

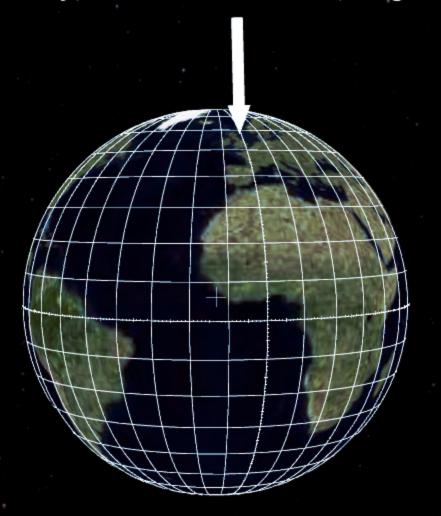
Latitude, Longitude, and



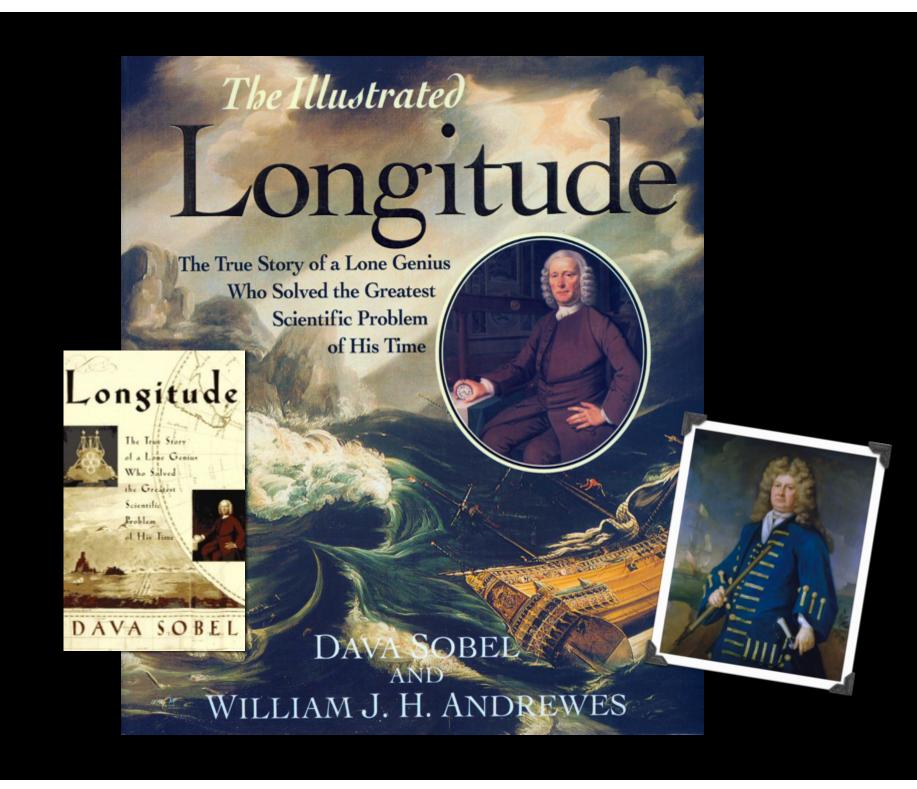
WorldWide Telescope



But Longitude has no natural "zero." Today, we use Greenwich, England.



