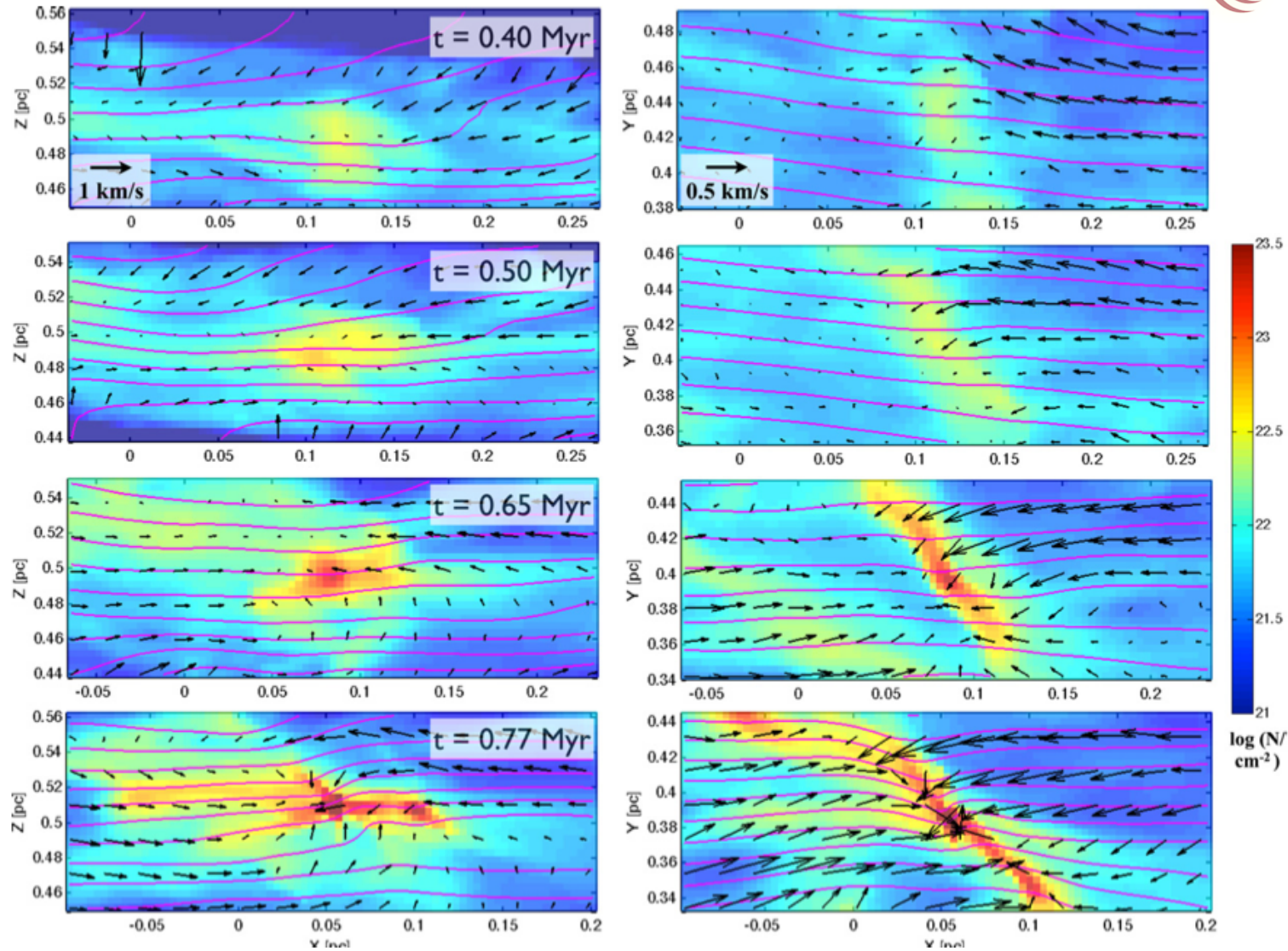


Offner (priv. comm.) 2014

B5-ISH? SIMULATION (WITH MAGNETIC FIELD)

