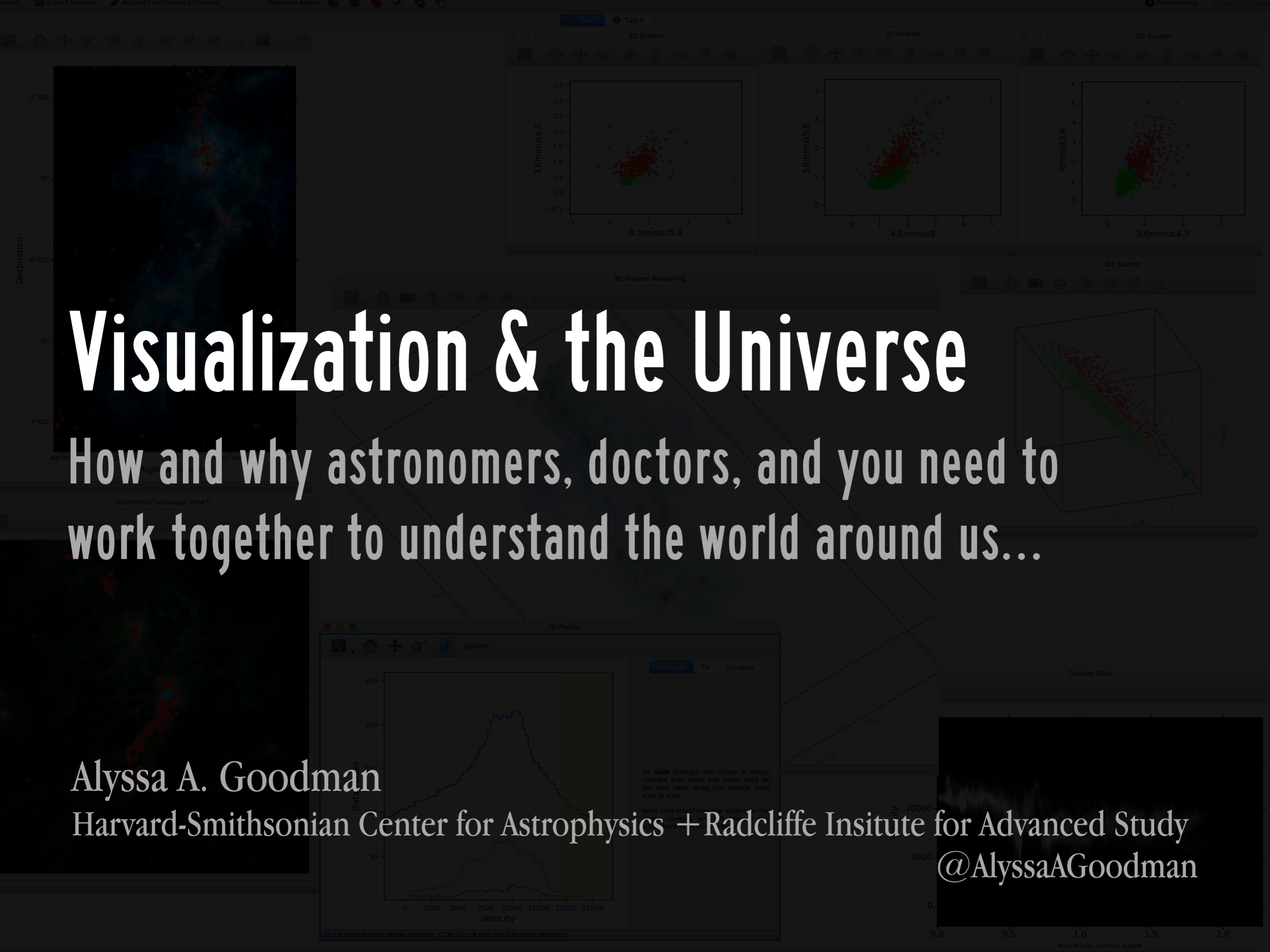


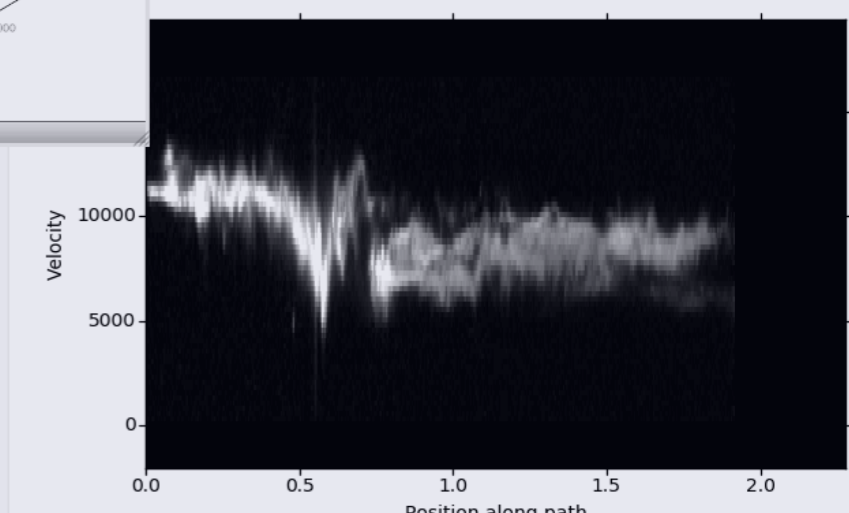
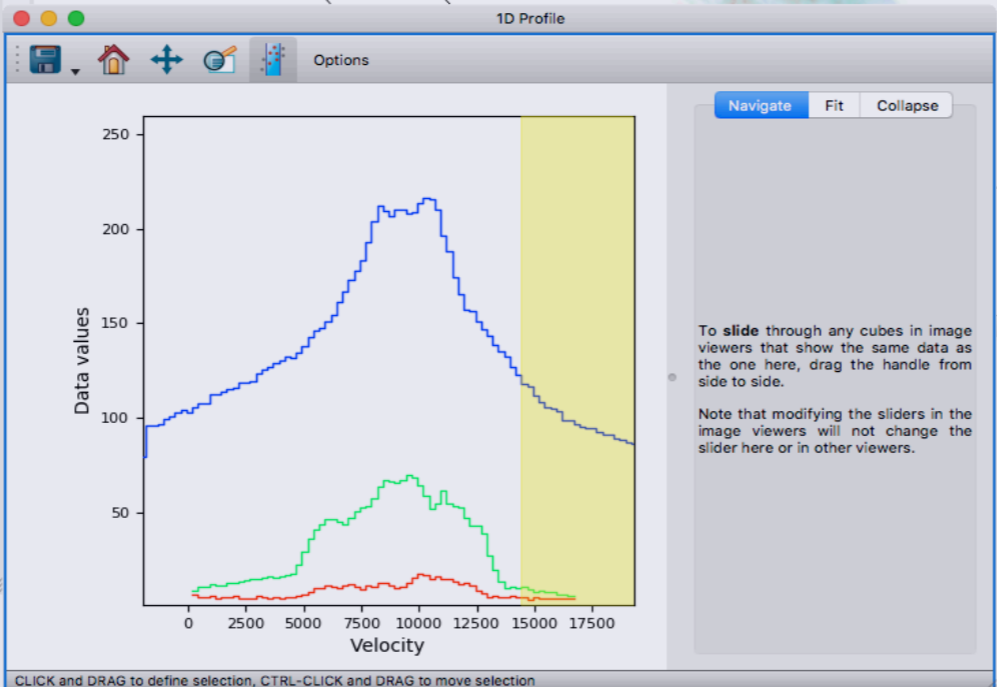
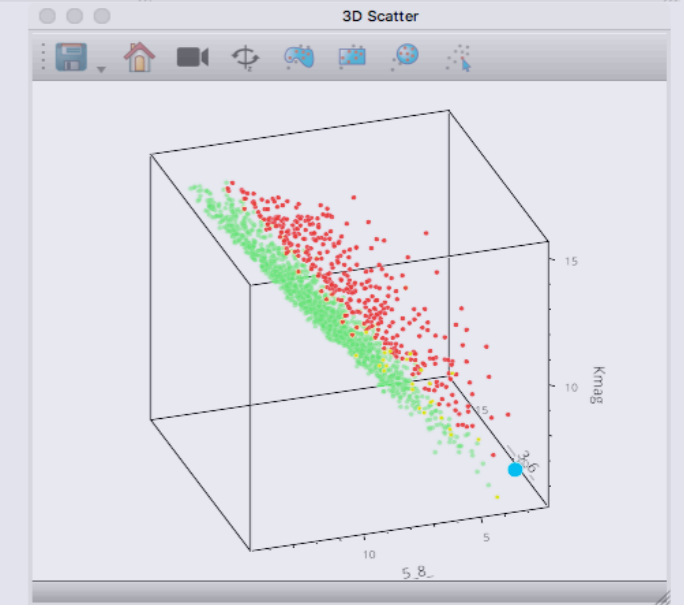
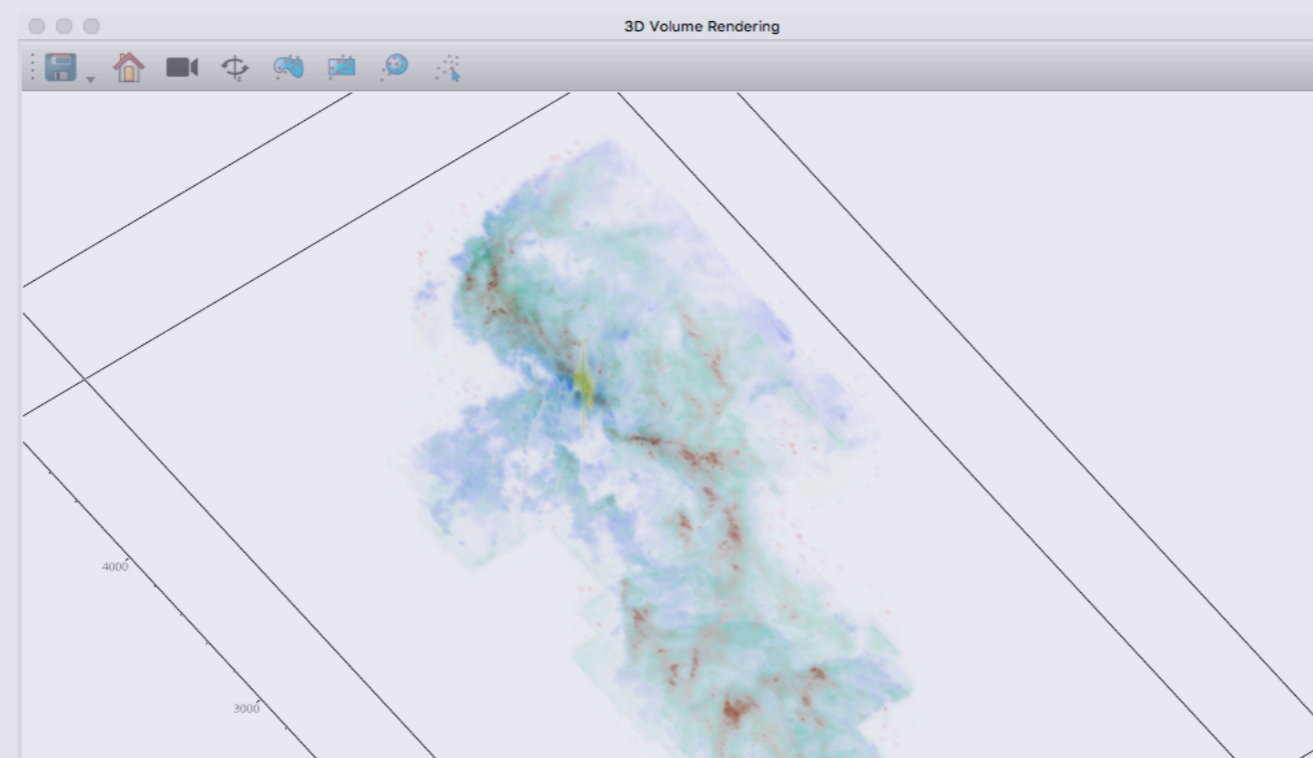
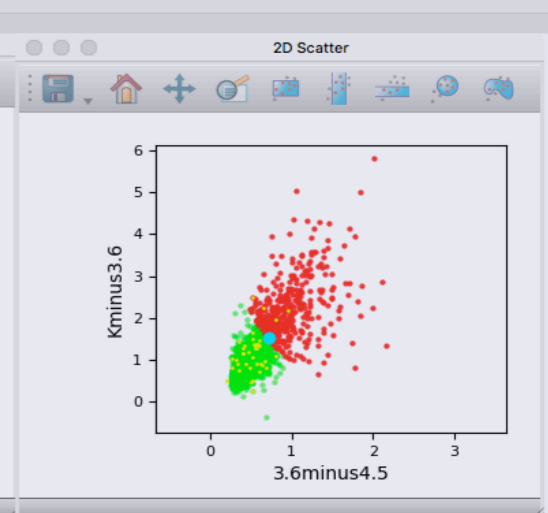
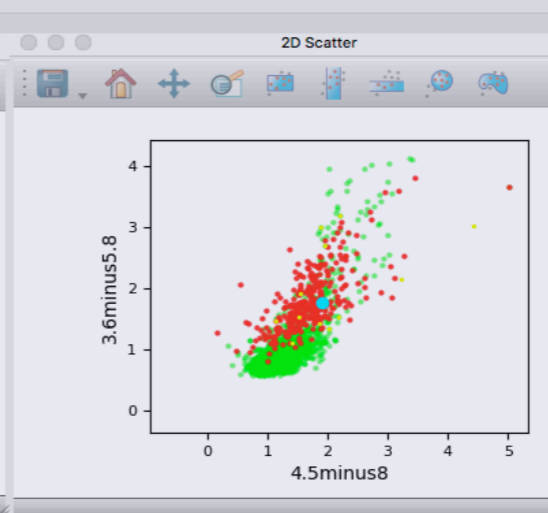
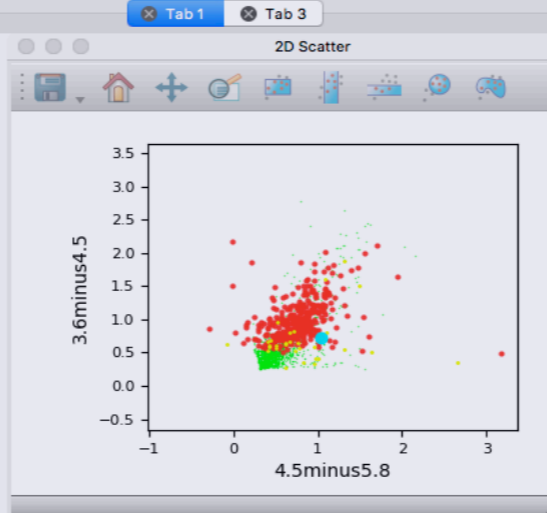
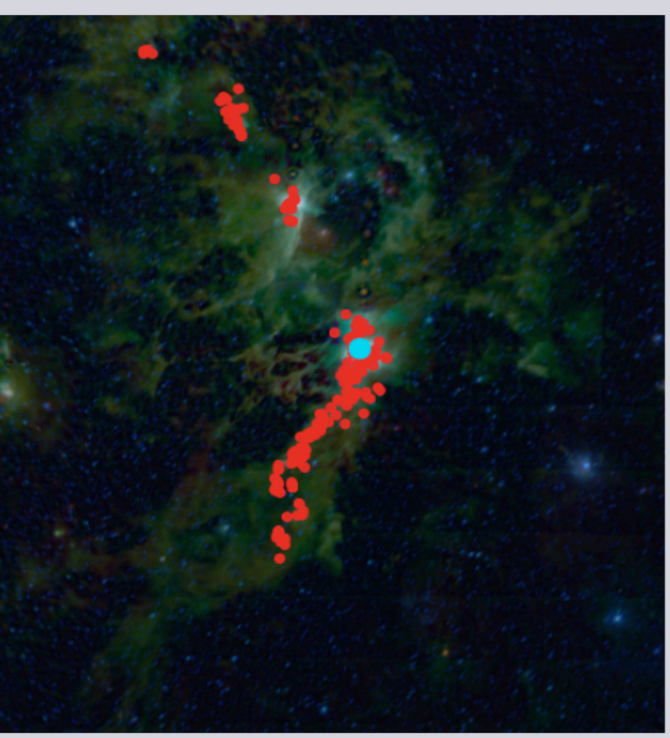
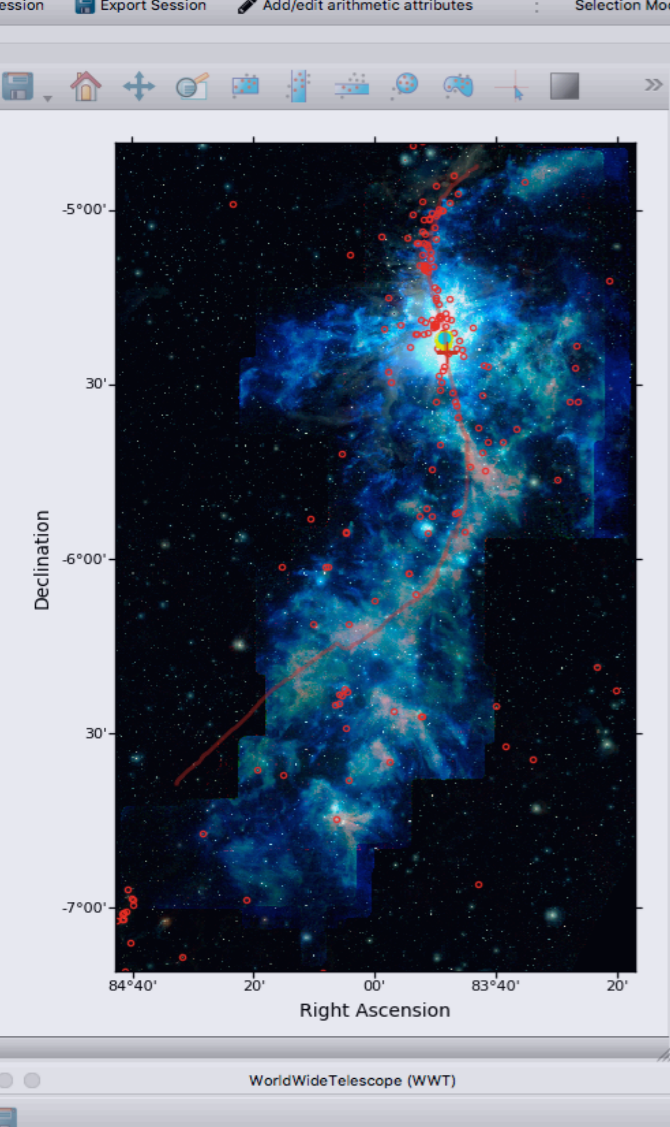
Visualization & the Universe

How and why astronomers, doctors, and you need to work together to understand the world around us...