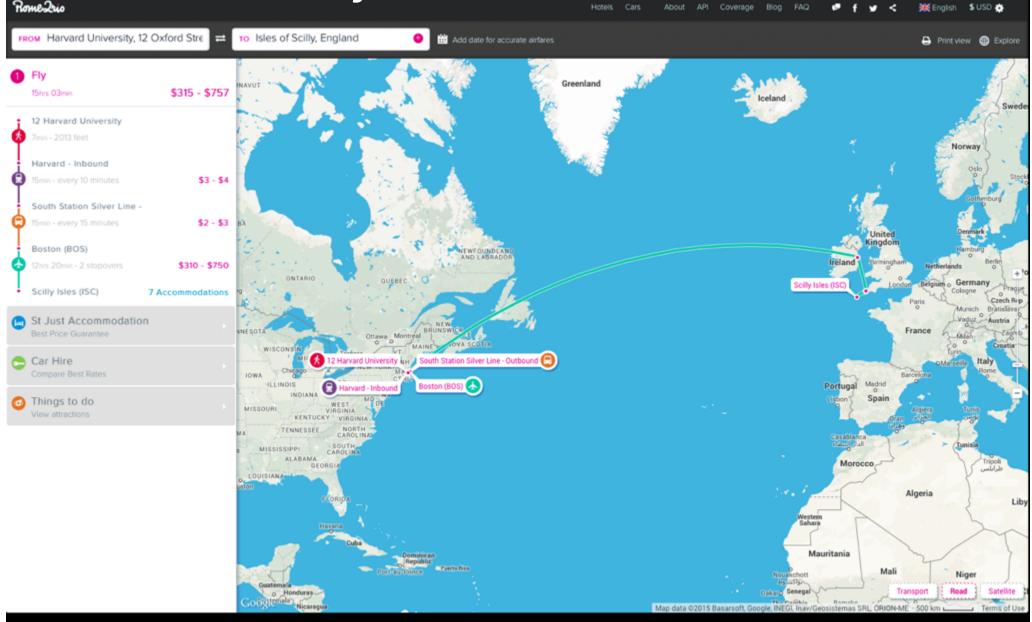


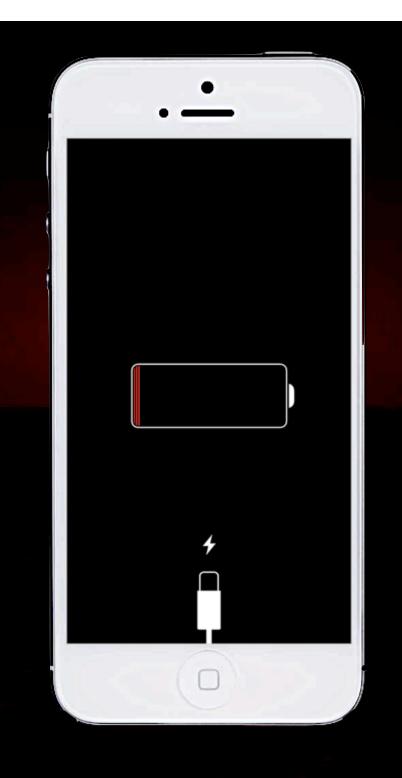




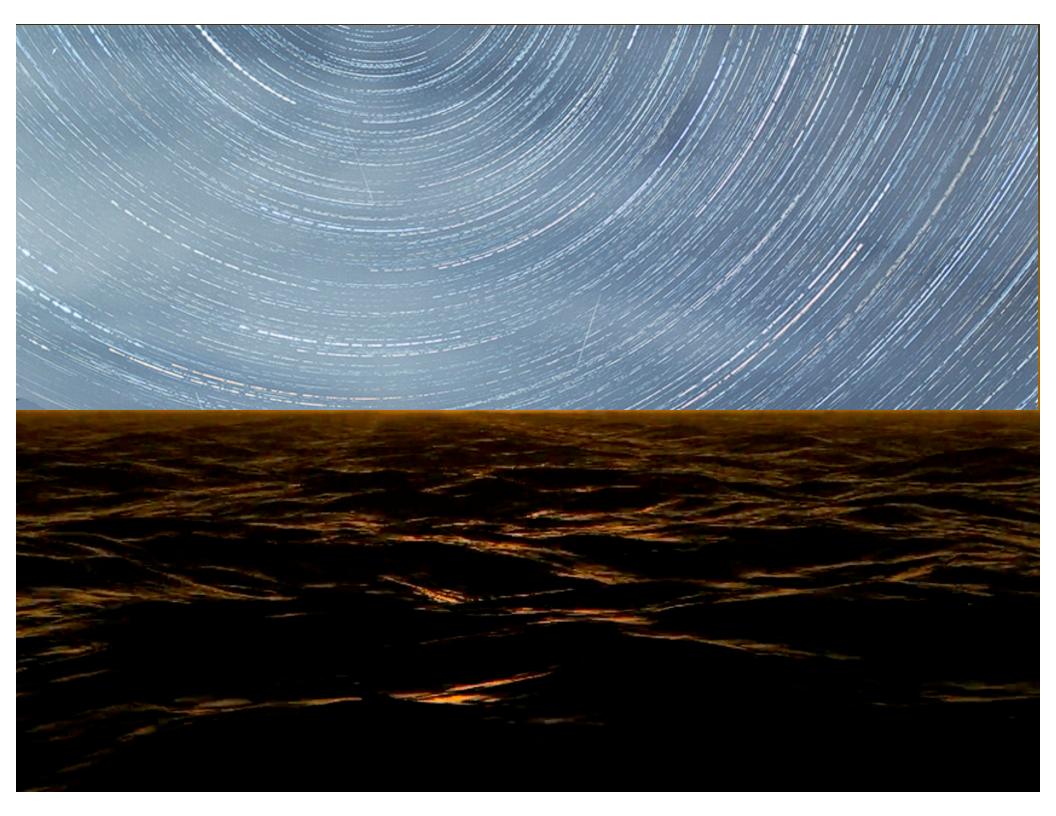


Isles of Scilly











Uncertainty

Accuracy of Maps

Direction

Measurement

Time

Measurement

Angular Measurement

Astronomical Measurement Measuring Speed









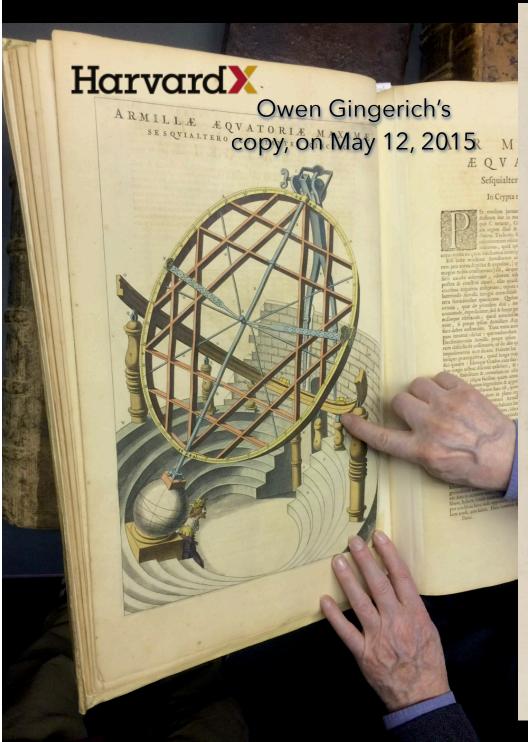
now

1200

1675

1750

2000





© Copyright Nederlands Scheepvaartmuseum Amsterdam www.geheugenvannederland.nl/?/en/items/NESA01xxCOLONxxL01-0370



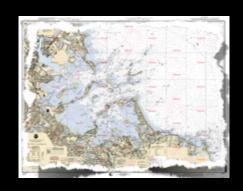




Dead Reckoning







"Once they were out of sight of land..."



from: The Clock That Changed the World (BBC History of the World)

presenter: Adam Hart-Davis; https://www.youtube.com/watch?v=T-g27KS0yiY

Latitude is "Easy"...

The Earth **spins** on its axis, once per ~24 hours.

Earth's **spin** causes distant stars to appear to go in **circular paths** (as if painted on to a very distant sphere).

Earth's spin axis is **tilted** by 23 degrees from the plane of its orbit.

Because of the tilt, the rise, **transit**, and set times for a star whose path dips below the horizon **depends** on where the Earth is in its orbit (**on date**).

The value of the maximum **altitude** (at transit) any star reaches on a particular day **depends on latitude**. (+Altitude of North Star=Latitude.)

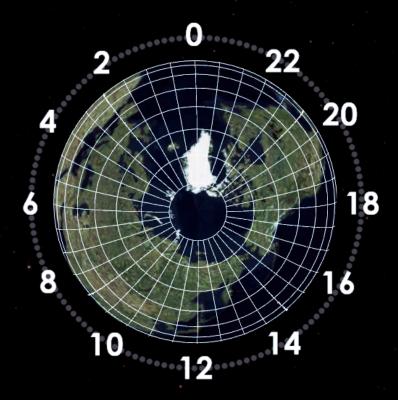
A star will reach **its maximum altitude ("transit")**, at the *same* (local) time on any given day, as viewed from any place with equal latitude.