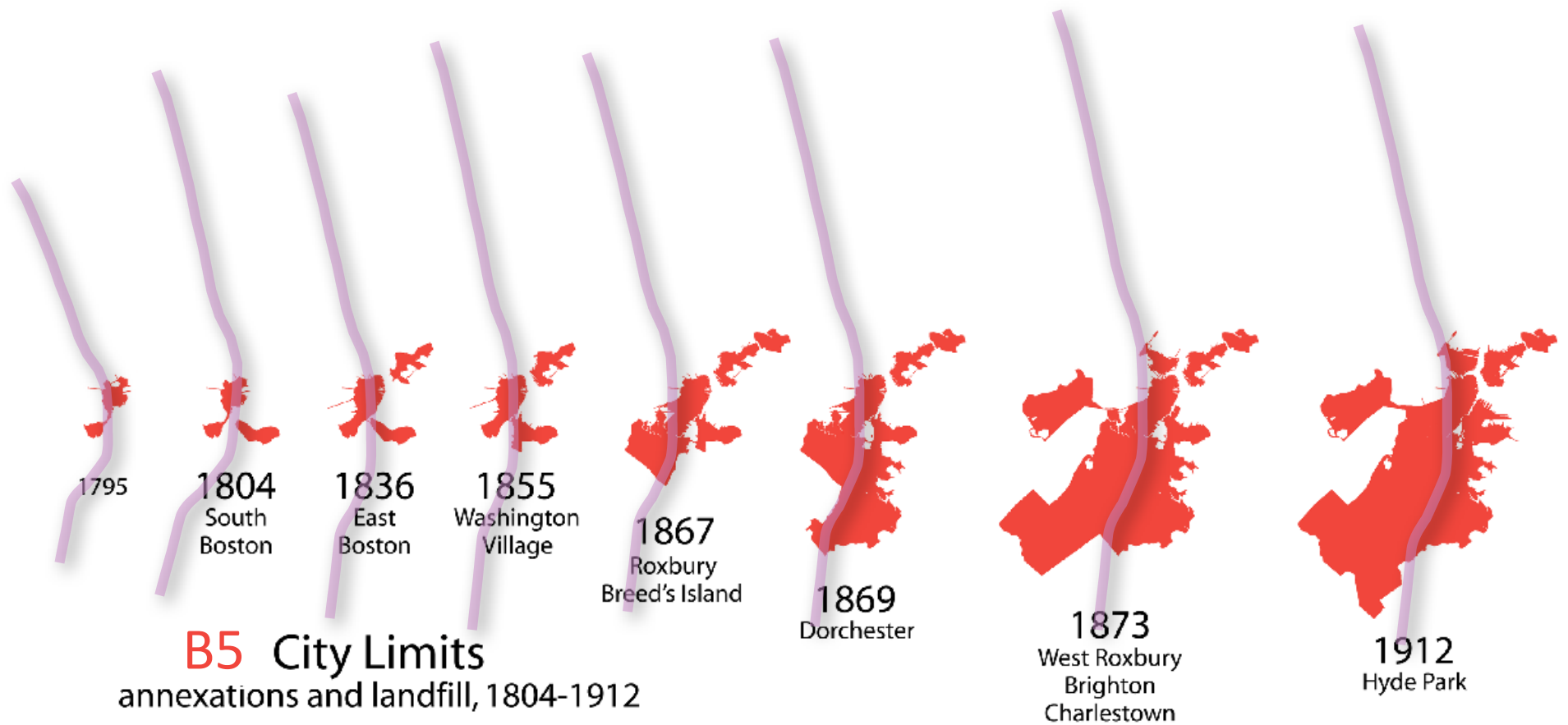


THE ASTROPHYSICAL JOURNAL, 785:69 (20pp), 2014 April 10