Alyssa A. Goodman

Harvard-Smithsonian Center for Astrophysics + Radcliffe Institute for Advanced Study
@AlyssaAGoodman





What anyone can do right now



What computer
scientists do

What scientists
do



What we should do?

3D PDF
2009

PotF
2015

10QViz
2018

glupyter
2019

+ more, together

1610: Galileo Discovers Jupiter's Moons

Scipio Principe

Galileo Galilei, Familiare Seruo della Ser.^a V.^a inuigilanza
 Do assistendo, et ad ogni spirito se benero no solo satisfare
 aluano che non della letura de Mathematici nelle Scu-
 ole di Padoua,

Inuere d'auere determinato di presentare al Scipio Principe
 l'Orchiale et il p. di Giouamento inestimabile di ogni
 negozio et in circa marittima o terrestre stimo di tenere quel
 che nouo artificio ne l' maggior segrete et solam a disposizione
 di V. Ser.^a L'Orchiale auato dalle piu u. di ite speculazioni di
 prospetua ha l'uantaggio di scoprire Legni et Vele dell' inimico
 di due hore et piu di tempo prima che egli sia sopra noi et distinguendo
 il numero et la qualita de i Vasselli giudicare la sua forte
 pallesirsi alla caccia al combattimento o alla fuga, o pure esser
 nella campagna aperta uedere et particolarmente distinguere ogni suo
 moto et preparatione.

Adi 7. di Gennaio
 Giove si uede a 7. * uici: 10. 11.
 Adi 8. uici: 7. * * * * *
 4. * * * * * ora d'uy diretto et no retrogrado
 Adi 12. si uede in tale uisione * * * * *
 N. 13. si uede non uicini a Giove 4 stelle * * * * * in ogni uici
 Adi 14. è angelo * * * * *
 N. 15. * * * * * la pressi a 7. ora in uici la 4. ora di =
 stante dalla 3.^a il doppio la 1.^a
 Lo spazio delle 3. uici d'ora ad ora
 maggiore del diametro di 7. et e =
 uici in linea retta.

East		West	
7	* * ○ *	17	* ○
8	○ * * *	18	* ○ *
10	* * ○	19	* ○ * *
11	* * ○	19	* ○ * *
12	* ○ *	20	○ * ○ ○
13	* ○ * *	21	... ○
15	○ * * * *	22	* ○ * *
15	○ * * *	22	* ○ * *
16	* ○ *	23	* ○ *
17	* ○ *	24	* ○ *

SIDEREUS NUNCIUS 75

On the third, at the seventh hour, the stars were arranged in this sequence. The eastern one was 1 minute, 30 seconds from Jupiter; the closest western one 2 minutes; and the other western one was 10 minutes removed from this one. They were absolutely on the same straight line and of equal magnitude.

On the fourth, at the second hour, there were four stars around Jupiter, two to the east and two to the west, and arranged precisely on a straight line, as in the adjoining figure. The easternmost was distant 3 minutes from the next one, while this one was 40 seconds from Jupiter; Jupiter was 4 minutes from the nearest western one, and this one 6 minutes from the westernmost one. Their magnitudes were nearly equal; the one closest to Jupiter appeared a little smaller than the rest. But at the seventh hour the eastern stars were only 30 seconds apart. Jupiter was 2 minutes from the nearer eastern one, while he was 4 minutes from the next western one, and this one was 3 minutes from the westernmost one. They were all equal and extended on the same straight line along the ecliptic.

On the fifth, the sky was cloudy.

On the sixth, only two stars appeared flanking Jupiter, as is seen in the adjoining figure. The eastern one was 2 minutes and the western one 3 minutes from Jupiter. They were on the same straight line with Jupiter and equal in magnitude.

On the seventh, two stars stood near Jupiter, both to the east, arranged in this manner.

Notes for & re-productions of Siderius Nuncius

WorldWide Telescope: Explaining Galileo's Discovery

GALILEO'S "NEW ORDER"

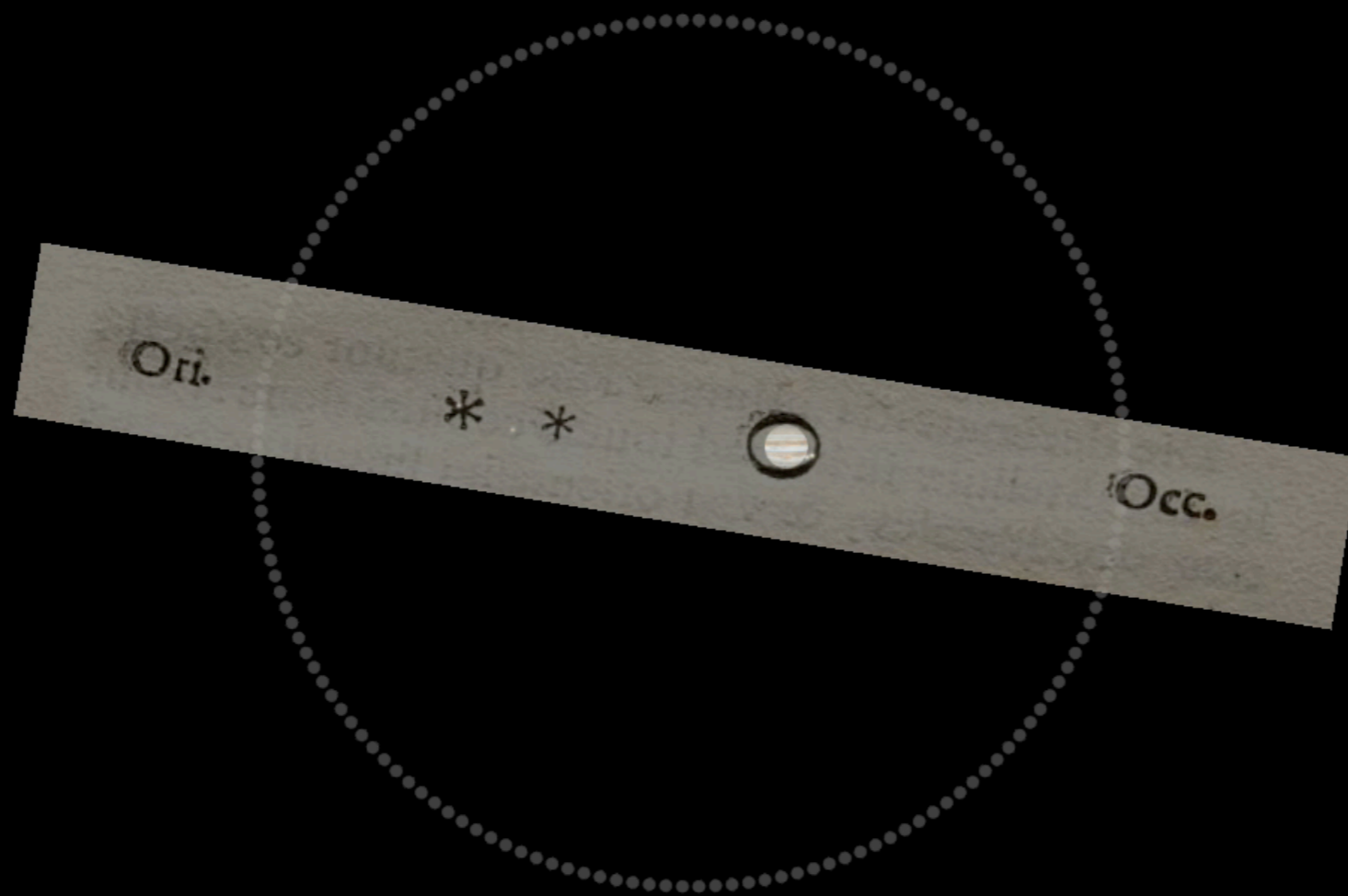
Created by Alyssa Goodman, Curtis Wong and Pat Udomprasert,
with advice from Owen Gingerich and David Malin



Galileo's New Order, A WorldWide Telescope Tour by Goodman, Wong & Udomprasert 2010
WWT Software Wong (inventor, MS Research), Fay (architect, MS Research), et al., now open source, hosted by AAS
see wwtambassadors.org for more on WWT Outreach

WorldWide Telescope: Explaining Galileo's Discovery

January 11, 1610



Galileo's New Order, A WorldWide Telescope Tour by Goodman, Wong & Udomprasert 2010
WWT Software Wong (inventor, MS Research), Fay (architect, MS Research), et al., now open source, hosted by AAS
see wwtambassadors.org for more on WWT Outreach

WorldWide Telescope: Exploring the Sky at Many Wavelengths

Home Explore Guided Tours Search Communities View Settings [Install Windows Client](#) [Sign Out](#)

Use Layer Manager to Control User Settings View From This Location

Name My Location
Lat 47:43:01 Alt 100 au
Lng -123:05:08

2009/07/29 06:01:17
X 10000 : Paused
Now

Galactic Plane Mode

Layers

- Sun
 - Mercury
 - Venus
- Earth
- Mars
- Jupiter
- Saturn
- Uranus
- Neptune
- Pluto

Sky

- Overlays
 - Constellations
 - Constellation Pictures
 - Constellation Figures
 - Constellation Boundaries
 - Constellation Names
- Grids
 - Equatorial Grid
 - Galactic Grid
 - AltAz Grid
 - Ecliptic Grid
 - Ecliptic Overview
 - Precession Chart
- 2d Sky

Time Scrubber

Look At: Sky Imagery: Digitized Sky Survey (Color) Image Crossfade

WorldWide Telescope

Tracking SFD Dust Map (Infrared)

RA: 16h26m05s Dec: -25:20:17

Cetus 14:56:51

[demo]

Visualization in Computer Science



VIS 2018 SCHEDULE

WEDNESDAY, 24 OCTOBER

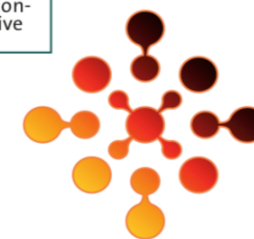
	VAST Conv 1, Sec C	INFOVIS Conv 1, Sec D	SCIVIS Estrel A+B	VIS Estrel C	VIS Room IV
9:00 AM	Ensemble and Provenance DS	Time	Biomedical Visualization BA	VAST: Text DS BA	VisAP: Arts Program Session 1: Arts and Society
10:40 AM BREAK	BREAK				
11:00 AM	Applications AP	Graphs & Trees	Volume Visualization	InfoVis: Devices: Small & Large UX	CG&A Session 1
12:40 PM	Restructuring IEEE VIS for the Future (1:00–2:20 PM) @ Conv 1, Sec C		LUNCH		
LUNCH					
2:20 PM	High Dimensional Data DS	Text & Communication AP	Space and Physics SS	SciVis Contest SS	Panel: Perspectives in Color Research for Scientific Visualization
4:00 PM BREAK	BREAK				
4:20 PM	AP Fast Forward (Thu & Fri Sessions) (4:20–5:20 PM) @ Estrel A+B				
6:00 PM	AP Posters + Networking + Hiring events (5:20–7:00 PM) @ Foyers				
7:00 PM	VIS Dinner Banquet (7:00–9:00 PM) @ Conv 1, Sec C+D				
9:00 PM					

THURSDAY, 25 OCTOBER

	VAST Conv 1, Sec C	INFOVIS Conv 1, Sec D	SCIVIS Estrel A+B	VIS Estrel C	VIS Room IV
9:00 AM	Security, Privacy, and Anomaly	Immersive Analytics UX	Tensors	VAST: Interactive Analytics and Design DS	VisAP: Arts Program Session 2: Paths and Memories
10:40 AM BREAK	BREAK				
11:00 AM	Deep Learning DS	Design & Storytelling DJ	Scalable Techniques SS	InfoVis: Interaction UX	CG&A Session 2 SS BA
12:40 PM	LUNCH		VIS 2019 Kick-off Meeting (1:00–2:20 PM) @ Estrel C		
LUNCH					
2:20 PM	Graph and Image	Perception & Cognition 1	Topology, Geometry, and Precision	SciVis Short: Visual Abstractions, Perceptual Study and Immersive Vis SS UX	Panel: Meet the Founders: How to Start and Sustain a Business in the Vis Space AP
4:00 PM BREAK	BREAK				
4:20 PM	Explainable ML DS	Perception & Cognition 2	Interaction and Multivariate Data UX	SciVis Short: Flow, Astrophysics, and Computationally Intensive Data Vis	

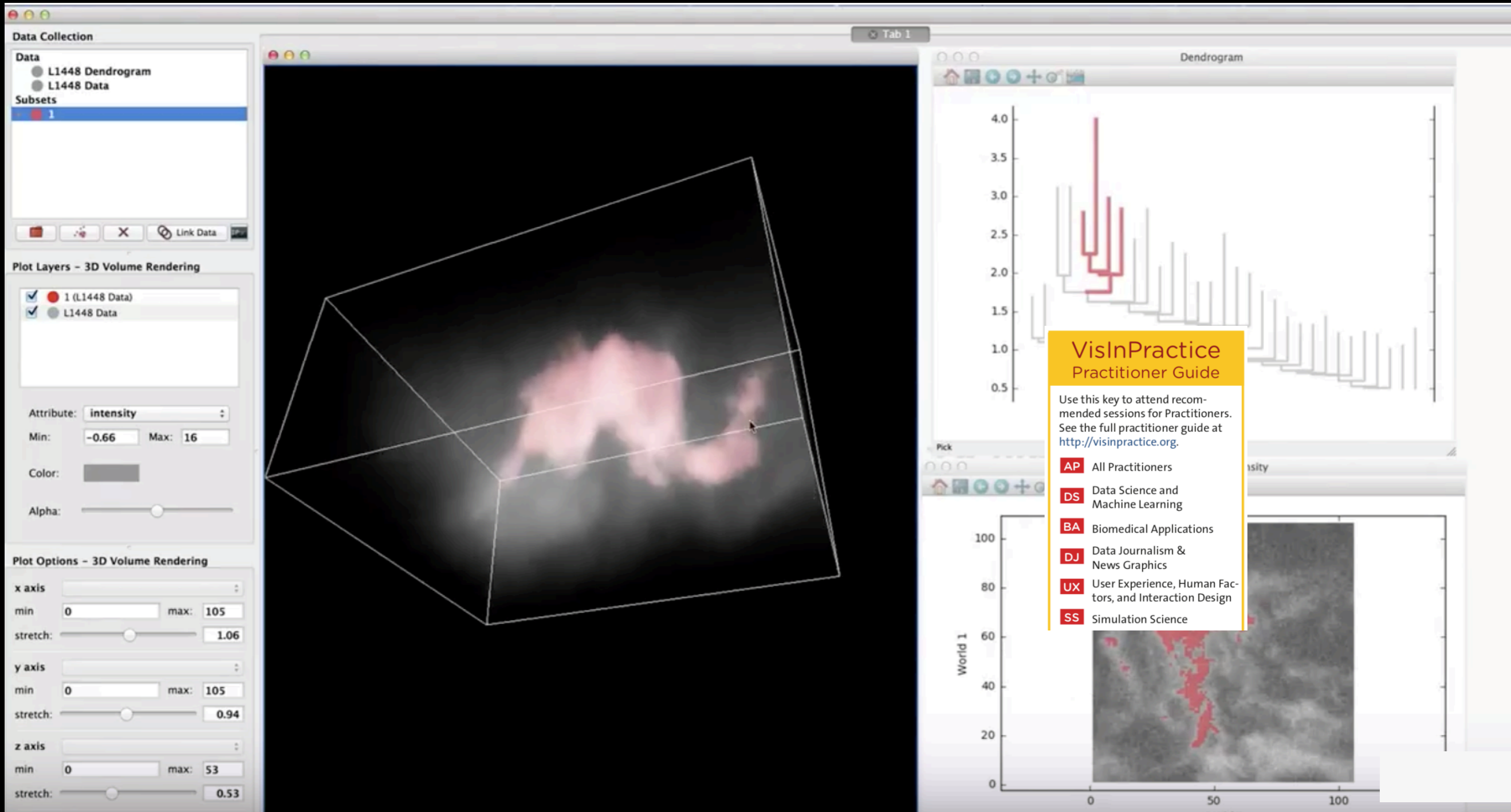
FRIDAY, 26 OCTOBER

	VAST Conv 1, Sec C	INFOVIS Conv 1, Sec D	SCIVIS Estrel C	
9:00 AM	Event, Sequence, and ML DS	Uncertainty & Error	Time-varying Data SS	9:00 AM
10:40 AM BREAK	BREAK			10:40 AM BREAK
11:00 AM	AP VIS Capstone (11:00 AM–12:00 PM) Can I Believe What I See?—Information-Theoretic Algorithm Validation Joachim M. Buhmann, <i>ETH Zurich</i>			11:00 AM
12:40 PM	AP VIS Closing (12:00–12:30 PM) @ Conv 1, Sec C			12:40 PM
LUNCH				LUNCH
2:20 PM	VisInPractice Practitioner Guide Use this key to attend recommended sessions for Practitioners. See the full practitioner guide at http://visinpractice.org .			2:20 PM
4:00 PM BREAK	Posters Located in Foyer 1 & Foyer Estrel Hall Sunday—Thursday, 9:00 AM–6:00 PM			4:00 PM BREAK
4:20 PM	Exhibitions Located in Foyer 3 Tuesday—Thursday, 10:40 AM–6:00 PM			4:20 PM
6:00 PM	VIS Arts Program Located in Room Paris Opening, Tuesday 7:00 PM Exhibition, Wednesday—Thursday, 9:00 AM–6:00 PM			6:00 PM
7:00 PM				7:00 PM
9:00 PM				9:00 PM