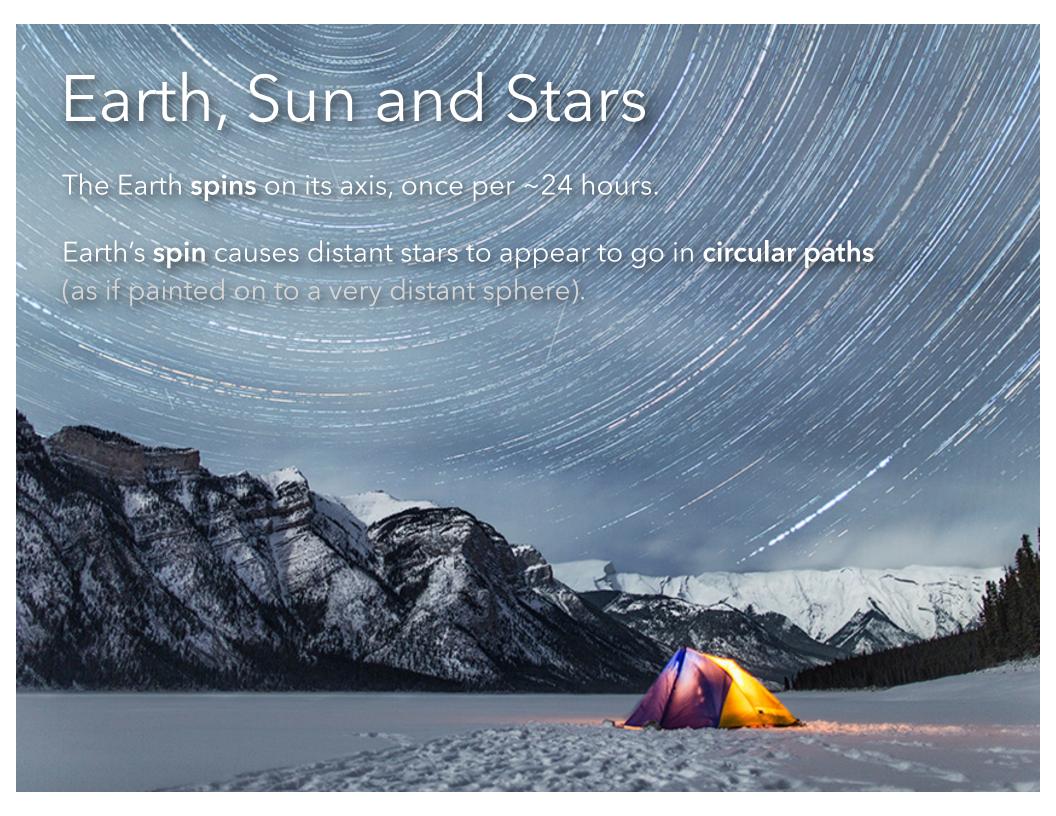
The Earth **spins** on its axis, once per ~24 hours.



1 Rotation per 24 Hours means

15 degrees of longitude=1 hour



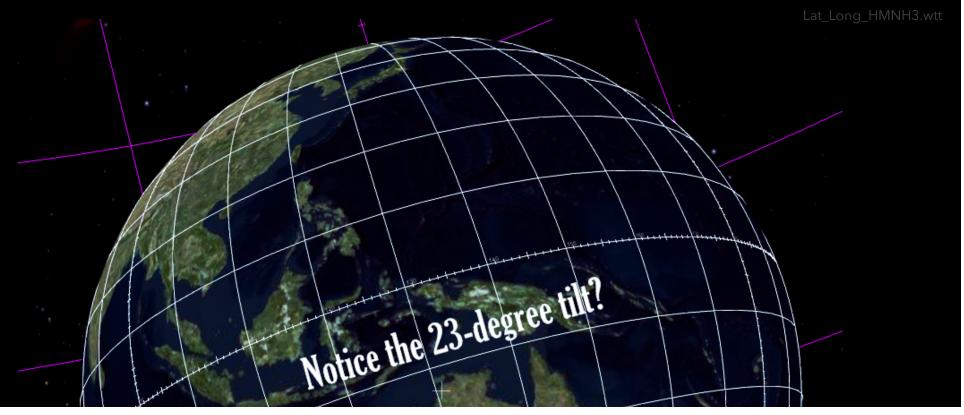


The Earth **spins** on its axis, once per ~24 hours.