Chen & Ostriker 2014

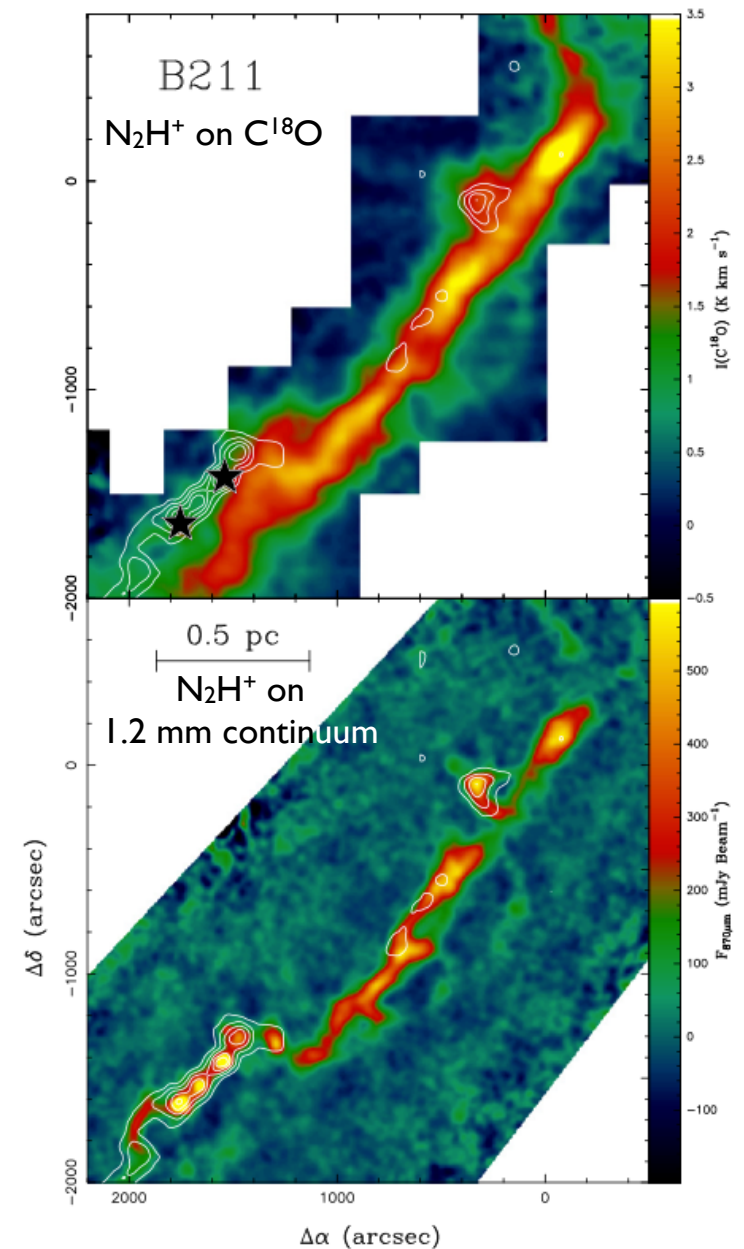
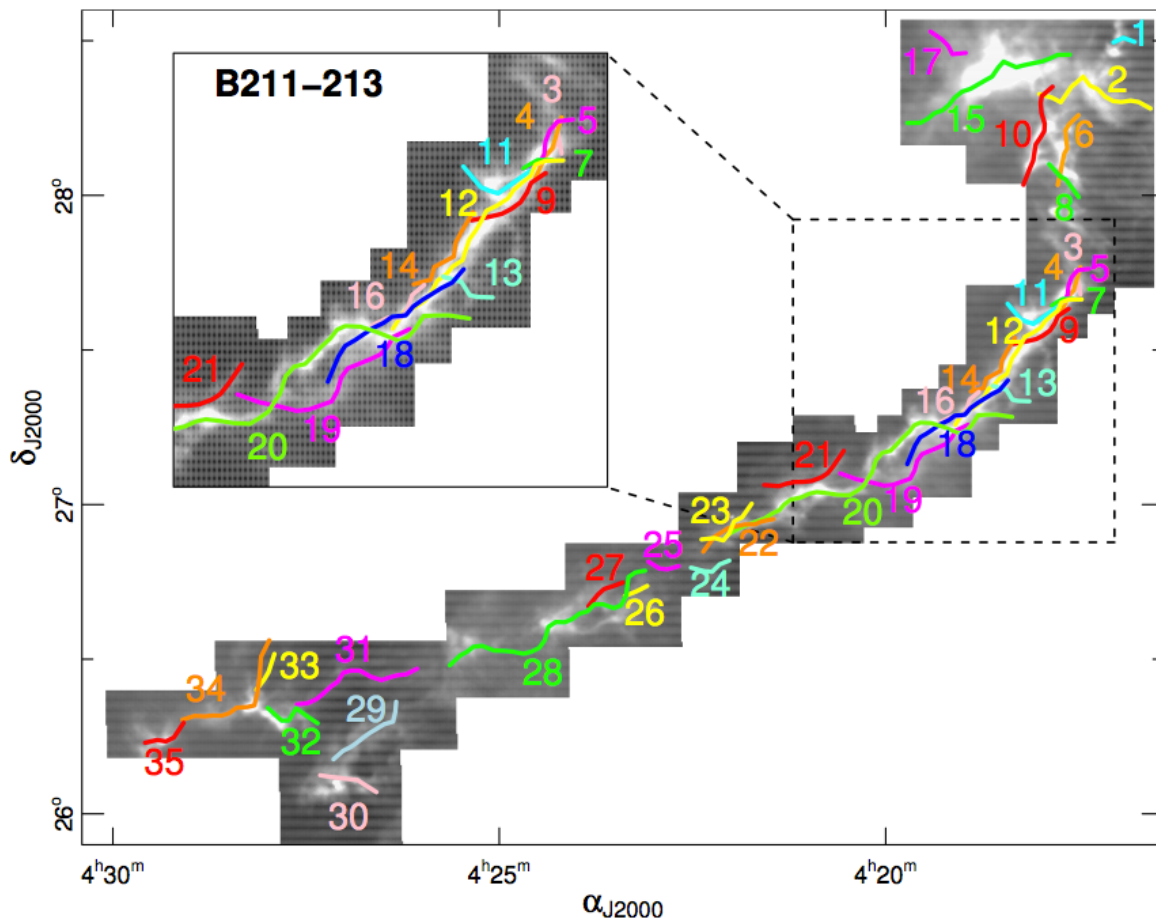
SHHH...