“Information Visualization” “VAST” “Scientific Visualization”

Visualization in Computer Science



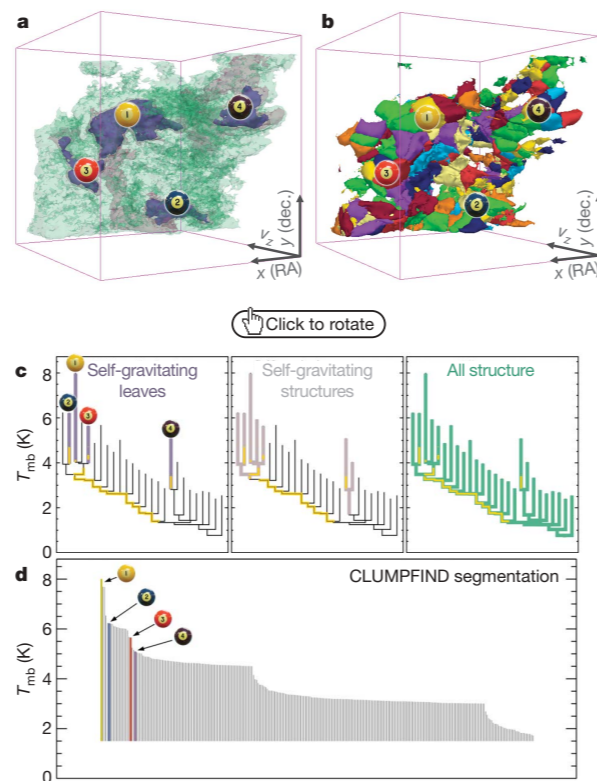


Figure 2 | Comparison of the 'dendrogram' and 'CLUMPFIND' feature-identification algorithms as applied to ^{13}CO emission from the L1448 region of Perseus. **a**, 3D visualization of the surfaces indicated by colours in the dendrogram shown in **c**. Purple illustrates the smallest scale self-gravitating structures in the region corresponding to the leaves of the dendrogram; pink shows the smallest surfaces that contain distinct self-gravitating leaves within them; and green corresponds to the surface in the data cube containing all the significant emission. Dendrogram branches corresponding to self-gravitating objects have been highlighted in yellow over the range of T_{mb} (main-beam temperature) test-level values for which the virial parameter is less than 2. The x - y locations of the four 'self-gravitating' leaves labelled with billiard balls are the same as those shown in Fig. 1. The 3D visualizations show position-position-velocity (p - p - v) space. RA, right ascension; dec., declination. For comparison with the ability of dendrograms (**c**) to track hierarchical structure, **d** shows a pseudo-dendrogram of the CLUMPFIND segmentation (**b**), with the same four labels used in Fig. 1 and in **a**. As 'clumps' are not allowed to belong to larger structures, each pseudo-branch in **d** is simply a series of lines connecting the maximum emission value in each clump to the threshold value. A very large number of clumps appears in **b** because of the sensitivity of CLUMPFIND to noise and small-scale structure in the data. In the online PDF version, the 3D cubes (**a** and **b**) can be rotated to any orientation, and surfaces can be turned on and off (interaction requires Adobe Acrobat version 7.0.8 or higher). In the printed version, the front face of each 3D cube (the 'home' view in the interactive online version) corresponds exactly to the patch of sky shown in Fig. 1, and velocity with respect to the Local Standard of Rest increases from front (-0.5 km s^{-1}) to back (8 km s^{-1}).

data, CLUMPFIND typically finds features on a limited range of scales, above but close to the physical resolution of the data, and its results can be overly dependent on input parameters. By tuning CLUMPFIND's two free parameters, the same molecular-line data set⁸ can be used to show either that the frequency distribution of clump mass is the same as the initial mass function of stars or that it follows the much shallower mass function associated with large-scale molecular clouds (Supplementary Fig. 1).

Four years before the advent of CLUMPFIND, 'structure trees'⁹ were proposed as a way to characterize clouds' hierarchical structure

using 2D maps of column density. Very recently, the work of Fluke et al. as inspiration, we have developed a structure tree that abstracts the hierarchical structure of the data into an easily visualized representation. This method, well developed in other data-intensive fields, is the application of tree methodology to the analysis of almost exclusively within the context of astronomy. 'merger trees' are being used to analyze the evolution of galaxy clusters.

Figure 3 and its legend explain the dendrogram process schematically. The dendrogram of the Perseus emission merge with the Perseus emission is explained in Supplementary Methods and Supplementary Fig. 2. The dendrogram is determined almost entirely by the data and its sensitivity to algorithm parameters is limited to what is possible on paper and 2D data (see Fig. 3 and its legend). The dendrogram is a cross, which eliminates the need for a separate step of preserving all information. The dendrogram is a 'Numbered 'billiard ball' dendrogram' (see Supplementary online) and a sorted dendrogram.

A dendrogram of a physical property, such as surfaces, such as radius (R), velocity (v), or luminosity (L). The volumes can have any shape, and in our case, the significance of the especially elongated features seen in Fig. 2a). The luminosity is an approximate proxy for mass, such that $M_{\text{lum}} = X_{13\text{CO}} L_{13\text{CO}}$, where $X_{13\text{CO}} = 8.0 \times 10^{20} \text{ cm}^2 \text{ K}^{-1} \text{ km}^{-1} \text{ s}$ (ref. 15; see Supplementary Methods and Supplementary Fig. 2). The derived values for size, mass and velocity dispersion can then be used to estimate the role of self-gravity at each point in the hierarchy, via calculation of an 'observed' virial parameter, $\alpha_{\text{obs}} = 5\sigma_v^2 R / GM_{\text{lum}}$. In principle, extended portions of the tree (Fig. 2, yellow highlighting) where $\alpha_{\text{obs}} < 2$ (where gravitational energy is comparable to or larger than kinetic energy) correspond to regions of p - p - v space where self-gravity is significant. As α_{obs} only represents the ratio of kinetic energy to gravitational energy at one point in time, and does not explicitly capture external over-pressure and/or magnetic fields¹⁶, its measured value should only be used as a guide to the longevity (boundedness) of any particular feature.

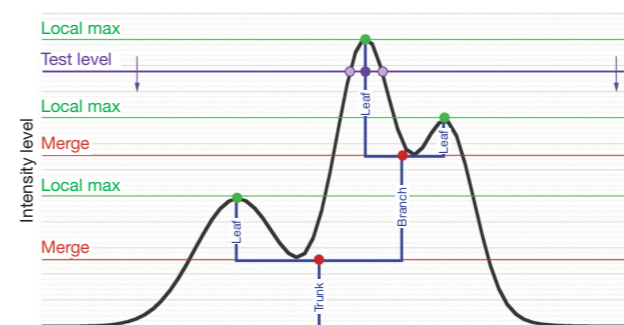
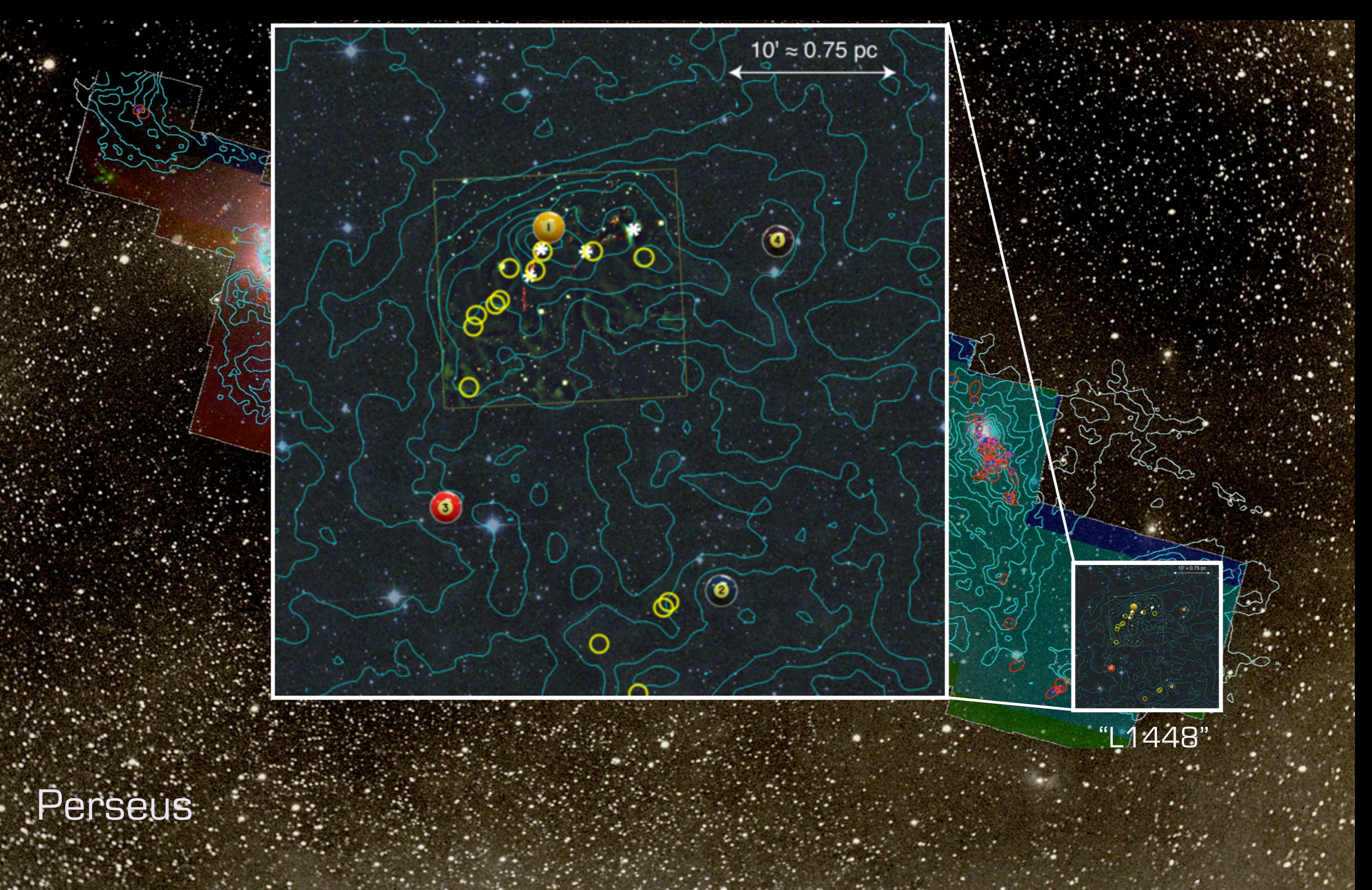
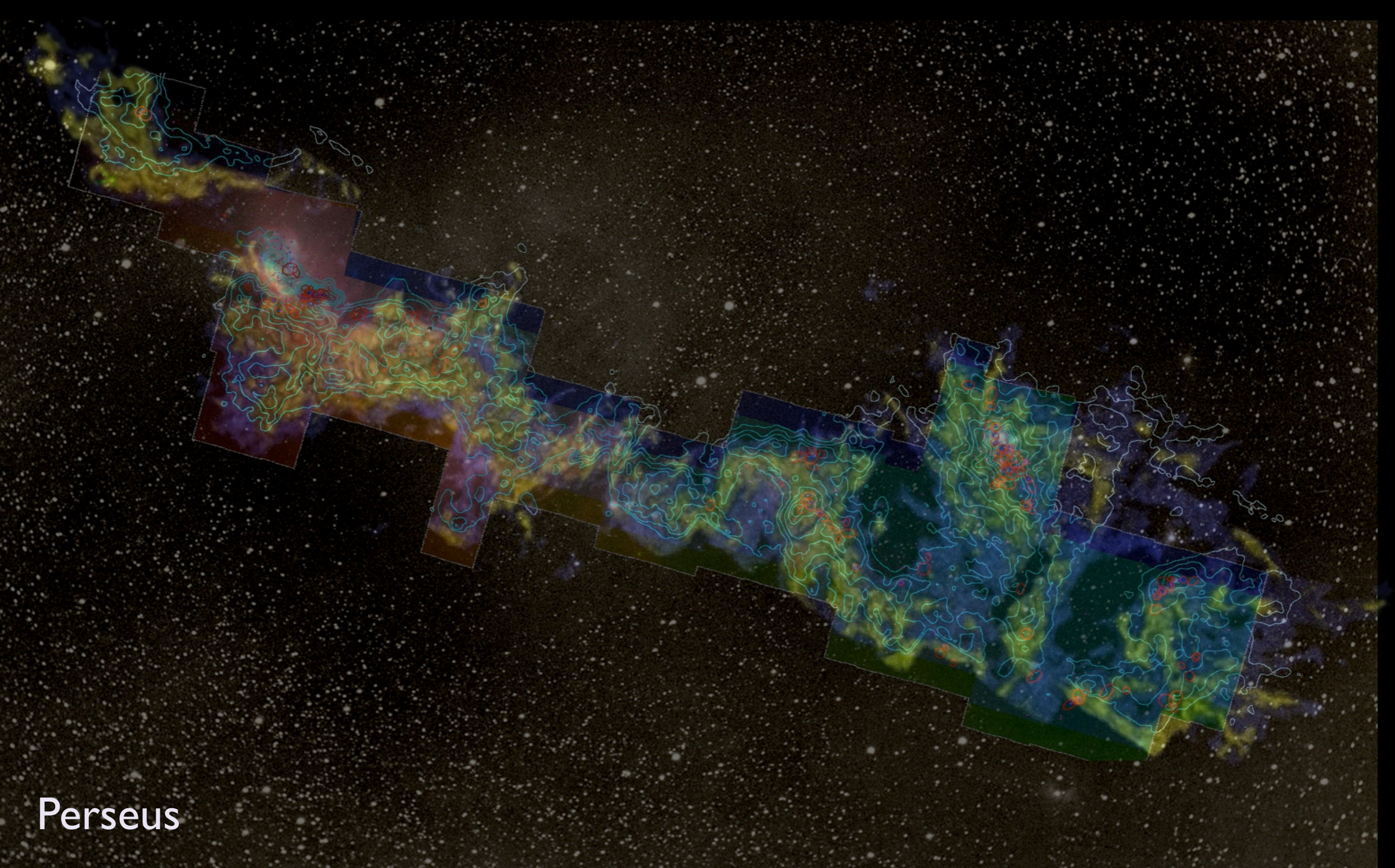


Figure 3 | Schematic illustration of the dendrogram process. Shown is the construction of a dendrogram from a hypothetical one-dimensional emission profile (black). The dendrogram (blue) can be constructed by 'dropping' a test constant emission level (purple) from above in tiny steps (exaggerated in size here, light lines) until all the local maxima and mergers are found, and connected as shown. The intersection of a test level with the emission is a set of points (for example the light purple dots) in one dimension, a planar curve in two dimensions, and an isosurface in three dimensions. The dendrogram of 3D data shown in Fig. 2c is the direct analogue of the tree shown here, only constructed from 'isosurface' rather than 'point' intersections. It has been sorted and flattened for representation on a flat page, as fully representing dendrograms for 3D data cubes would require four dimensions.

Visualization in Science



Perseus



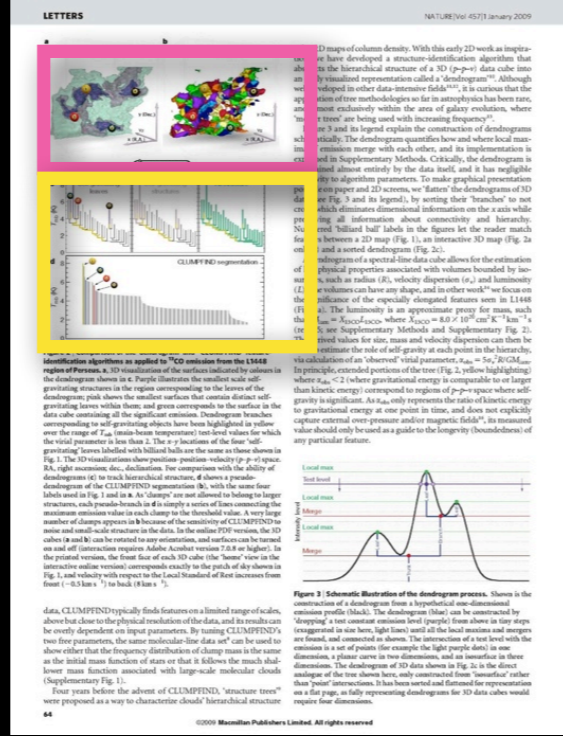
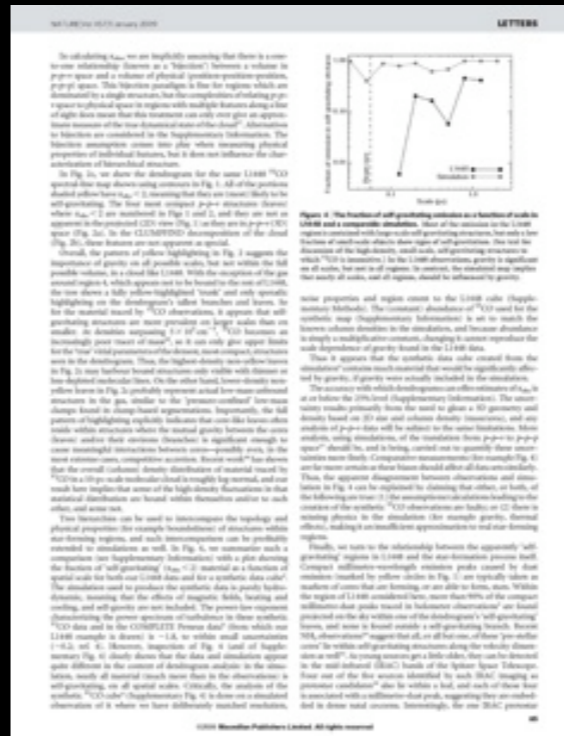
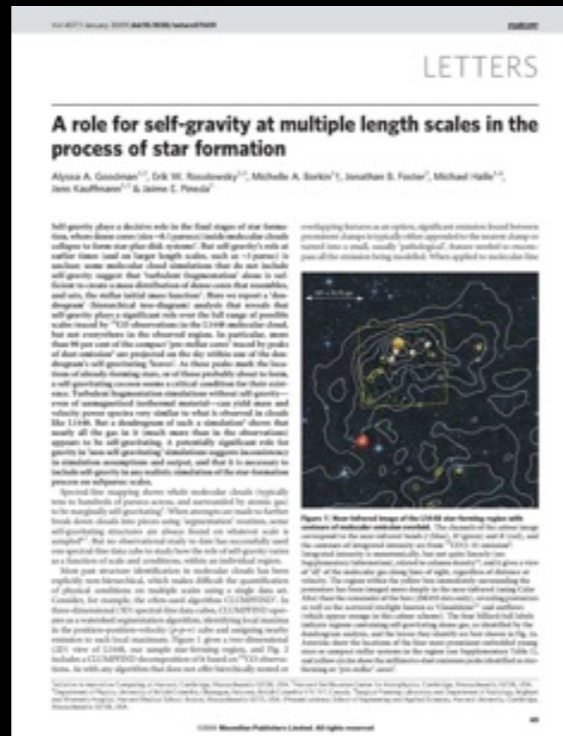
Perseus

3D Viz made with VolView

Why Astronomical **Medicine**?