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The Earth **spins** on its axis, once per ~24 hours.

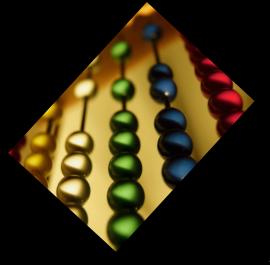
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"Sail the Parallel"







The Earth **spins** on its axis, once per ~24 hours.

Earth's **spin** causes distant stars to appear to go in **circular paths** (as if painted on to a very distant sphere).

Earth's spin axis is **tilted** by 23 degrees from the plane of its orbit.

Because of the tilt, the rise, **transit**, and set times for a star whose path dips below the horizon **depends** on where the Earth is in its orbit (**on date**).

The value of the maximum **altitude** (at transit) any star reaches on a particular day **depends on latitude**. (+Altitude of North Star=Latitude.)

A star will reach its maximum altitude ("transit"), at the same (local) time on any given day, as viewed from any place with equal latitude.

The value of the maximum **altitude** (at transit) any star reaches on a particular day **depends on latitude**. (+Altitude of North Star=Latitude.)

Great! We can easily use the stars to find latitude.

A star will reach **its maximum altitude ("transit")**, at the *same* (local) time on any given day, as viewed from any place with equal latitude.

Drat! We can only use the stars to find longitude if we know the time, very well.

Uncertainty → Tragedy

Accuracy of Maps

Direction

Measurement

Time

Measurement

Angular Measurement

Astronomical Measurement Measuring Speed









now

1200

1675

1750

2000

The Longitude Act



from: The Clock That Changed the World (BBC History of the World)

presenter: Adam Hart-Davis; https://www.youtube.com/watch?v=T-g27KS0yiY

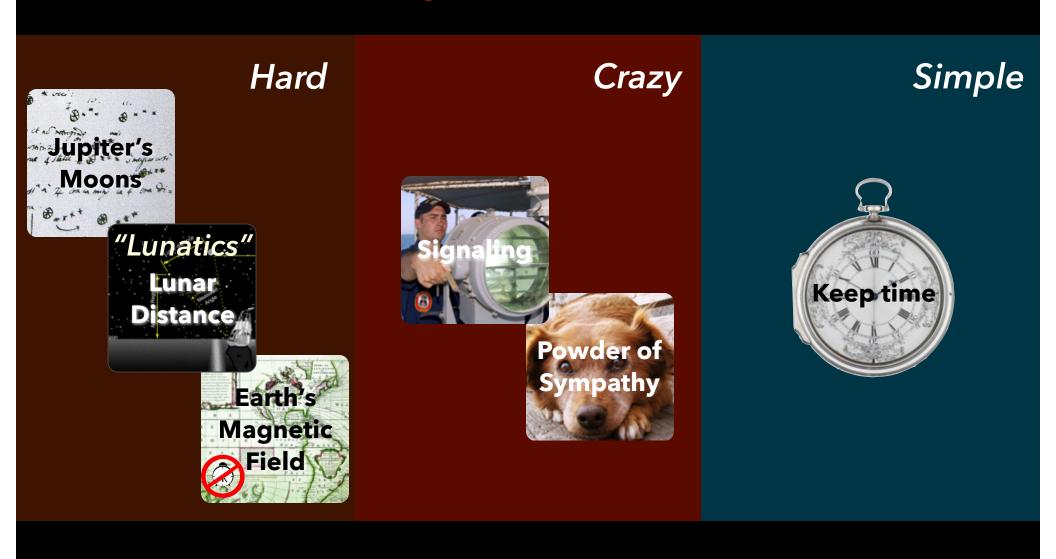
Measuring Astronomical Objects' Positions



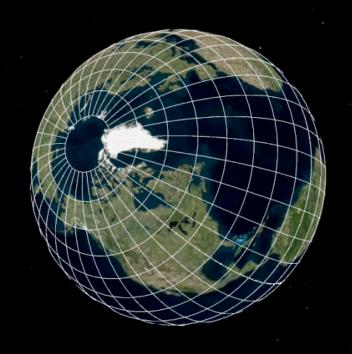
Videos Copyright 1992, The Mariners' Museum. featuring Captain Eric A. Speth, maritime program manager, of the Jamestown Yorktown foundation.