Filaments offer pre-existing density enhancement.

Collapse is rapid enough that aboriginal filament is not erased, even within a "coherent core."

In B5, small bound cluster will form c. 40K years from now.



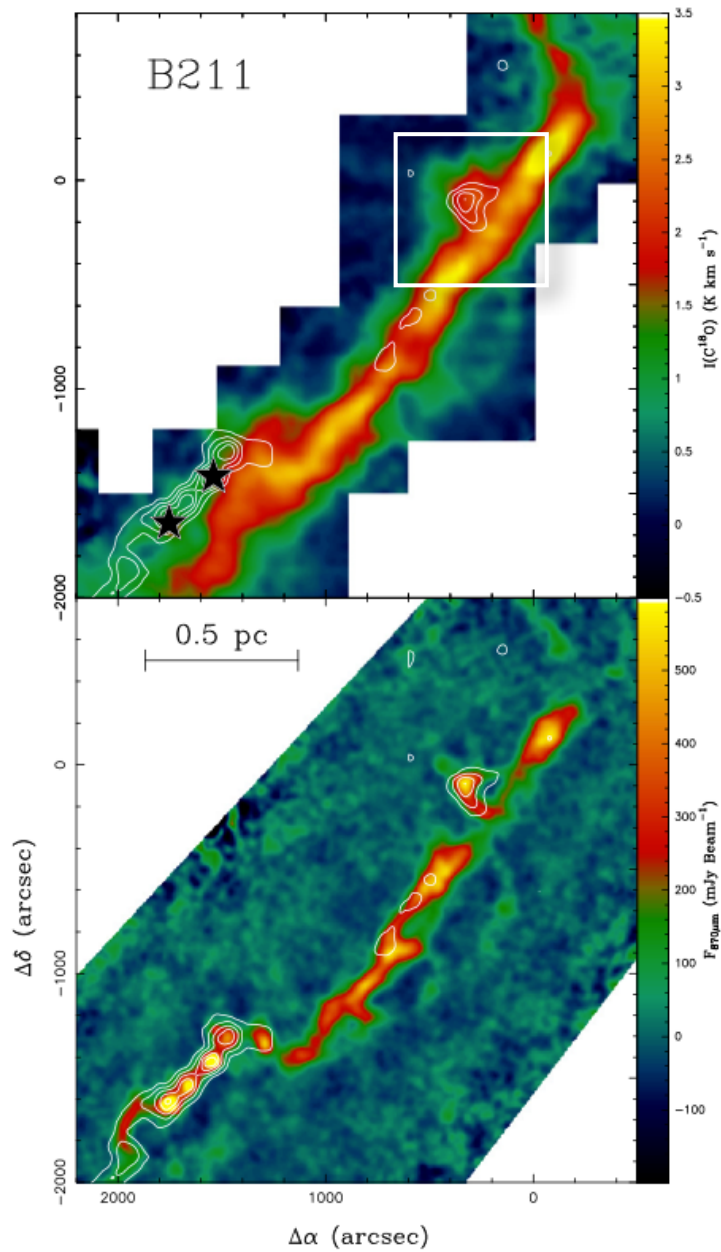
Now, we (all!) need to try FIVE, from Hacar et al. 2013, to study “coherent” core-filament relation.

Filaments offer pre-existing density enhancement.