Visualization in Computer Science



“Scientific Visualization”
 “Information Visualization”

“VAST”

1 / 4 131% Tools Comment Share

Vol 457 | 1 January 2009 | doi:10.1038/nature07609 nature

LETTERS

A role for self-gravity at multiple length scales in the process of star formation

Alyssa A. Goodman^{1,2}, Erik W. Rosolowsky^{2,3}, Michelle A. Borkin^{1†}, Jonathan B. Foster², Michael Halle^{1,4}, Jens Kauffmann^{1,2} & Jaime E. Pineda²

Self-gravity plays a decisive role in the final stages of star formation, where dense cores (size ~ 0.1 parsecs) inside molecular clouds collapse to form star-plus-disk systems¹. But self-gravity's role at earlier times (and on larger length scales, such as ~ 1 parsec) is unclear; some molecular cloud simulations that do not include self-gravity suggest that ‘turbulent fragmentation’ alone is sufficient to create a mass distribution of dense cores that resembles, and sets, the stellar initial mass function². Here we report a ‘denrogram’ (hierarchical tree-diagram) analysis that reveals that self-gravity plays a significant role over the full range of possible scales traced by ¹³CO observations in the L1448 molecular cloud, but not everywhere in the observed region. In particular, more than 90 per cent of the compact ‘pre-stellar cores’ traced by peaks of dust emission³ are projected on the sky within one of the denrogram's self-gravitating ‘leaves’. As these peaks mark the locations of already-forming stars, or of those probably about to form, a self-gravitating cocoon seems a critical condition for their exist-

overlapping features as an option, significant emission found between prominent clumps is typically either appended to the nearest clump or turned into a small, usually ‘pathological’, feature needed to encompass all the emission being modelled. When applied to molecular-line

The image shows a field of stars with white density contours overlaid. A yellow box highlights a specific region. A scale bar at the top left indicates 10 arcminutes, which is approximately 0.75 parsecs. Several stars are marked with numbers and symbols, including a yellow circle with '1', a red circle with '4', and several white stars.

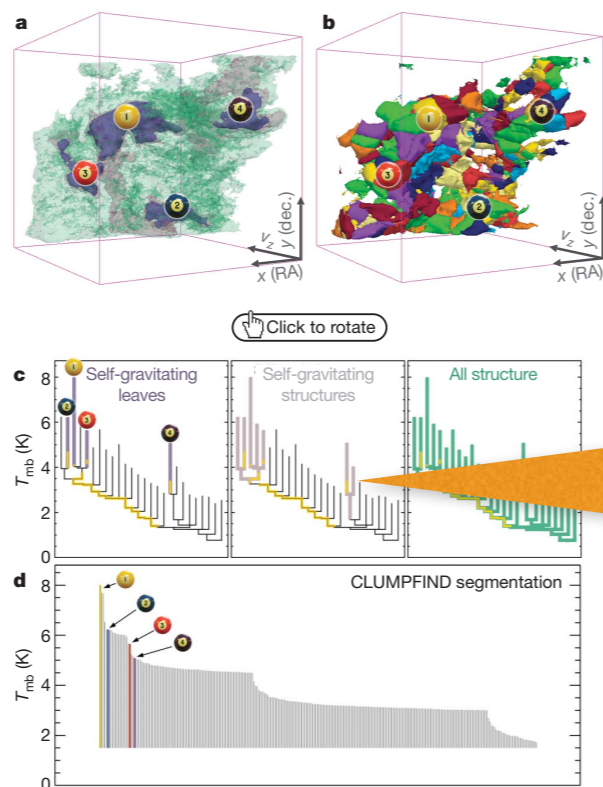


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Four years before the advent of CLUMPFIND, 'structure trees'⁹ were proposed as a way to characterize clouds' hierarchical structure

using 2D maps of column density. With this 2D work as inspiration, we have developed a structure-identification algorithm that abstracts the hierarchical structure of a data set into an easily visualized representation called a dendrogram, well developed in other data-intensive fields. The application of tree methodologies so far has been almost exclusively within the astronomical domain, and 'merger trees' are being used with increasing frequency.

Figure 3 and its legend explain the dendrogram process schematically. The dendrogram and its associated

These are "dead" panels! That's not good enough.

used to estimate the role of self-gravity at each point in the hierarchy, via calculation of an 'observed' virial parameter, $\alpha_{\text{obs}} = 5\sigma_v^2 R / GM_{\text{lum}}$. In principle, extended portions of the tree (Fig. 2, yellow highlighting) where $\alpha_{\text{obs}} < 2$ (where gravitational energy is comparable to or larger than kinetic energy) correspond to regions of p - p - v space where self-gravity is significant. As α_{obs} only represents the ratio of kinetic energy to gravitational energy at one point in time, and does not explicitly capture external over-pressure and/or magnetic fields¹⁶, its measured value should only be used as a guide to the longevity (boundedness) of any particular feature.

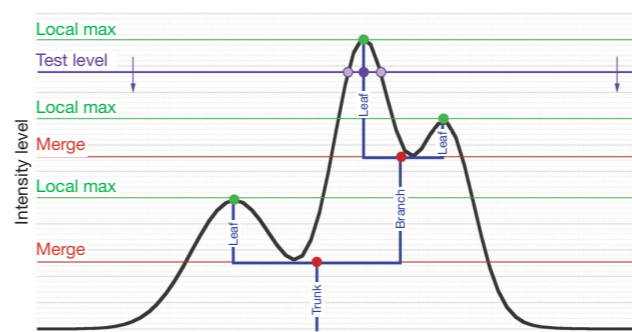
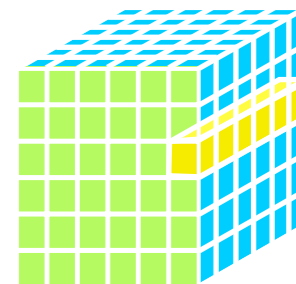


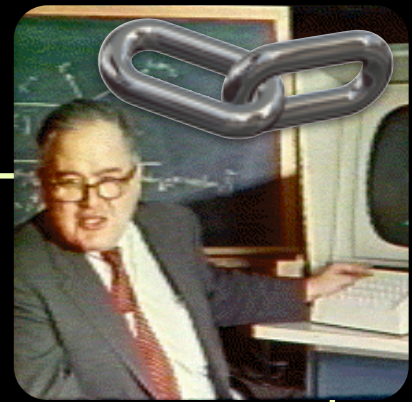
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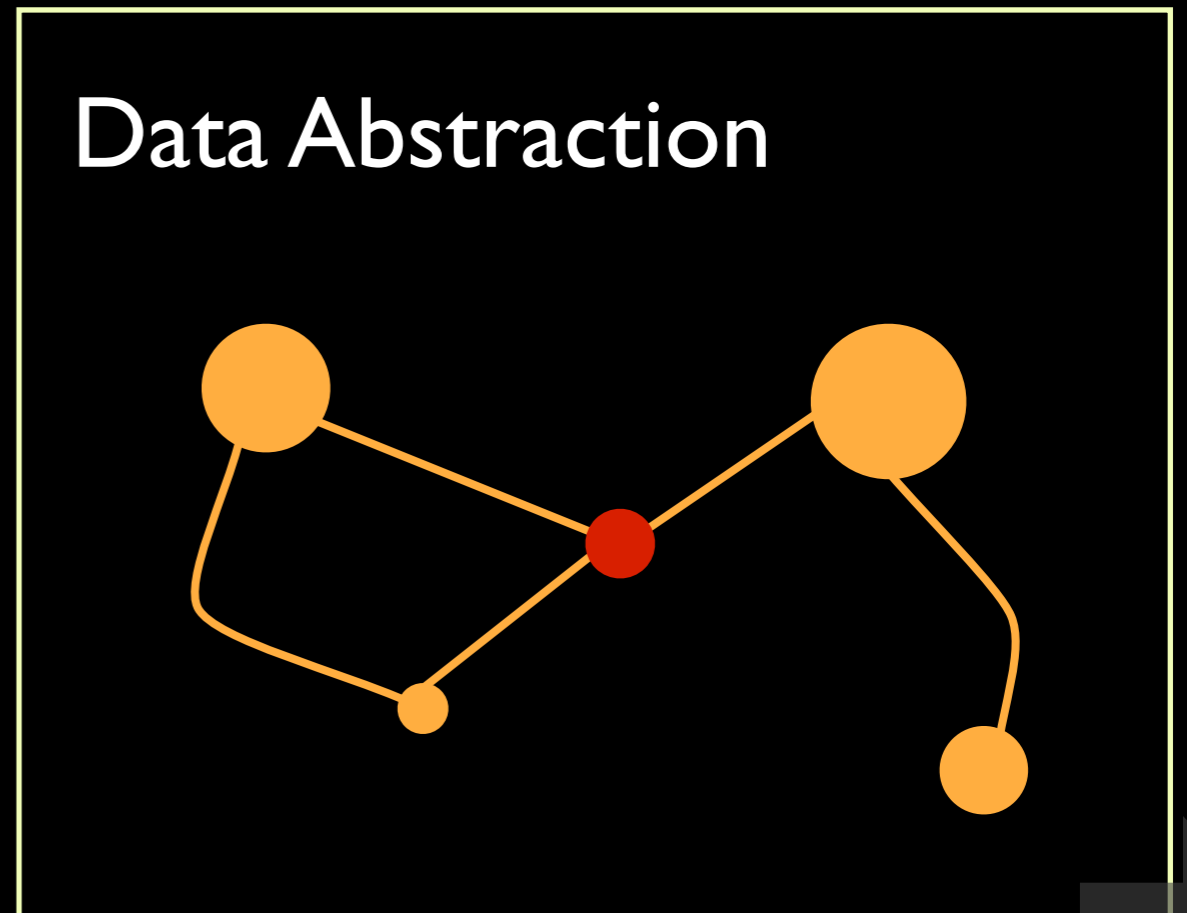
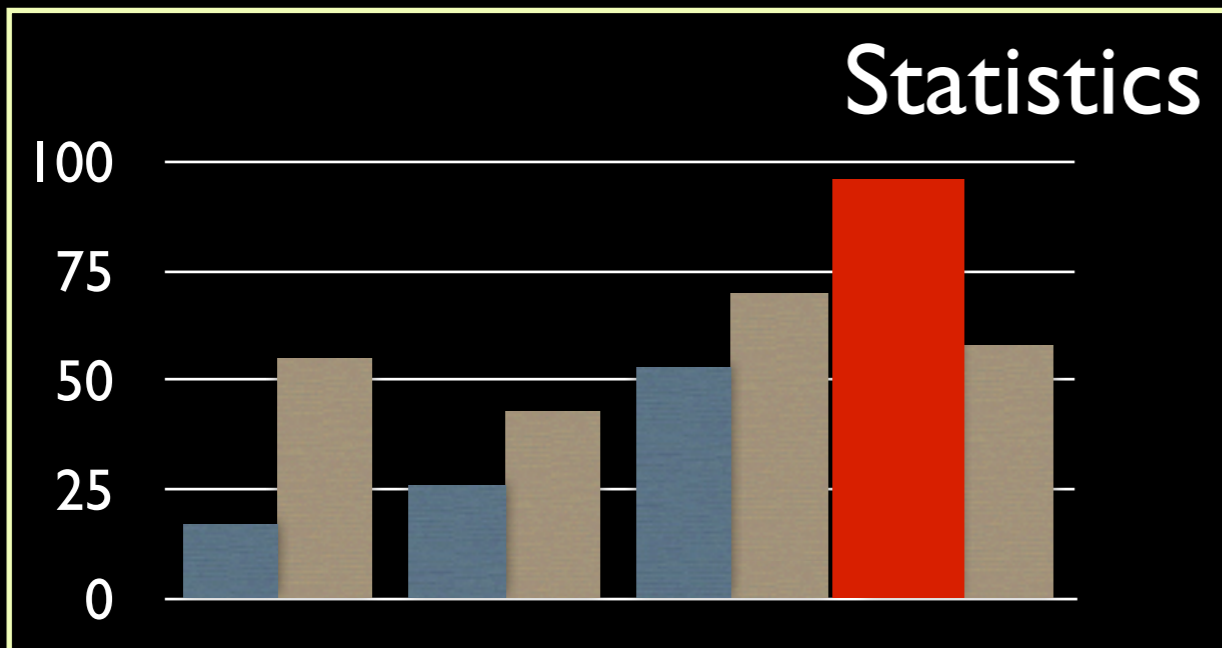
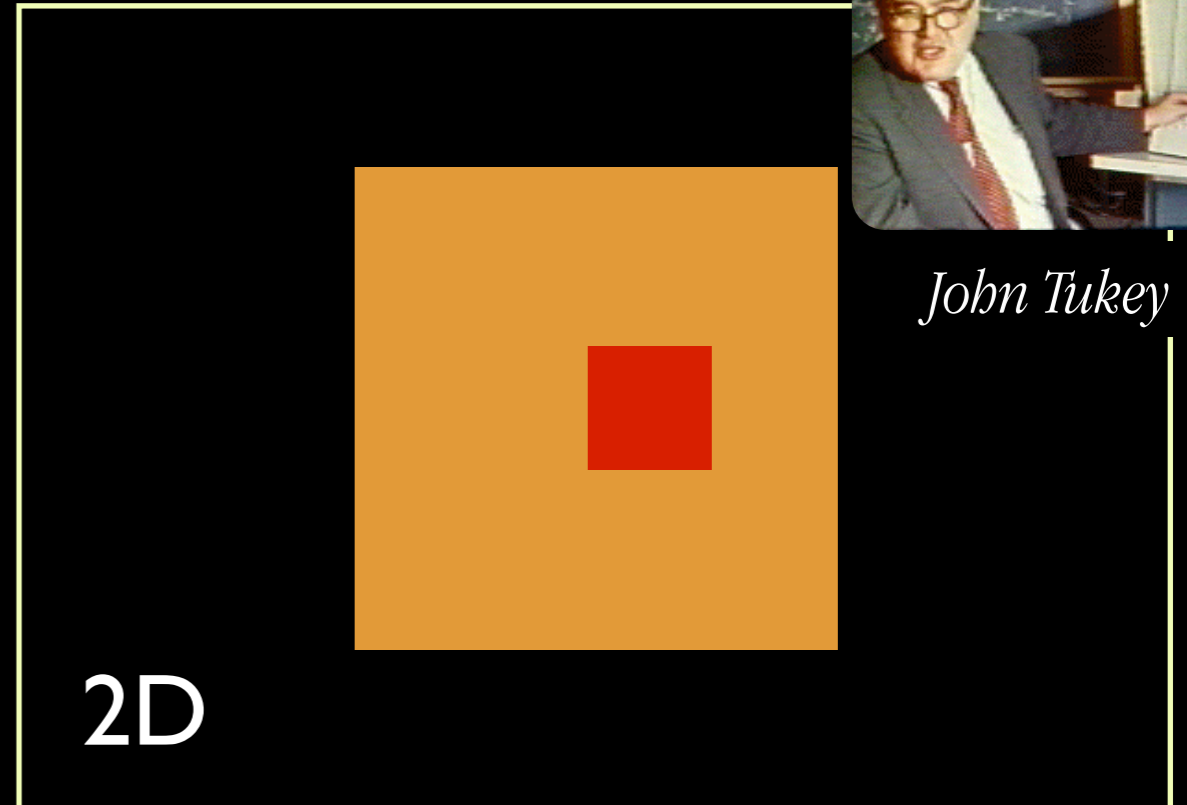
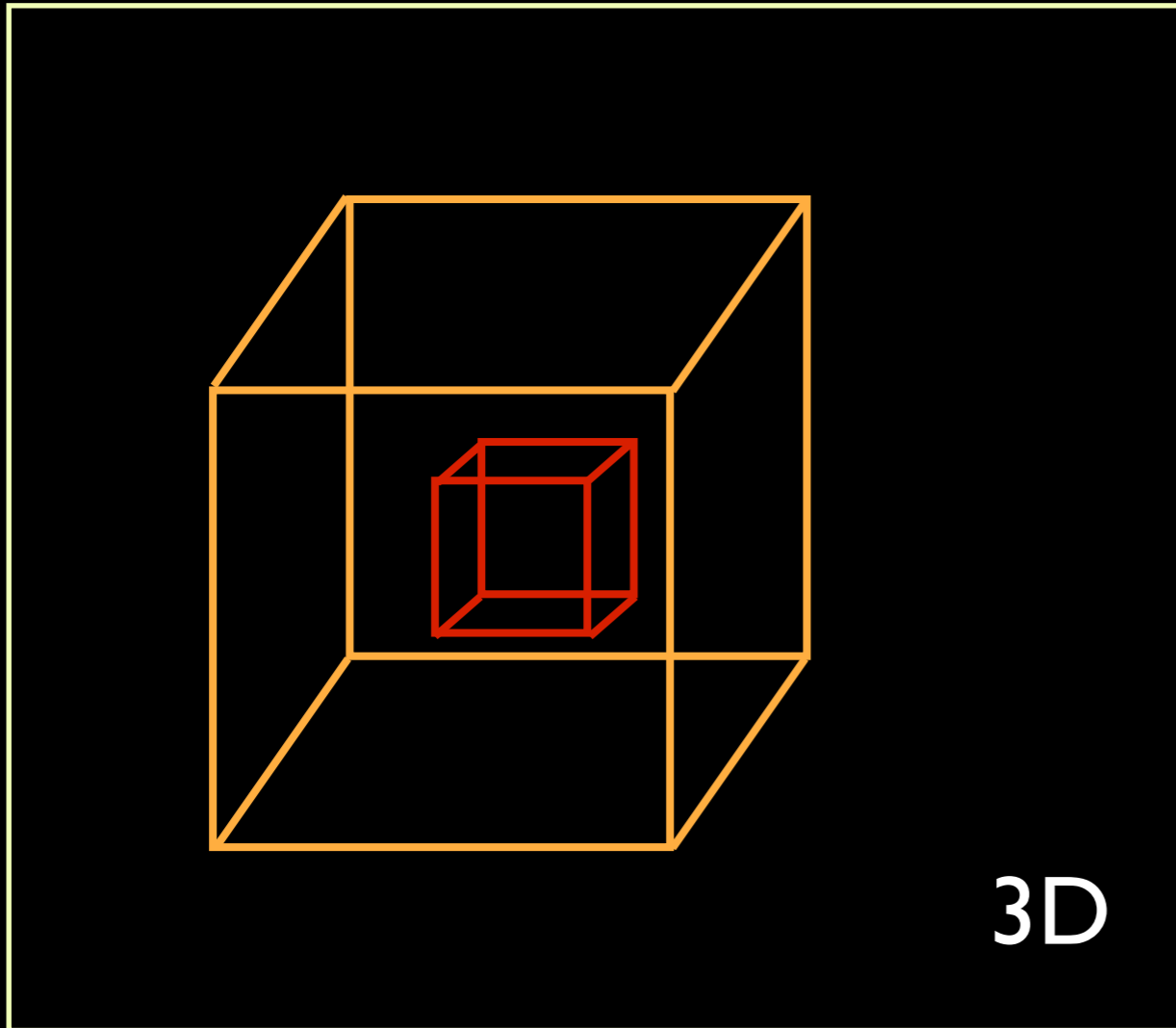
Goodman et al. 2009, Nature, cf: Fluke et al. 2009