Finding Longitude

...use the stars to find longitude if we know the time, very well...



Time-based navigation relies on known speed... of Earth's rotation.







scene from "The Longitude", reproduced from YouTube, without permission https://www.youtube.com/watch?v=XUhhdwvLuyk





scene from "The Longitude", reproduced from YouTube, without permission https://www.youtube.com/watch?v=XUhhdwvLuyk

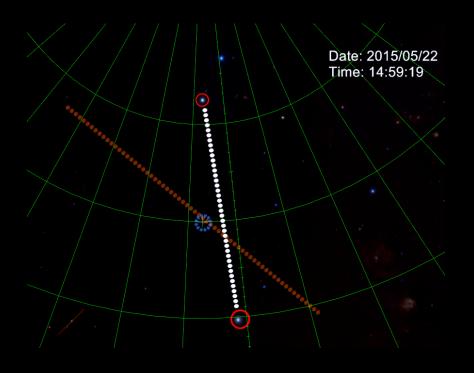




The "Lunar Distance Method"

To find time: Calculate the exact path of the moon, as a function of time, read off position, calculate time.

Complications: Relies on VERY accurate records, predictions, measurements and calculations.





lunars_hard.wtt

Galileo & Jupiter's Moons

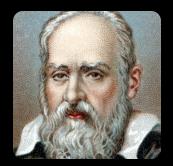


To find time: Calculate the positions of Jupiter's moons as a function of time. Observe moons, know time.

Complications: Relies on high-resolution observations, requiring a very stable observing platform (not a ship!!)







Galileo Galilei (1564-1642)



Jex Praye.
Golibo Galily Humlis Serus Della Ser. V. inwigitan
To senduame et to ogni chinio to buere no solam satisfas aliario che nera della cheum di Madamatito nello seu
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On the third, at the seventh hour, the stars were arranged in this quence. The eastern one was 1 minute, 30 seconds from Jupiter 2 closest western one 2 minutes; and the other western one was

minutes removed from this one. They were absolutely on the ame straight line and of equal magnitude.

On the fourth, at the second hour, there were four stars arour upiter, two to the east and two to the west, and arranged precise

east we

on a straight line, as in the adjoining figure. The easternmost walistant 3 minutes from the next one, while this one was 40 second from Jupiter; Jupiter was 4 minutes from the nearest western one of this one 6 minutes from the westernmost one. Their magnitudes are nearly equal; the one closest to Jupiter appeared a little smaller can the rest. But at the seventh hour the eastern stars were only of seconds apart. Jupiter was 2 minutes from the nearer eastern

st ** O * * West

me, while he was 4 minutes from the next western one, and this one was 3 minutes from the westernmost one. They were all equal and extended on the same straight line along the ecliptic.

On the fifth, the sky was cloudy.

On the sixth, only two stars appeared flanking Jupiter, as is seen

* O * West

in the adjoining figure. The eastern one was 2 minutes and the vestern one 3 minutes from Jupiter. They were on the same straight fine with Jupiter and equal in magnitude.

On the seventh, two stars stood near Jupiter. both to the east.



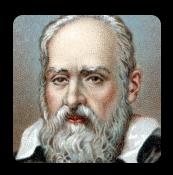
Galileo & Jupiter's Moons



GALILEO'S "NEW ORDER"

Created by Alyssa Goodman, Curtis Wong and Pat Udomprasert, with advice from Owen Gingerich and David Malin