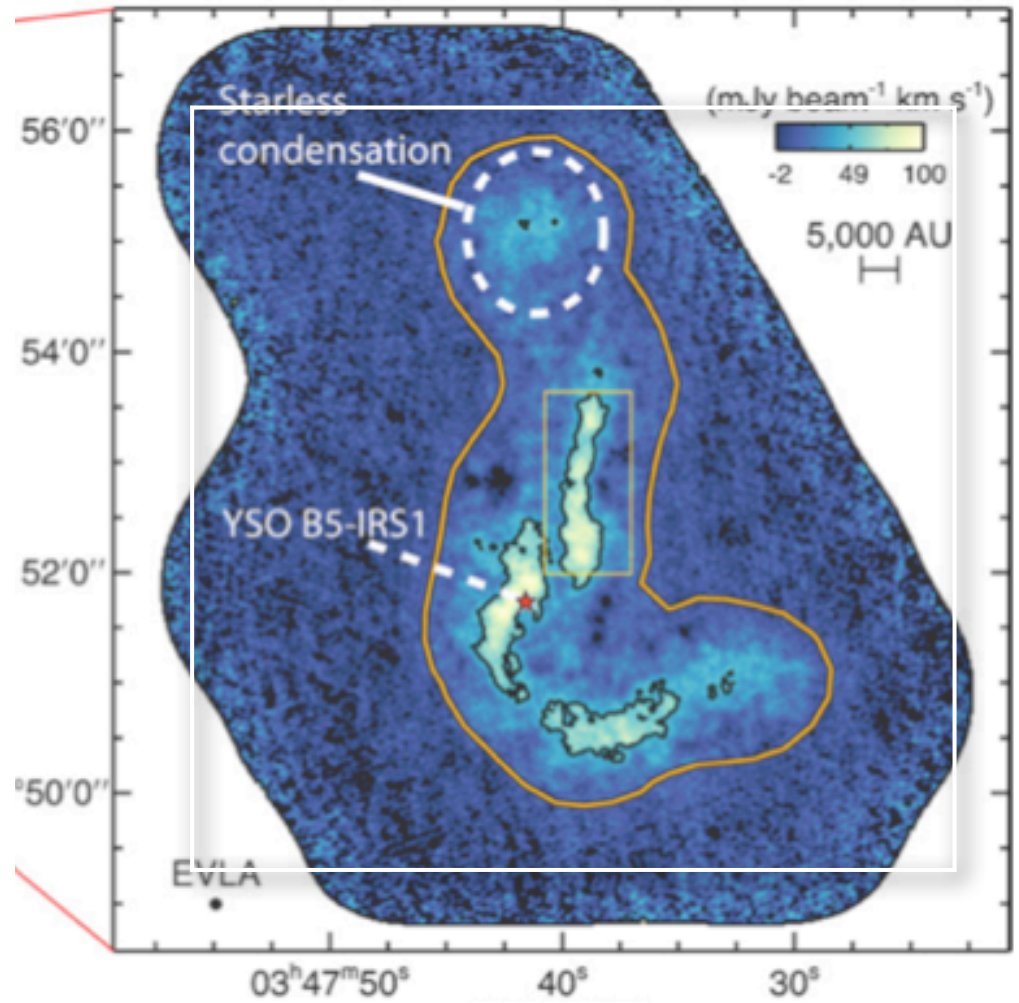
Collapse is rapid enough that aboriginal filament is not erased, even within a “coherent core.”

In B5, small bound cluster will form c. 40K years from now.

COMPARING SCALES



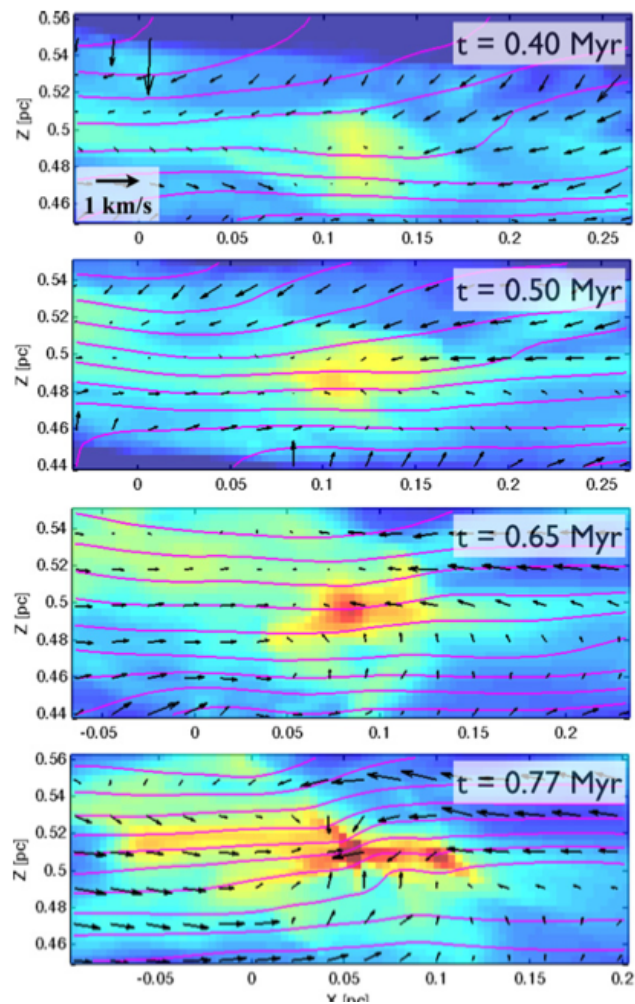
Taurus (Hacar et al.)