2009
3D PDF
High-Dimensional
data in a
"Paper"
on its way
to the Future

Linked Views of High-dimensional Data



John Tukey



JOHN TUKEY'S LEGACY



PRIM-9

PRIM-H

DataDesk®



XGobi

GGobi

RGGobi



Polaris



1970

1980

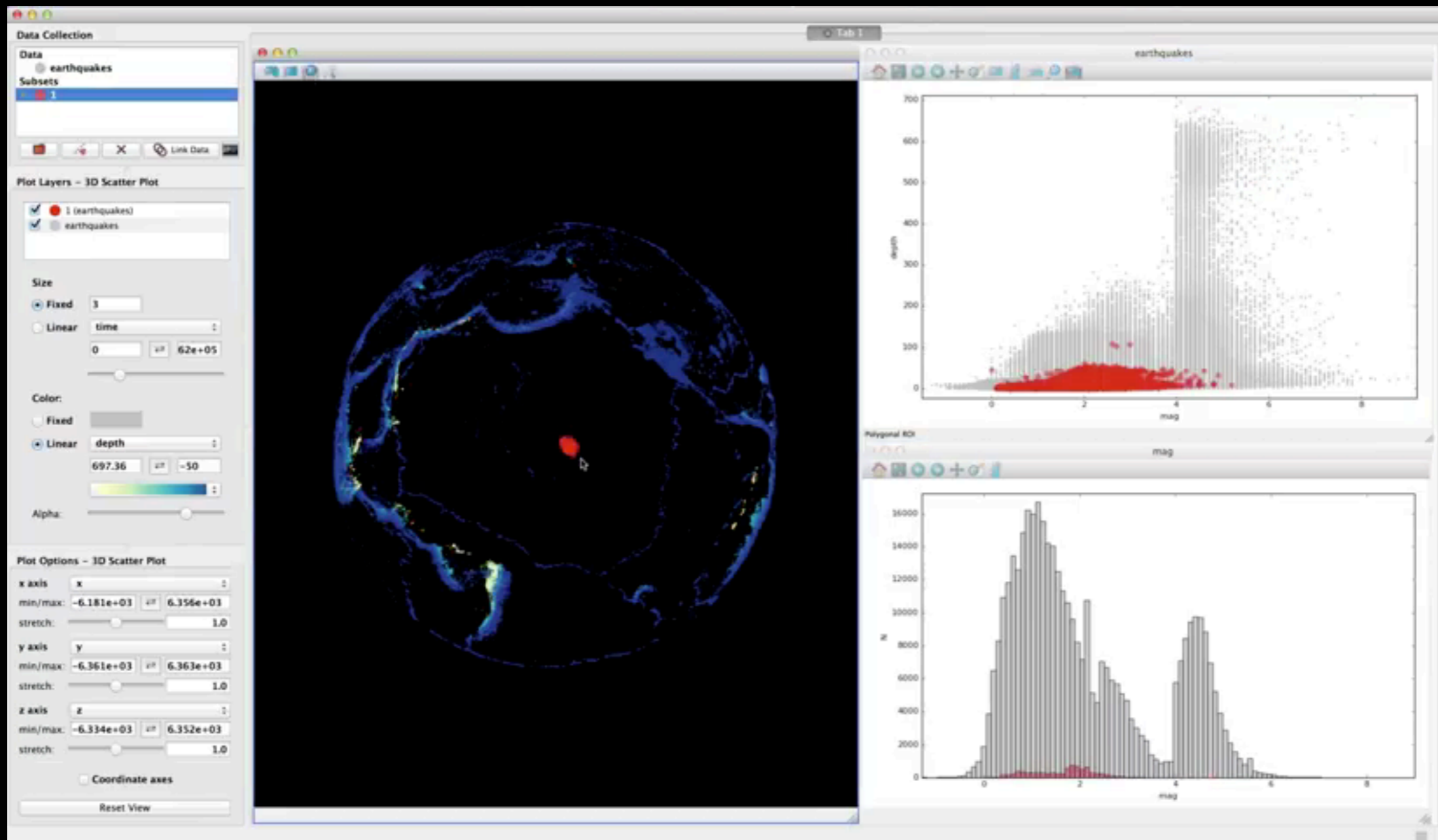
1990

2000

2010

Linked Views of High-dimensional Data (in Python)

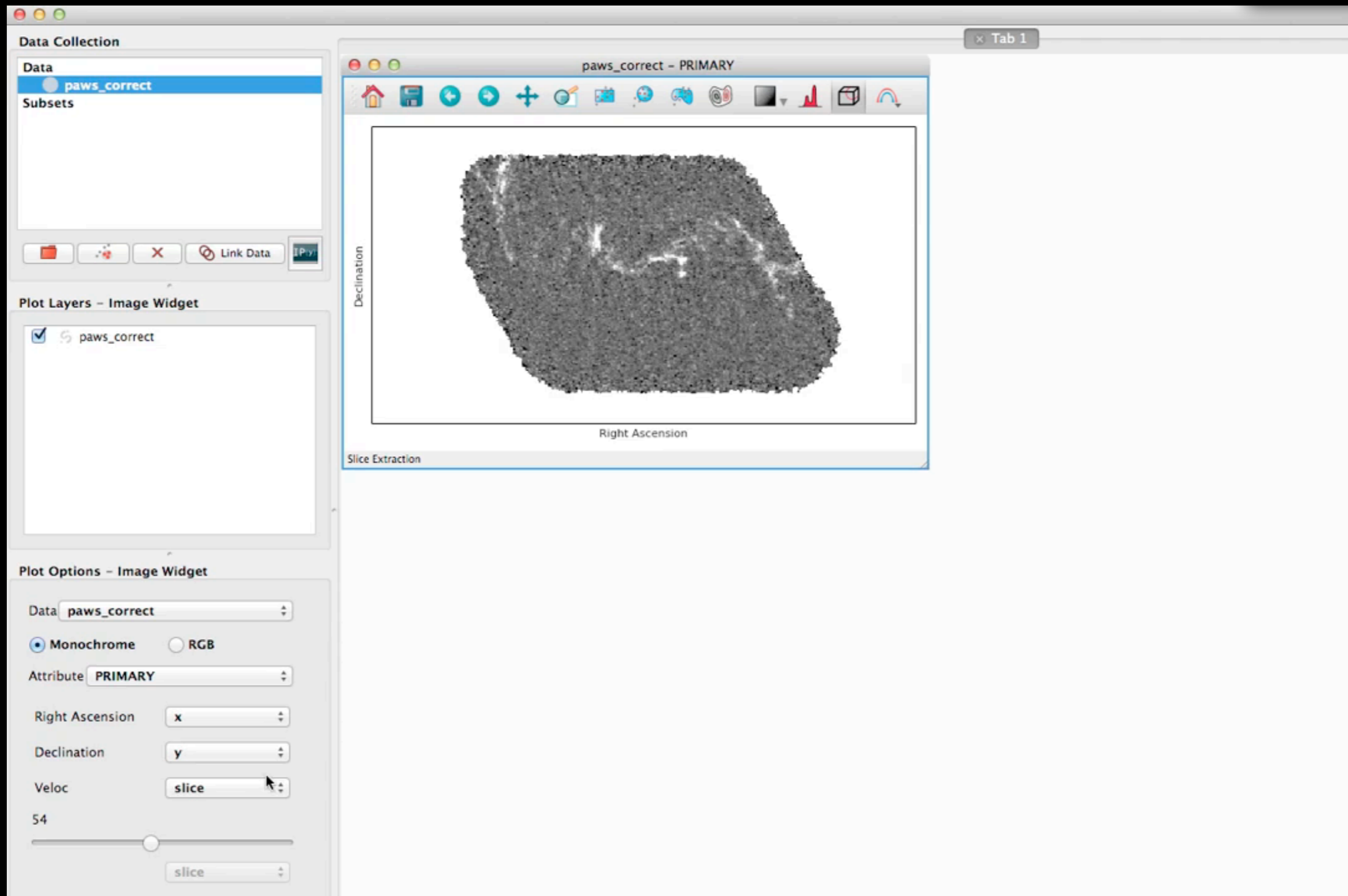
glue



video by Tom Robitaille, lead glue developer
glue created by: C. Beaumont, M. Borkin, P. Qian, T. Robitaille, and A. Goodman, PI

Linked Views of High-dimensional Data (in Python)

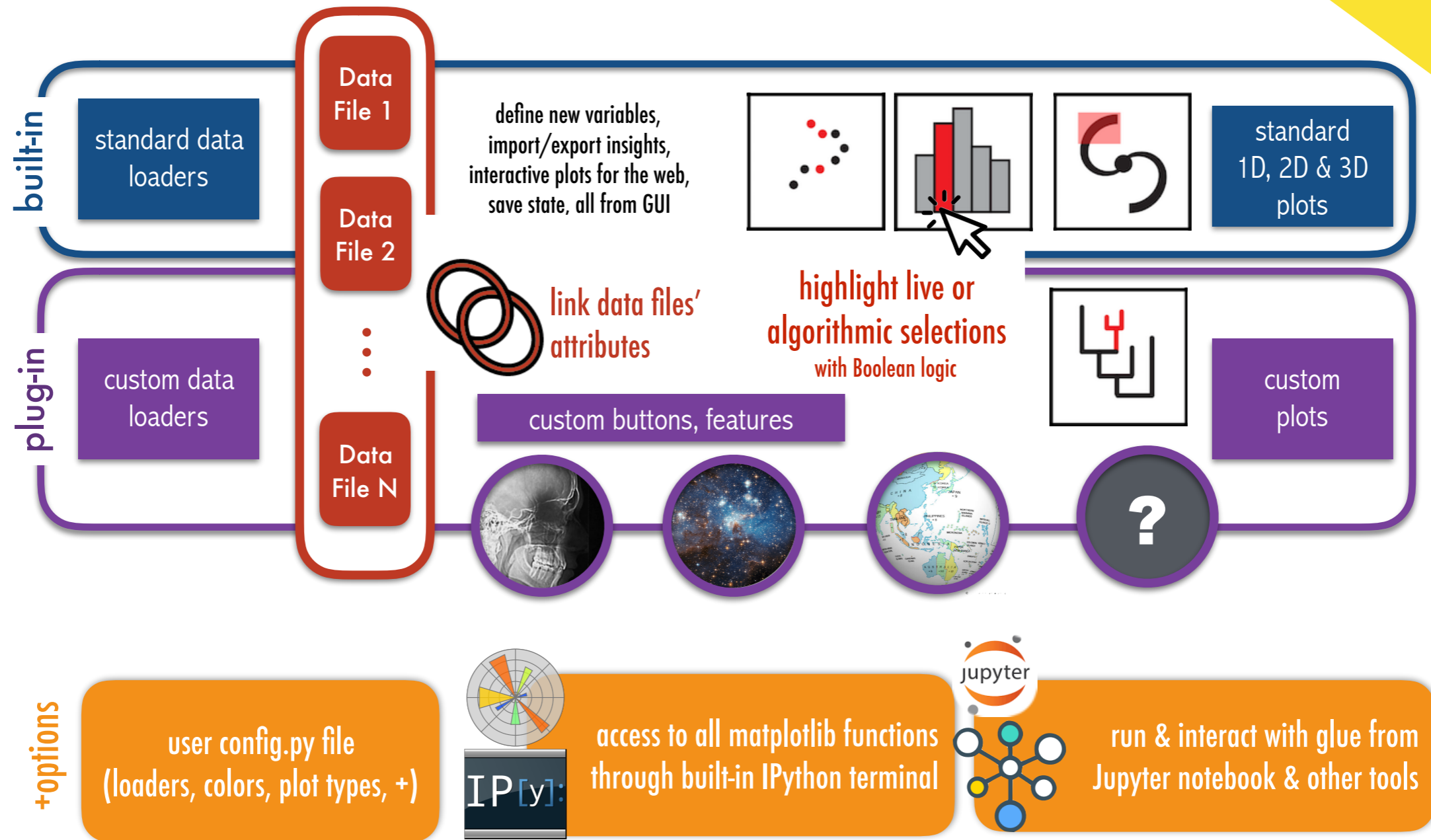
glue



*video by Chris Beaumont, glue developer
glue created by: C. Beaumont, M. Borkin, P. Qian, T. Robitaille, and A. Goodman, PI*



handout



glueviz.org

[demo]
[ALMA]
[publishing]

No merging of data sets—just glue them.

The screenshot displays the Glueviz application window. The top menu bar includes 'python', 'File', 'Edit', 'View', 'Canvas', 'Data Manager', 'Plugins', and 'Help'. The title bar shows 'Glue' and the user name 'Jorma Harju'. The main toolbar contains buttons for 'Open Data', 'Export Data/Subsets', 'Link Data', 'IPython Terminal', 'Open Session', 'Export Session', and 'Add/edit arithmetic attributes'. The 'Data Collection' panel on the left lists two data sources: 'meth_cube_hdrfixed' and 'onh2d_cube_hdrfixed', with the latter selected. Below this are 'Plot Layers' and 'Plot Options' panels, both currently empty. The central canvas area is a large white space with the text 'Drag Data To Plot' centered in a light gray font. A 'Tab 1' label is visible in the top right corner of the canvas area. The bottom left corner of the image features the text 'An ALMA core' in a bold, black font.

python File Edit View Canvas Data Manager Plugins Help

Glue Jorma Harju

Open Data Export Data/Subsets Link Data IPython Terminal Open Session Export Session Add/edit arithmetic attributes Selection Mode: Preferences

Data Collection

Data

- meth_cube_hdrfixed
- onh2d_cube_hdrfixed

Subsets

Plot Layers

Plot Options

Drag Data To Plot

Tab 1

An ALMA core

Just drag to visualize, e.g. series of 2D "channel maps."

The screenshot shows the Glueviz application window. The top menu bar includes 'python', 'File', 'Edit', 'View', 'Canvas', 'Data Manager', 'Plugins', and 'Help'. The top toolbar contains various icons for file operations and data management. Below the toolbar, there are buttons for 'Open Data', 'Export Data/Subsets', 'Link Data', 'IPython Terminal', 'Open Session', 'Export Session', and 'Add/edit arithmetic attributes'. On the right side of the toolbar, there are 'Selection Mode' options and a 'Preferences' button. The main interface is divided into three vertical panels on the left: 'Data Collection', 'Plot Layers', and 'Plot Options'. The 'Data Collection' panel is active and shows a list of data sources: 'meth_cube_hdrfixed' (selected) and 'onh2d_cube_hdrfixed'. The 'Plot Layers' and 'Plot Options' panels are currently empty. The large central plot area is mostly blank, with the text 'Drag Data To Plot' centered in a large, light gray font. A tab labeled 'Tab 1' is visible at the top of the plot area.