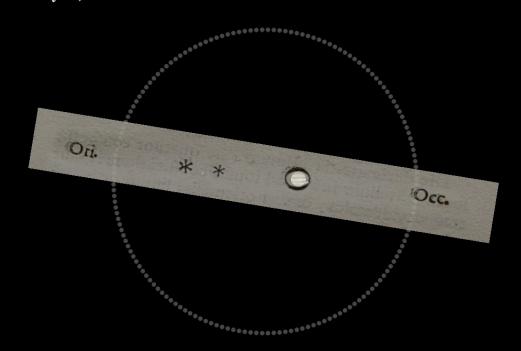




Galileo & Jupiter's Moons

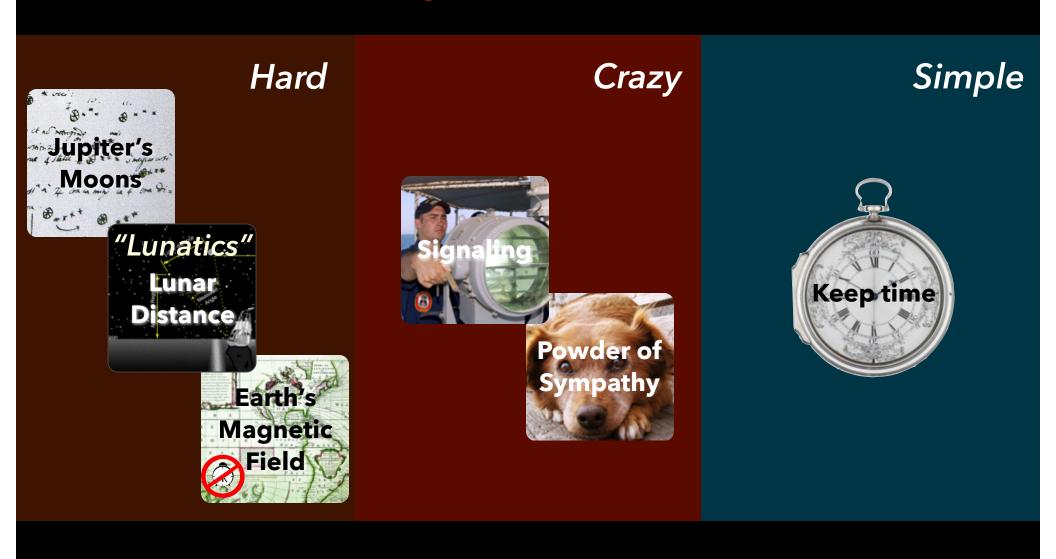


January 11, 1610



Finding Longitude

...use the stars to find longitude if we know the time, very well...



"do it with a clock..."

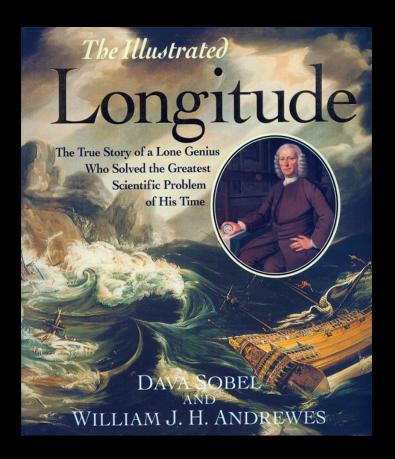




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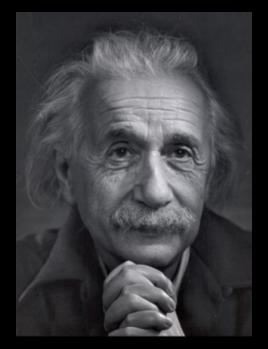