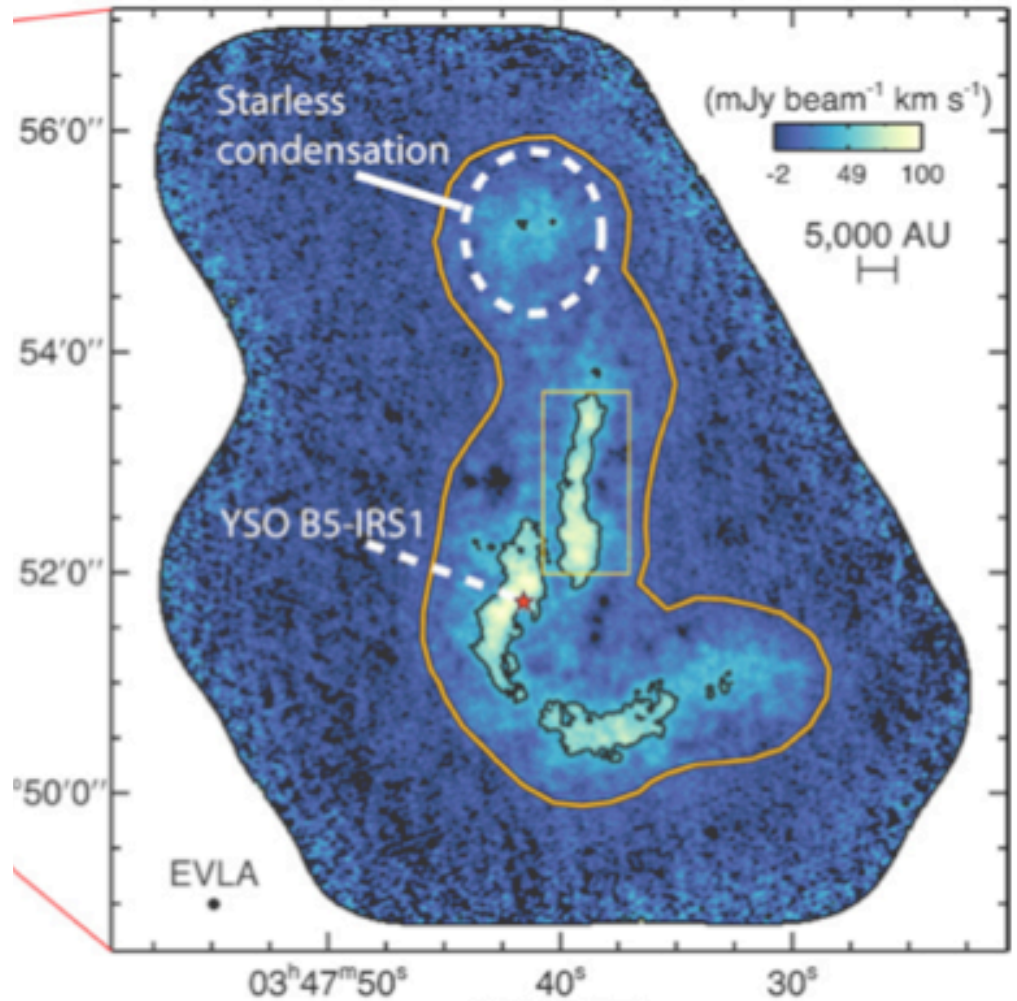
B5 (Pineda et al.)

TO THE SAME SCALE...

THE ASTROPHYSICAL JOURNAL, 785:69 (20pp), 2014 April 10



MHD (Chen & Ostriker 2014)



B5 (Pineda et al.)



Nessie to B5, the movie.