An ALMA core

Adjust so each tracer is a different color.

python File Edit View Canvas Data Manager Plugins Help

Open Data Export Data/Subsets Link Data IPython Terminal Open Session Export Session Add/edit arithmetic attributes Selection Mode: Preferences

Data Collection

Data

- meth_cube_hdrfixed
- onh2d_cube_hdrfixed

Subsets

Plot Layers - 2D Image

- meth_cube_hdrfixed (PRIMARY)

attribute: PRIMARY

limits: Custom Arcsinh

0 1.1412

color/opacity: [red color swatch] Sync

contrast/bias: [sliders] Reset

Plot Options - 2D Image

General Limits Axes

mode: One color per layer

aspect: Square Pixels

reference: meth_cube_hdrfixed

x axis: Right Ascension

y axis: Declination

Vrad: Show real coordinates

4300.0 m/s

2D Image

methanol

Declination

Right Ascension

2D Image

Declination

Right Ascension

o-NH₂D

Create 3D views...

python File Edit View Canvas Data Manager Plugins Help

Open Data Export Data/Subsets Link Data IPython Terminal Open Session Export Session Add/edit arithmetic attributes Selection Mode: Preferences

Data Collection

Data

- meth_cube_hdrfixed
- onh2d_cube_hdrfixed

Subsets

Plot Layers - 3D Volume Rendering

- meth_cube_hdrfixed

Attribute: PRIMARY

Limits: 0 1.1412

Color: [red color swatch]

Plot Options - 3D Volume Rendering

x axis: Pixel Axis 2 [x]

min/max: -0.5 511.5

stretch: 1.00

y axis: Pixel Axis 1 [y]

min/max: -0.5 511.5

stretch: 1.00

z axis: Pixel Axis 0 [z]

min/max: -0.5 12.5

stretch: 10.00

resolution: 256

- Native aspect ratio
- Perspective
- Show axes
- Downsample when panning

2D Image

methanol

Declination

Right Ascension

3D Volume Rendering

Pixel Axis 0 [z]

Pixel Axis 1 [y]

Pixel Axis 2 [x]

...see clearly "veil" of wind-blown methanol.

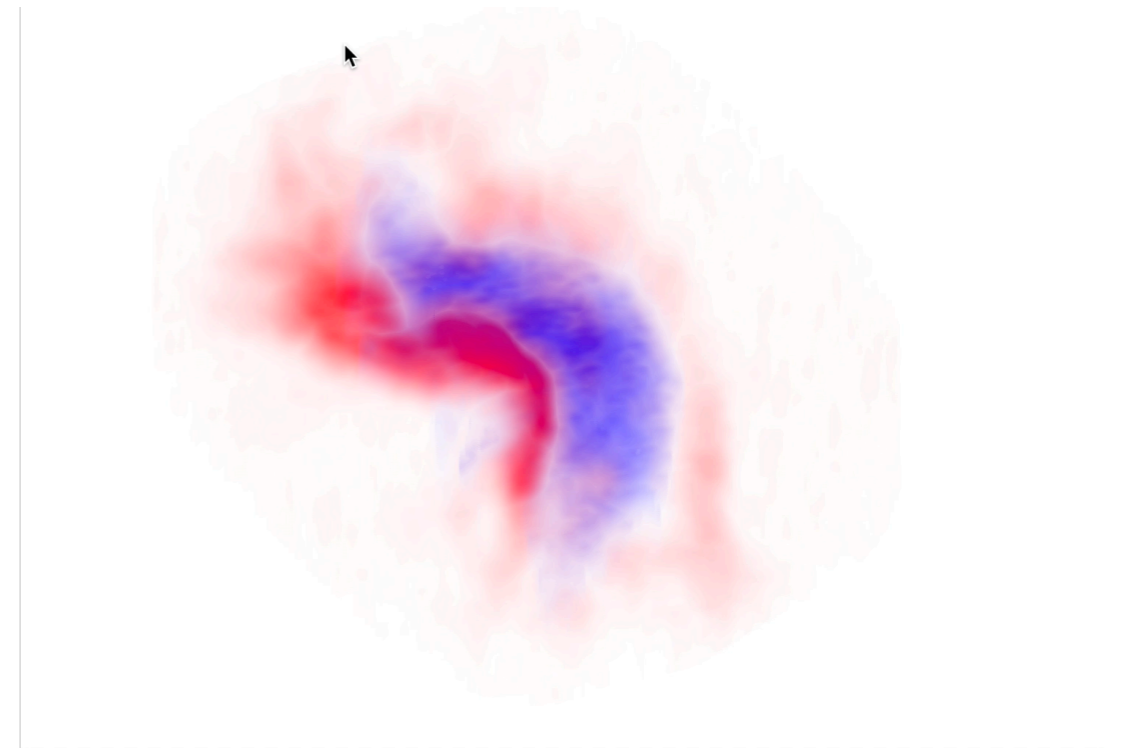
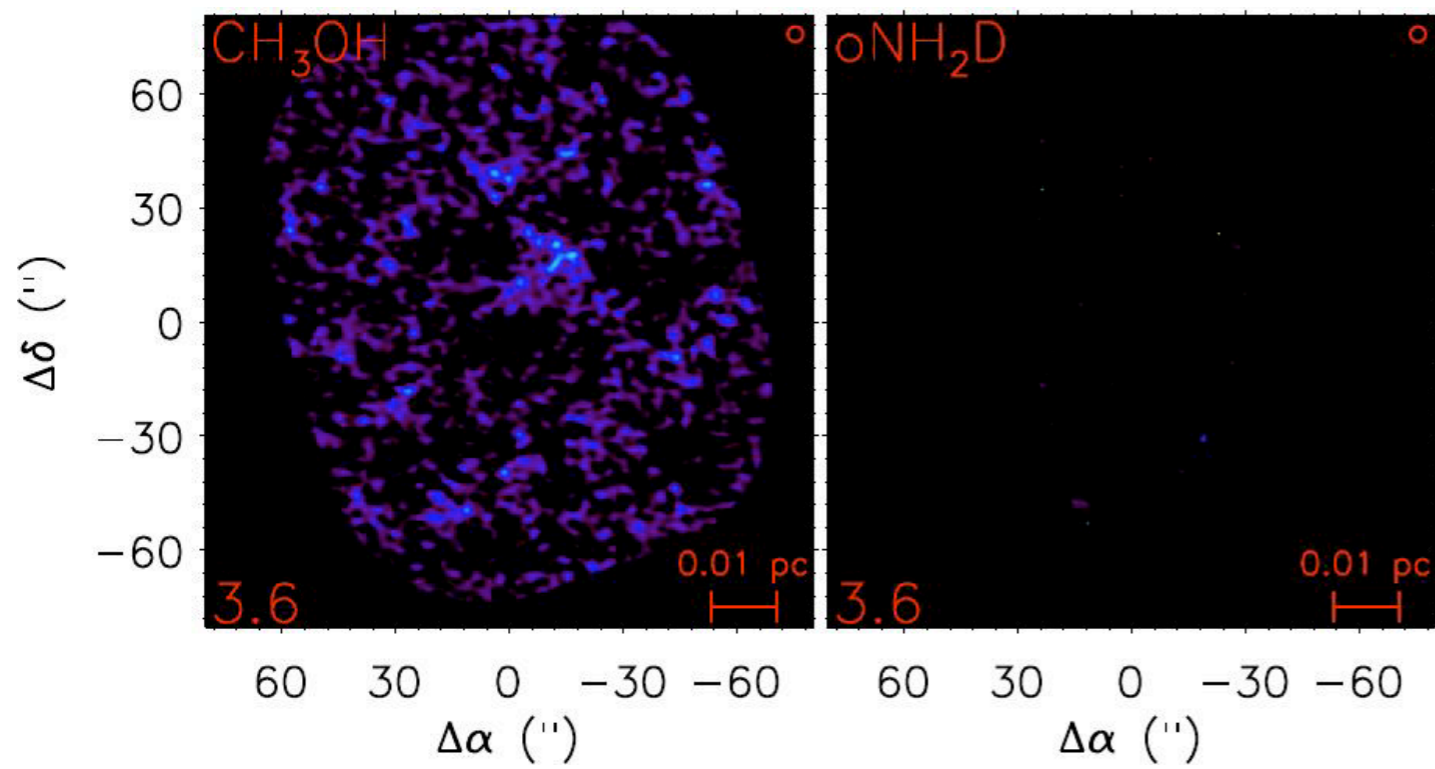
Declination

Right Ascension

o-NH₂D

Traditional Rainbow Channel maps

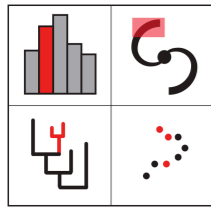
glue



2015: The "Paper" of the Future

The screenshot shows the Authorea website interface. At the top, the Authorea logo is on the left, and navigation links for 'FEATURED ARTICLES', 'ABOUT', 'PLANS', 'BLOG', 'FEEDBACK', 'HELP', and 'ALYSSA GOODMAN' are on the right. Below the navigation, there are status indicators for 'PUBLIC' and 'ROUGH DRAFT', and utility links for 'Index', 'Settings', 'Fork', 'Quickedit', 'Word Count', '42 Comments', 'Export', and 'Unfollow'. The main content area features the title 'The "Paper" of the Future' with a list of authors: Alyssa Goodman, Josh Peek, Alberto Accomazzi, Chris Beaumont, Christine L. Borgman, How-Huan Hope Chen, Merce Crosas, Christopher Erdmann, August Muench, Alberto Pepe, and Curtis Wong. There are buttons for 'Add author' and 'Re-arrange authors'. A note mentions a 5-minute video demonstration available at a YouTube link. The text begins with a '1 Preamble' section, discussing human cognition and the future of scholarly communication. A large image at the bottom shows a brain with the word 'Cognition' and the title 'Paper of the Future'. On the right, a comment section shows three comments from Konrad Hinsien, Merce Crosas, and Christine L. Borgman.





glue

multidimensional data exploration

enabled by d3.js (javascript) outputs



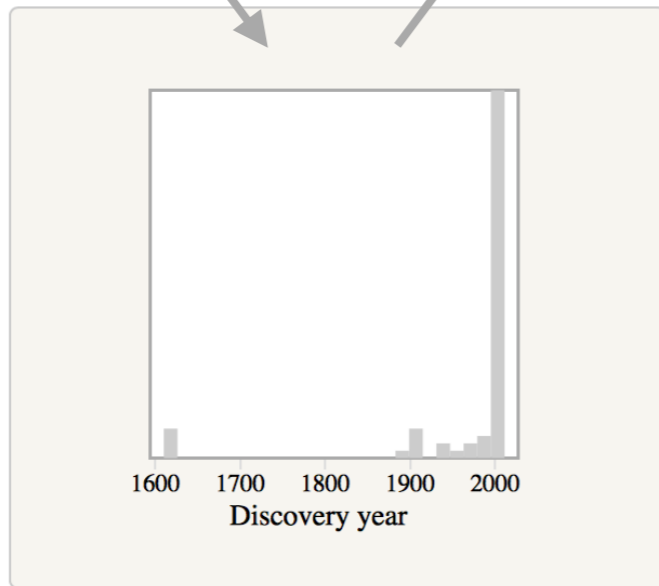
d3po

d3po is a project designed to allow an astronomer (or anyone), with no special data visualization skills, to make an interactive, publication-quality figure that has staged builds and linked brushing through scatter plots. Our current version can be previewed at d3po.org, and represents a figure from upcoming work by graduate student Elisabeth Newton. The figure describes how metallicity affects color in cool stars, and represents a nice use case for d3po. Try clicking and dragging in the scatter plots to understand the power of linked brushing in published figures.

Right now we are in search of alpha testers, who have figures that could be made interactive and who are willing to get their hands a little dirty (No javascript skills needed). In future versions, we plan to link to [glue](#) to allow the creation of d3po figures interactively. We are also exploring [implementation](#) of d3po within presentations and within [authorea](#). Full 1.0 version expected in January 2014.

Installing your own d3po server

```
git clone git@github.com:adrm/d3po.git
cd d3po
virtualenv --no-site-packages venv
source venv/bin/activate
pip install -r pip-requirements.txt
python run.py
```



- Four Centuries of Discovery
- A Chasm in Mass
- Little Siblings
- Close Cousins
- The Strangers

After Galileo discovered the first four moons of Jupiter, it took nearly three hundred years to discover the next one.

Authorea

The "Paper" of the Future

Authorea preprint 02/21/2017 DOI: 10.22541/au.148769949.92783646

Alyssa Goodman (Harvard University)
Josh Peek (Space Telescope Science Institute)
Alberto Accomazzi (Harvard-Smithsonian Center for Astrophysics (CFA))
Chris Beaumont (Harvard-Smithsonian Center for Astrophysics (CFA))
Christine L. Borgman (UCLA - University of California, Los Angeles)
Hope How-Huan Chen (Harvard University)
Merce Crosas (Harvard University)
Christopher Erdmann (North Carolina State University)

And 3 more...

A 5-minute video demonstration of this paper is available at [this YouTube link](#).

1 Preamble

A variety of research on human cognition demonstrates that humans learn and communicate best when more than one processing system (e.g. visual, auditory, touch) is used. And, related research also shows that, no matter how technical the material, most humans also retain and process information best when they can put a narrative "story" to it. So, when considering the future of scholarly communication, we should be careful not to do blithely away with the linear narrative format that articles and books have followed for centuries: instead, we should enrich it.

Much more than text is used to communicate in Science. Figures, which include images, diagrams, graphs, charts, and more, have enriched scholarly articles since the time of Galileo, and ever-growing volumes of data underpin most scientific papers. When scientists communicate face-to-face, as in talks or small discussions, these figures are often the focus of the conversation. In the best discussions, scientists have the ability to manipulate the figures, and to access underlying data, in real-time, so as to test out various what-if scenarios, and to explain findings more clearly. **This short article explains—and shows with demonstrations—how scholarly "papers" can morph into long-lasting rich records of scientific discourse**, enriched with deep data and code linkages, interactive figures, audio, video, and commenting.

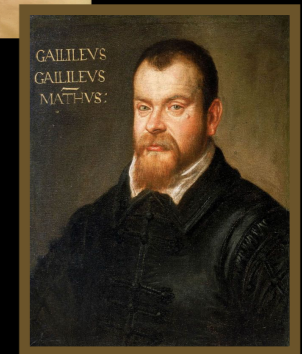
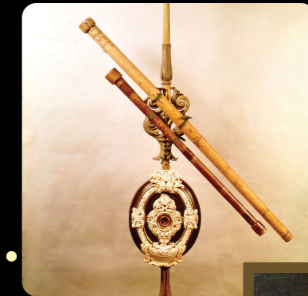
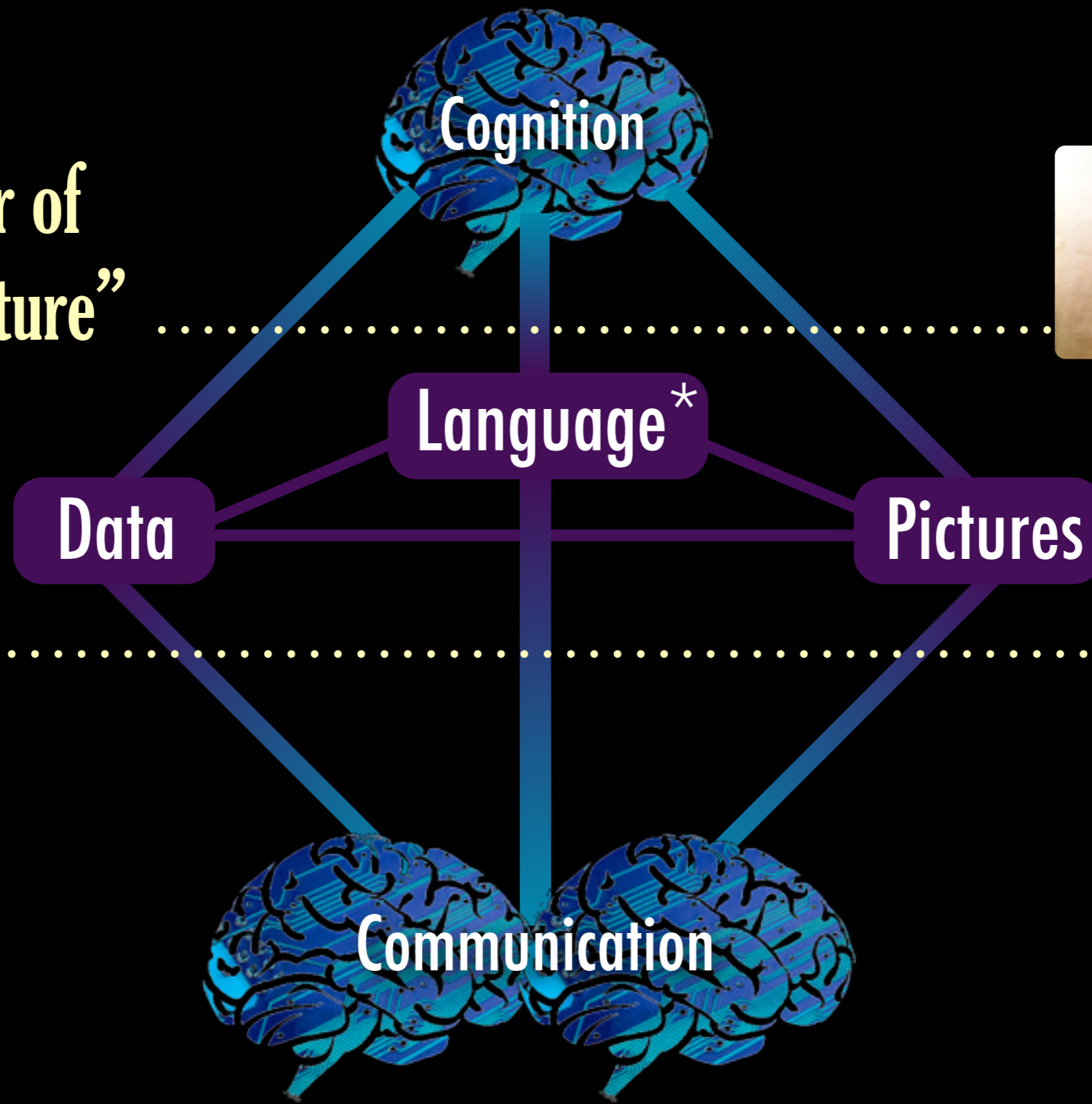
Fig. 1

The Paper of the Future should include seamless linkages amongst **data**, **pictures**, and **language**, where "language" includes both words and math. When an individual attempts to understand each of these kinds of information, different cognitive functions are utilized: communication is inefficient if the channel is restricted primarily to language, without easy interconnection to data and pictures.

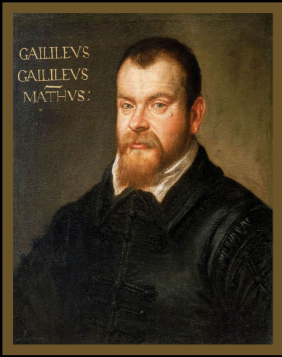
[demo]

Many thanks to Alberto Pepe, Josh Peek, Chris Beaumont, Tom Robitaille, Adrian Price-Whelan, Elizabeth Newton, Michelle Borkin & Matteo Cantiello for making this possible.

“Paper of
the Future”

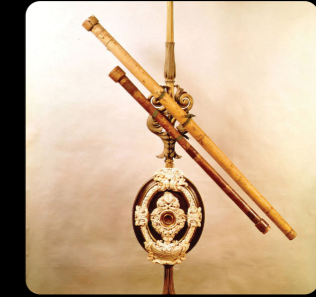


*“Language” includes words & math



1610: Why Galileo is My Hero

(Explore-Explain-Explore)



Sc. Principale.

Galileo Galilei, Familiari. Seruo della Ser. V. inuigilanza
 do assistendo, et de ogni spirito se benero no solo satisfare
 aluano che non della letura de Mathematici nelle stu-
 dio di Padoua,

Inuice dauere determinato di presentare al Sc. Principale
 l'Orchiale et il p. essere di giouamento inestimabile di ogni
 negocio et in iuncta marittima o terrestre stimo di tenere quel-
 sto nuovo artificio ne l' maggior segreto et solam a disposizione
 di V. Ser. L'Orchiale auato dalle piu u. di ite speculazioni di
 prospetua ha l'uantaggio di scoprire Legni et Vele dell' inimico
 di lue hore et pu di tempo prima che egli sia sopra noi et distinguendo
 il numero et la qualita de i Vasselli giudicare la sua forte
 pallesirsi alla caccia al combattimento o alla fuga, o pure esser
 nella campagna aperta uedere et particolarmente distinguere ogni suo
 moto et preparatione.

Adi 7. di Gennaio
 Giove si uede usti

Adi 8. usti

Adi 11. si uede in tale uisione

Adi 13. si uede uasi in Giove 4 stelle

Adi 14. è angelo

Adi 15. si uede la pross. a 4 ora in m. la 4. ora di =
 stante dalla 3. il gruppo la 1. =

La spatio delle 3 uidevoli ad ora
 maggiore del diametro di 7. et e =
 uasi in linea retta.

7. long. 28. lat. 115

East		West	
7	* * ○ *	17	* ○
8	○ * * *	18	* ○ *
10	* * ○	19	* ○ * *
11	* * ○	19	* ○ * *
12	* ○ *	20	○ * ○ ○
13	* ○ * *	21	... ○
15	○ * * *	22	* ○ * *
15	○ * *	22	* ○ * *
16	* ○ *	23	* ○ *
17	* ○ *	24	* ○ *

SIDEREUS NUNCIUS 75

On the third, at the seventh hour, the stars were arranged in this sequence. The eastern one was 1 minute, 30 seconds from Jupiter; the closest western one 2 minutes; and the other western one was 10 minutes removed from this one. They were absolutely on the same straight line and of equal magnitude.

On the fourth, at the second hour, there were four stars around Jupiter, two to the east and two to the west, and arranged precisely on a straight line, as in the adjoining figure. The easternmost was distant 3 minutes from the next one, while this one was 40 seconds from Jupiter; Jupiter was 4 minutes from the nearest western one, and this one 6 minutes from the westernmost one. Their magnitudes were nearly equal; the one closest to Jupiter appeared a little smaller than the rest. But at the seventh hour the eastern stars were only 30 seconds apart. Jupiter was 2 minutes from the nearer eastern one, while he was 4 minutes from the next western one, and this one was 3 minutes from the westernmost one. They were all equal and extended on the same straight line along the ecliptic.

On the fifth, the sky was cloudy.

On the sixth, only two stars appeared flanking Jupiter, as is seen in the adjoining figure. The eastern one was 2 minutes and the western one 3 minutes from Jupiter. They were on the same straight line with Jupiter and equal in magnitude.

On the seventh, two stars stood near Jupiter, both to the east, arranged in this manner.

Notes for & re-productions of Siderius Nuncius

[More later?]

2018: 10QViz.org

≡ MENU



TEN QUESTIONS TO ASK WHEN CREATING A VISUALIZATION

The 10 Questions

1. **Who** | Who is your audience? How expert will they be about the subject and/or display conventions?
2. **Explore-Explain** | Is your goal to explore, document, or explain your data or ideas, or a combination of these?
3. **Categories** | Do you want to show or explore pre-existing, known, human-interpretable, categories?
4. **Patterns** | Do you want to identify new, previously unknown or undefined patterns?
5. **Predictions & Uncertainty** | Are you making a comparison between data and/or predictions? Is representing uncertainty a concern?
6. **Dimensions** | What is the intrinsic number of dimensions (not necessarily spatial) in your data, and how many do you want to show at once?
7. **Abstraction & Accuracy** | Do you need to show all the data, or is summary or abstraction OK?
8. **Context & Scale** | Can you, and do you want to, put the data into a standard frame of reference, coordinate system, or show scale(s)?
9. **Metadata** | Do you need to display or link to non-quantitative metadata? (including captions, labels, etc.)
10. **Display Modes** | What display modes might be used in experiencing your display?

 **Join the 10QViz Conversation!** 

To learn more about this site, please visit the [About](#) page.

To read an in-process manuscript giving the scholarship behind the recommendations on this site, see [Coltekin & Goodman 2018](#).

2019: glue in the browser

The screenshot displays the JupyterLab interface with several components:

- Code Editor:** Contains the command `app.scatter3d('x', 'y', 'z');` and a mode selector with options like 'lasso+', 'circle', and 'rect'.
- Output View:** Shows a scatter plot with a 'brush x' mode selected.
- Modal Menu:** A 'Start a new activity' dialog box is open, offering options for 'Notebook', 'Code Console', and 'Text Editor'.
- Visualizations:** A 2D histogram at the top left and a 3D scatter plot at the bottom left. A 2D histogram at the bottom right shows a distribution of data points with a green vertical highlight.

The 'gluepyter' logo is visible in the bottom left corner, featuring icons for a histogram, a scatter plot, and a code editor.

Video courtesy of Maarten Breddels, consulting developer

What anyone can do right now



What computer
scientists do

What scientists
do



What we should do?

3D PDF
2009

PotF
2015

10QViz
2018

glupyter
2019

+ more, together

+ more, together

Help us create a database of astro-referenced old Astronomy images.

Learn more Get started

1 person is talking about Astronomy Rewind right now.

Join in

ZOONIVERSE

WORDS FROM THE RESEARCHER

"Your contributions unlock the information from old astronomy journals. Thank you and enjoy the images!"

ABOUT ASTRONOMY REWIND

This project is part of an ongoing NASA-funded effort aimed at turning the SAO/NASA Astrophysics Data System (ADS) into a data resource. The result will be a database of astro-referenced images, i.e., images of the sky for which coordinates, orientation, and pixel scale will be publicly available through NASA data archives, the [Astronomy Image Explorer](#), and [WorldWideTelescope](#), thanks to your help!

ZOONIVERSE

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- How to Build
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- Acknowledgements
- Contact Us
- Zooniverse Talk
- Daily Zooniverse
- Blog

Privacy Policy | Jobs | System Status | Security

Custom Parts Organizer Box Included!

glue

Plot Layers - 1D Histogram

ROOTNAME	SCI	A,B,2	A,0
0 ic9-04cve	SCI	9.46198778625...	1.903870
1 ic9-04cve	SCI	1.0648009211...	2.360288
2 ic9-04cve	SCI	9.46198778625...	1.903870
3 ic9-04cve	SCI	1.0648009211...	2.360288
4 ic9-04d0c	SCI	9.46198778625...	1.903870
5 ic9-04d0c	SCI	1.0648009211...	2.360288
6 ic9-04d4c	SCI	9.46198778625...	1.903870
7 ic9-04d4c	SCI	1.0648009211...	2.360288

VAST • INFOVIS • SCIVIS

VIS 2018

21-26 October 2018
BERLIN, GERMANY

Plot Layers - 3D Volume Rendering

Attribute: intensity

Min: -0.66 Max: 16

Color: [gradient]

Alpha: [slider]

Plot Options - 3D Volume Rendering

X axis: min 0 max 105 stretch 1.06

Y axis: min 0 max 105 stretch 0.94

Z axis: min 0 max 53 stretch 0.53

Log axis labels in 1D histograms #1597

agoodman opened this issue a minute ago · 0 comments

agoodman commented a minute ago

The .glu file that can reproduce the behavior below is at <https://www.dropbox.com/5nVOp1wcd8pax5gwaACQ4v88Vp75oDtkb9hms?dl=0>.

What happens? When you make a histogram & ask for (both, or one) log axes, the tick labeling is not right, as shown below.

Miraculously, if you save the file as a .glu session, the situation improves somewhat. The same plot shown above opens with the x-axis labels fixed, but the y-axis labels no longer "log" and in fact just oddly wrap!

But then, if you click and off a few times below.

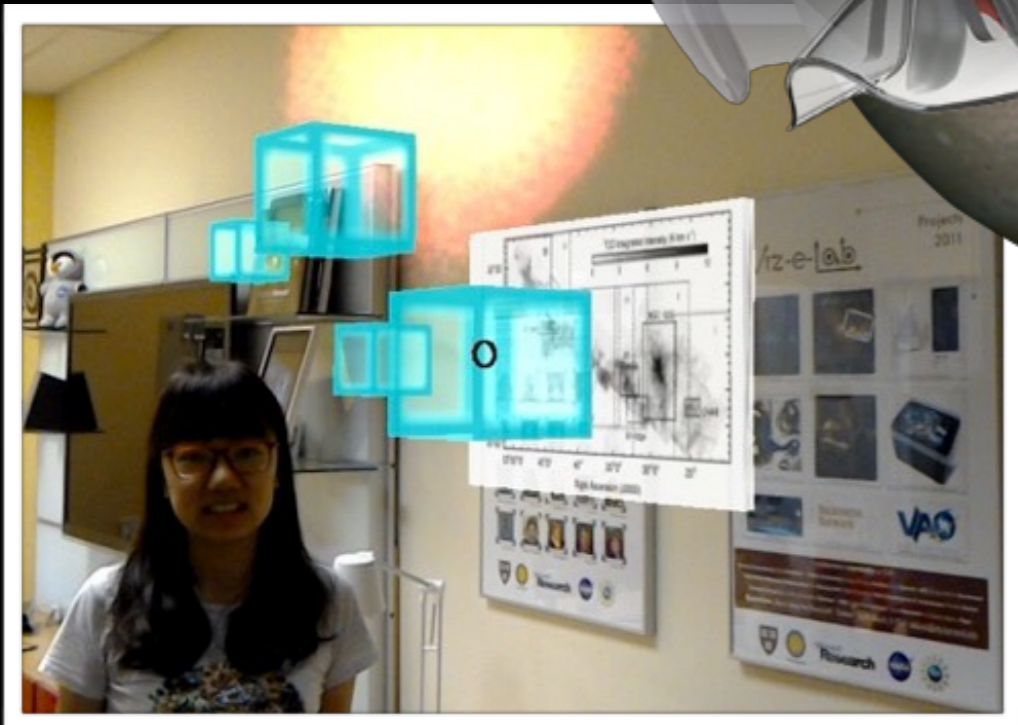
Oddly, at some point of all the on/off clicking, they can also look like this.

Write Preview

Attach files by dragging & dropping or selecting them.

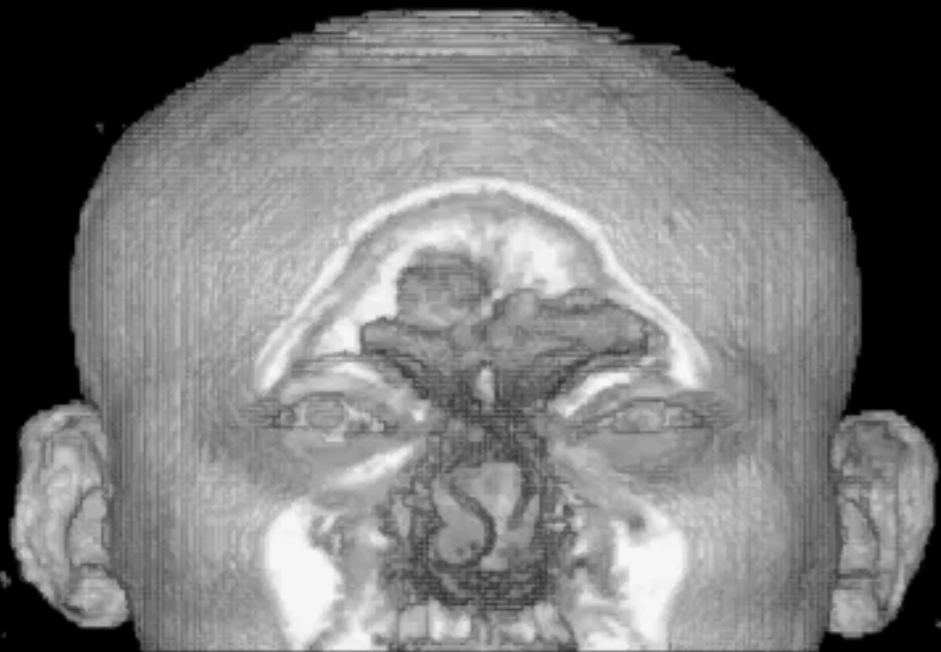
Close issue Comment

The challenge of 3D Selection



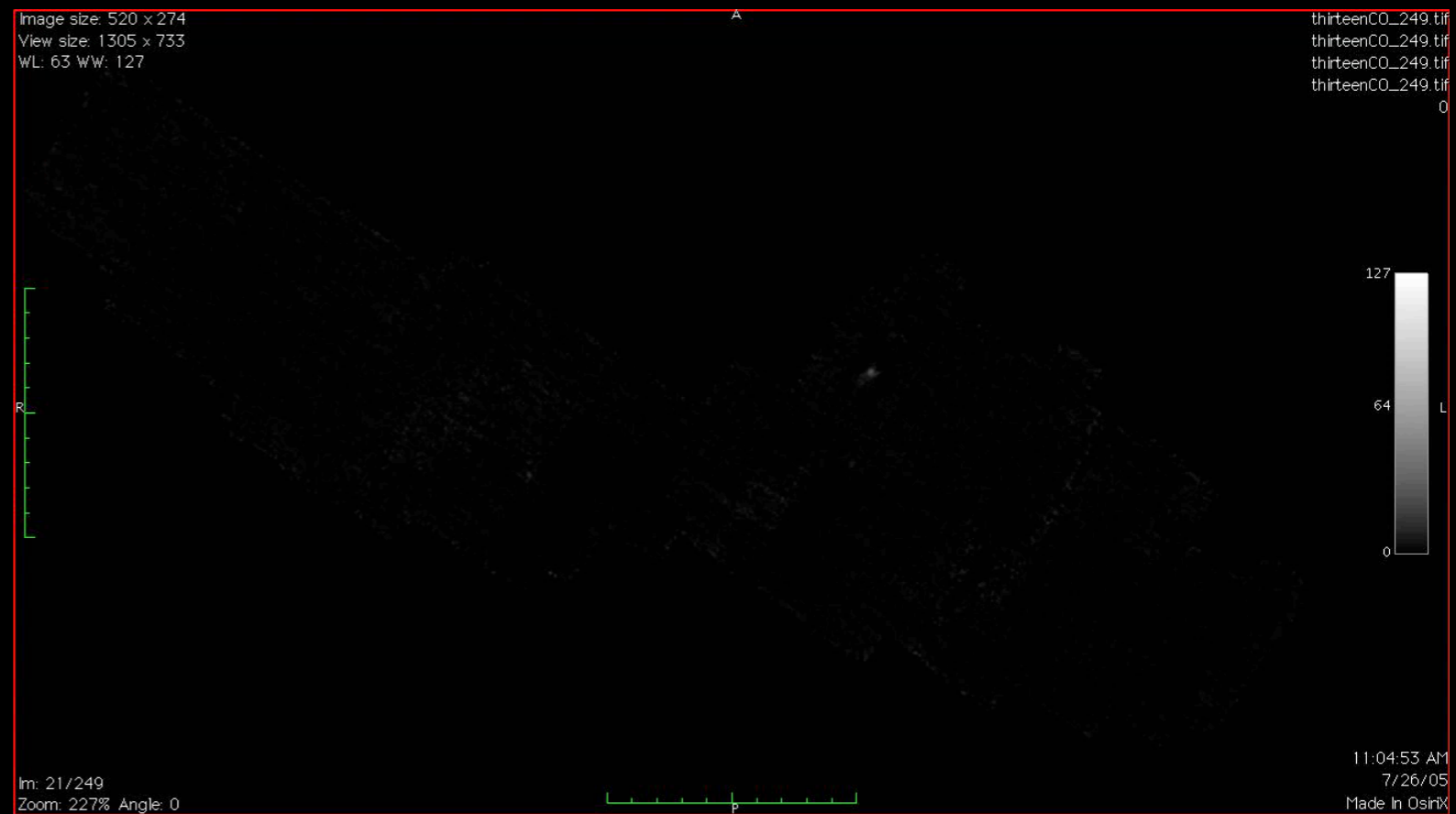
ASTRONOMICAL MEDICINE

"KEITH"



"z" is depth into head






"PERSEUS"

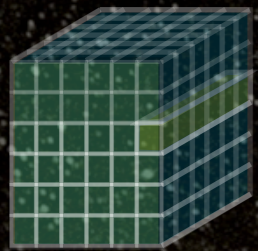


"z" is line-of-sight velocity

Image size: 520 x 274
View size: 1305 x 733
W/L: 63 WW: 127

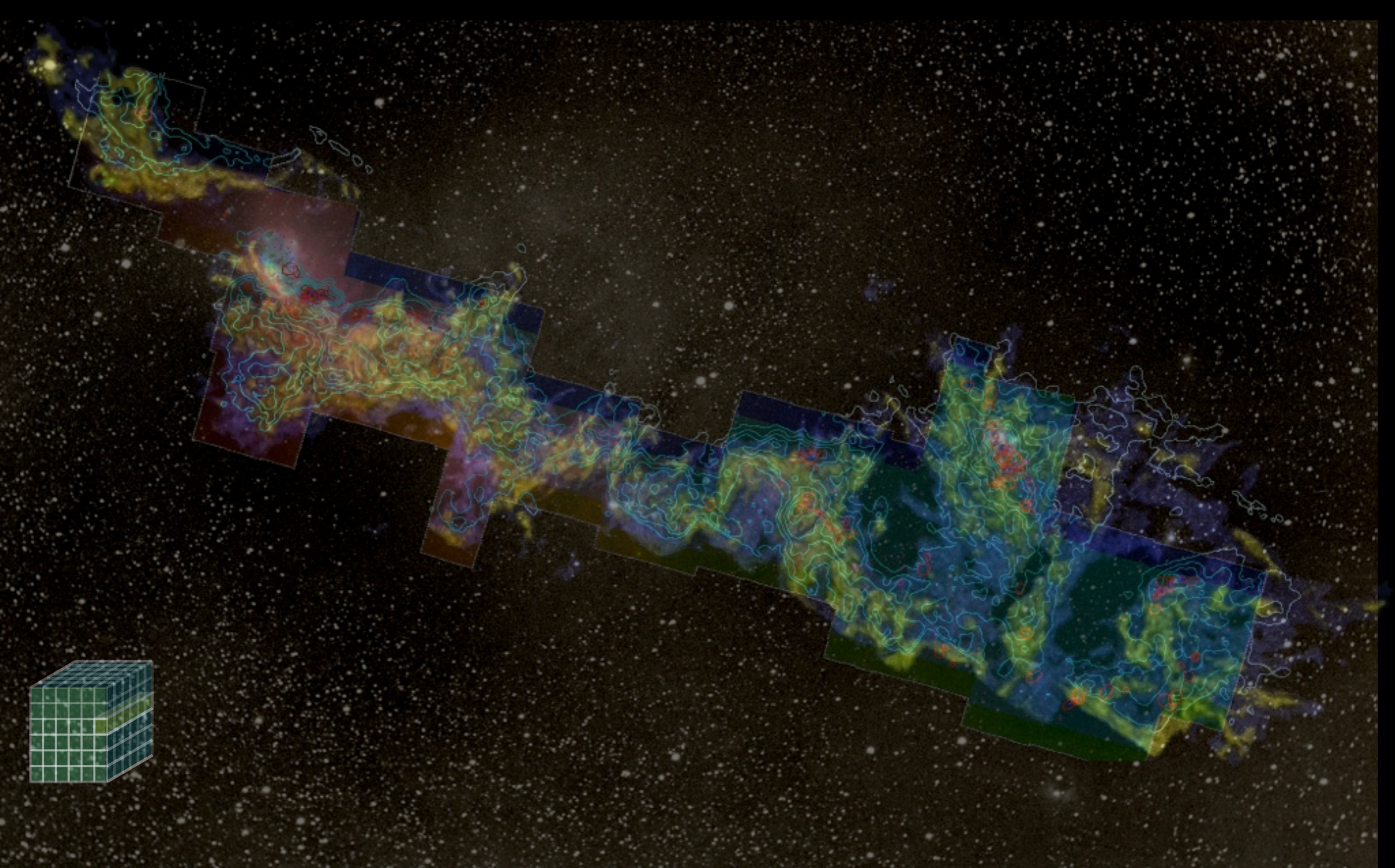
ASTRONOMICAL MEDICINE

-  mm peak (Enoch et al. 2006)
-  sub-mm peak (Hatchell et al. 2005, Kirk et al. 2006)
-  ^{13}CO (Ridge et al. 2006)
-  mid-IR IRAC composite from c2d data (Foster, Laakso, Ridge, et al.)
-  Optical image (Barnard 1927)



m: 1/249
Zoom: 227% Angle: 0





3D Viz made with VolView

DATA,
CODE,
COLLABORATION



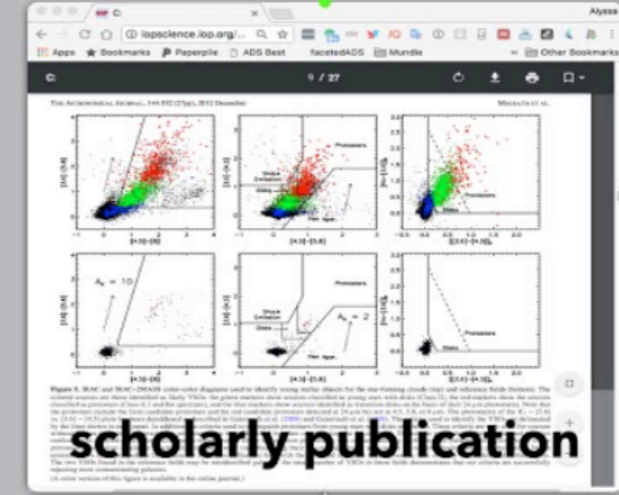
DATA-DRIVEN STORYTELLING



shared data



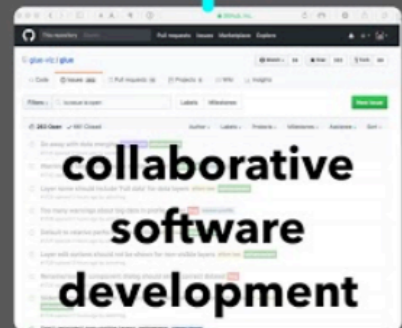
public outreach



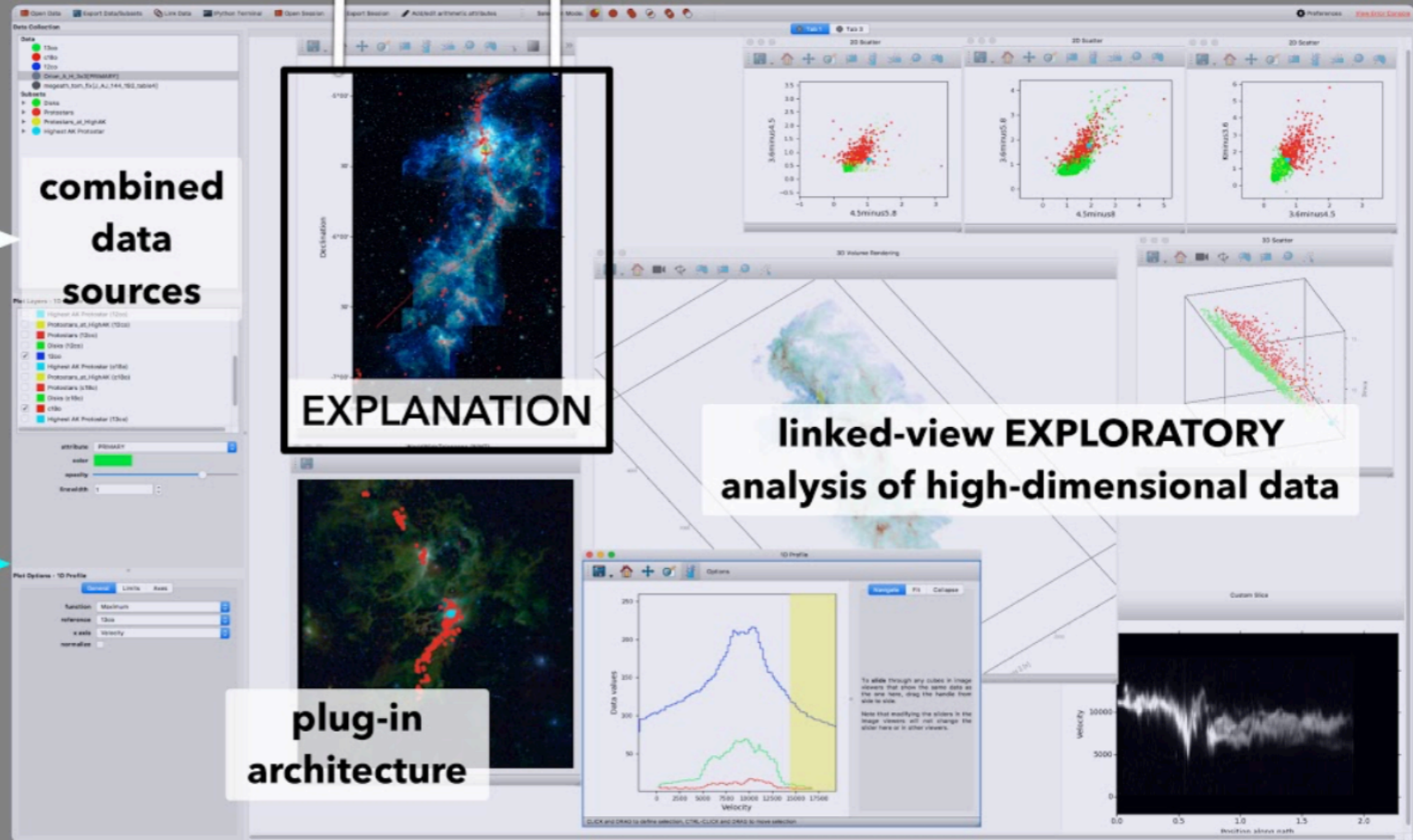
scholarly publication



open source,
modular,
software



collaborative
software
development



combined
data
sources

EXPLANATION

linked-view EXPLORATORY
analysis of high-dimensional data

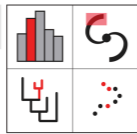
plug-in
architecture



EXPLORATION

EXPLANATION





glue
multidimensional data exploration

Linked-View Exploratory Visualization of High-Dimensional Data, for Everyone

Alyssa Goodman (PI, Harvard)
Michelle Borkin (PI, Northeastern)
Thomas Robitaille (Lead Architect)

The glue project was founded in 2012, with funding from NASA's James Webb Space Telescope (JWST) project. NASA contracts continue to support development of JWST-related (Astronomy) functionality.

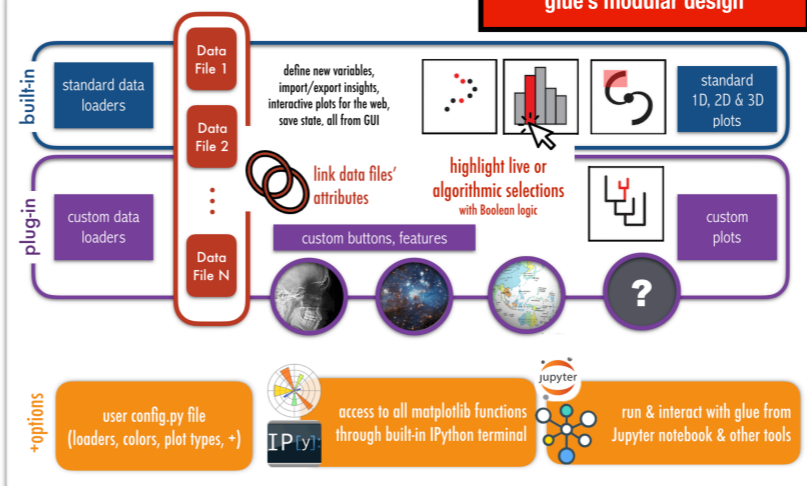
Beginning in 2017, glue has also been funded by the National Science Foundation, under SI2-SSE 1739657/1740229: Collaborative Research: A sustainable future for the glue multi-dimensional linked data visualization package. The goal of the NSF SSE funding is to expand glue's functionality into domains beyond its traditional strengths in Astronomy and Medicine, by broadening both its user and developer communities. All glue code is Open Source, at github.com/glue-viz

glueviz.org

github.com/glue-viz

glueviz.slack.com

glue's modular design



Linked Visualizations

With Glue, users can create scatter plots, histograms and images (2D and 3D) of their data. Glue is focused on the brushing and linking paradigm, where selections in any graph propagate to all others.



Flexible linking across data

Glue uses the logical links that exist between different data sets to overlay visualizations of different data, and to propagate selections across data sets. These links are specified by the user, and are arbitrarily flexible

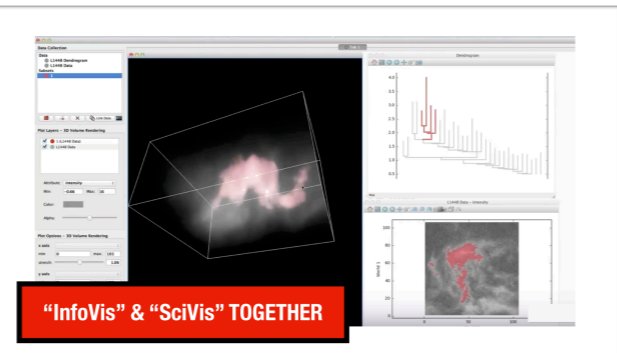


Full scripting capability

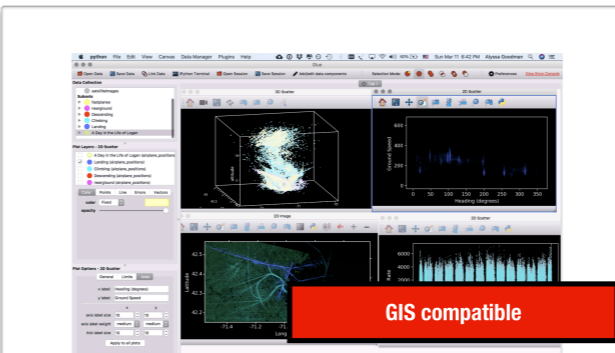
Glue is written in Python, and built on top of its standard scientific libraries (i.e., Numpy, Matplotlib, Scipy). Users can easily integrate their own python code for data input, cleaning, and analysis.

Want to plug-in your project or tool?
Consider joining us for **glue-con**, right after **JupyterCon**, August 27-29, 2018, at Harvard.

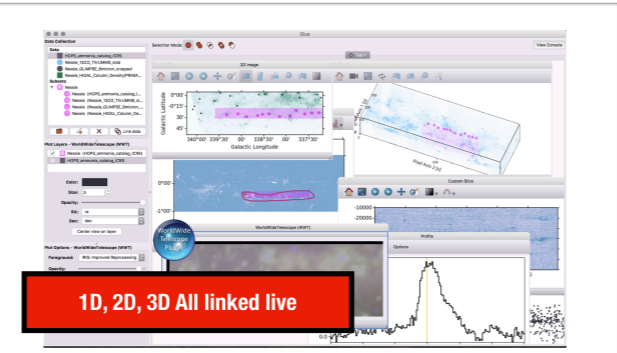
glue-con
2018, CAMBRIDGE, MA
projects.iq.harvard.edu/gluecon



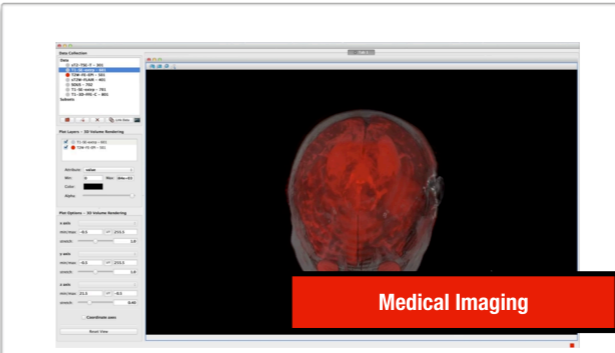
"InfoVis" & "SciVis" TOGETHER



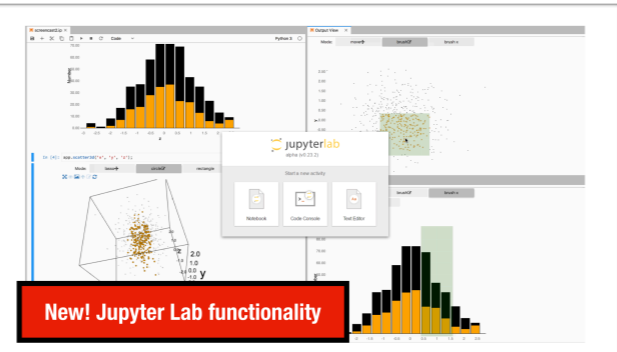
GIS compatible



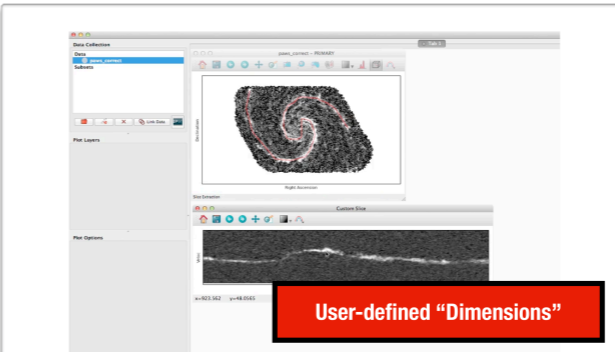
1D, 2D, 3D All linked live



Medical Imaging



New! Jupyter Lab functionality



User-defined "Dimensions"

Traditionally, travel from exploration to explanation is called “Scholarly Publishing” if its *dry*, and “Public Outreach,” if it’s *beautiful*.

Explore



Explain

Explore



Explain

It's much harder to go the other way.

Explore

And, the *best* roads are two-way.

Explain

Scripting

Glue

Open Data Export Data/Subsets Link Data IPython Terminal Open Session Export Session Add/edit arithmetic attributes Selection Mode: Preferences

Data Collection

Data

- x90y40_NH2_0deg
- x90y40_nH2
- x90y40_filmask_0deg

Subsets

- filmask_zero

Plot Layers - 3D Volume Rendering

- filmask_zero (x90y40_nH2)
- x90y40_nH2

Attribute: PRIMARY

Color: [Color Picker]

Subset: Data Outline

2D Image

3D Volume Rendering

```
~/Google Drive/Glue Stuff/SimFil_Glue] aagoodman% python3 simfil_startup.py 'x90y40'
```

x axis Pixel Axis 2 [x]

min/max: -28.5749 ⇌ 527.575

stretch: 1.00

y axis Pixel Axis 1 [y]

min/max: -28.5749 ⇌ 527.575

stretch: 1.00

z axis Pixel Axis 0 [z]

min/max: -28.5749 ⇌ 527.575

stretch: 1.00

resolution: 256

Native aspect ratio

Perspective Show axes

Downsample when panning