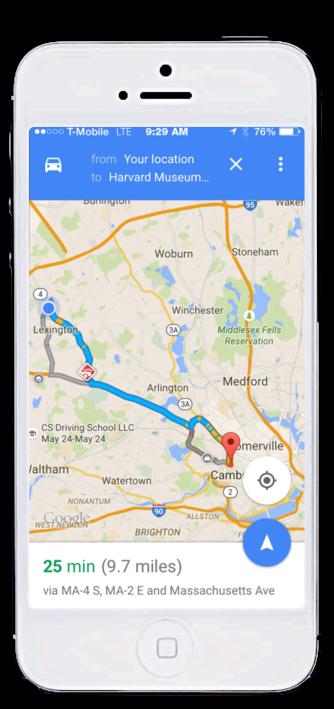


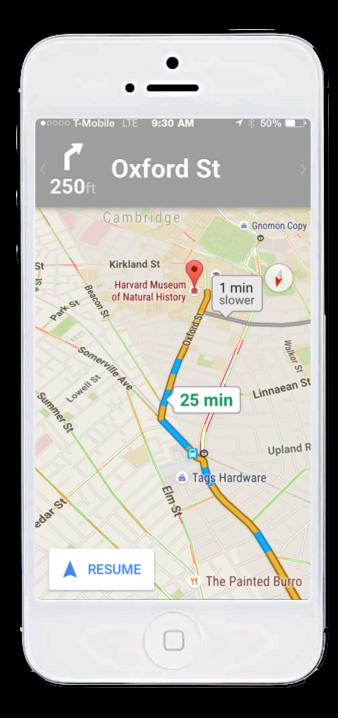


Einstein (, Jobs) & GPS

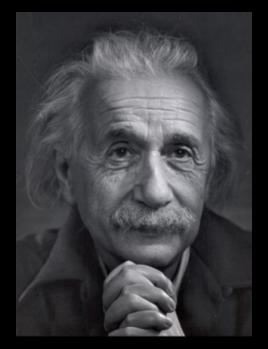












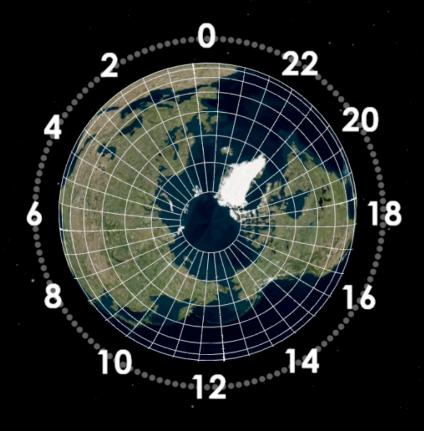


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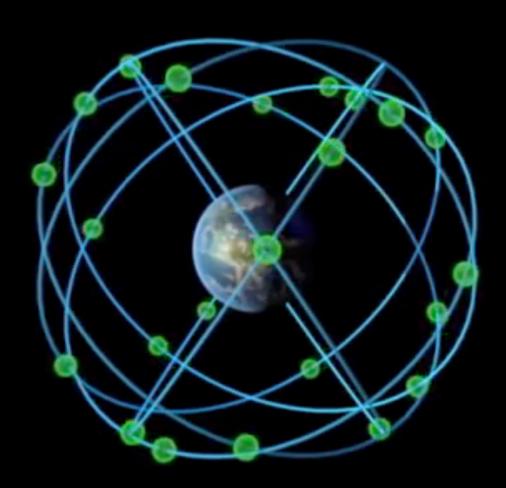


Time-based navigation relies on known speed... of Earth's rotation.



shown at 2 hours per second elapsed on-screen

GPS also relies on known speed...of light



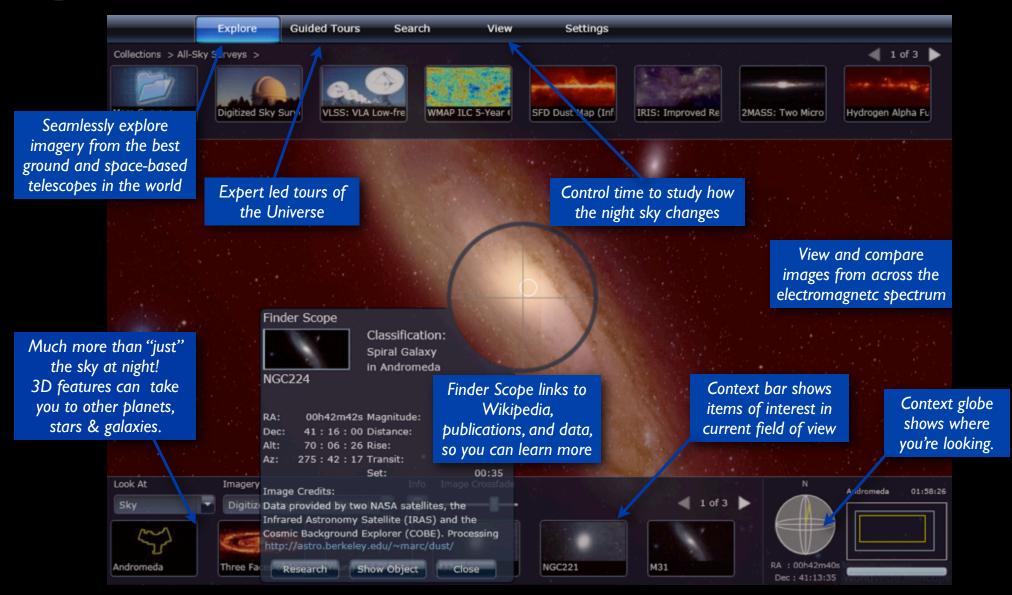




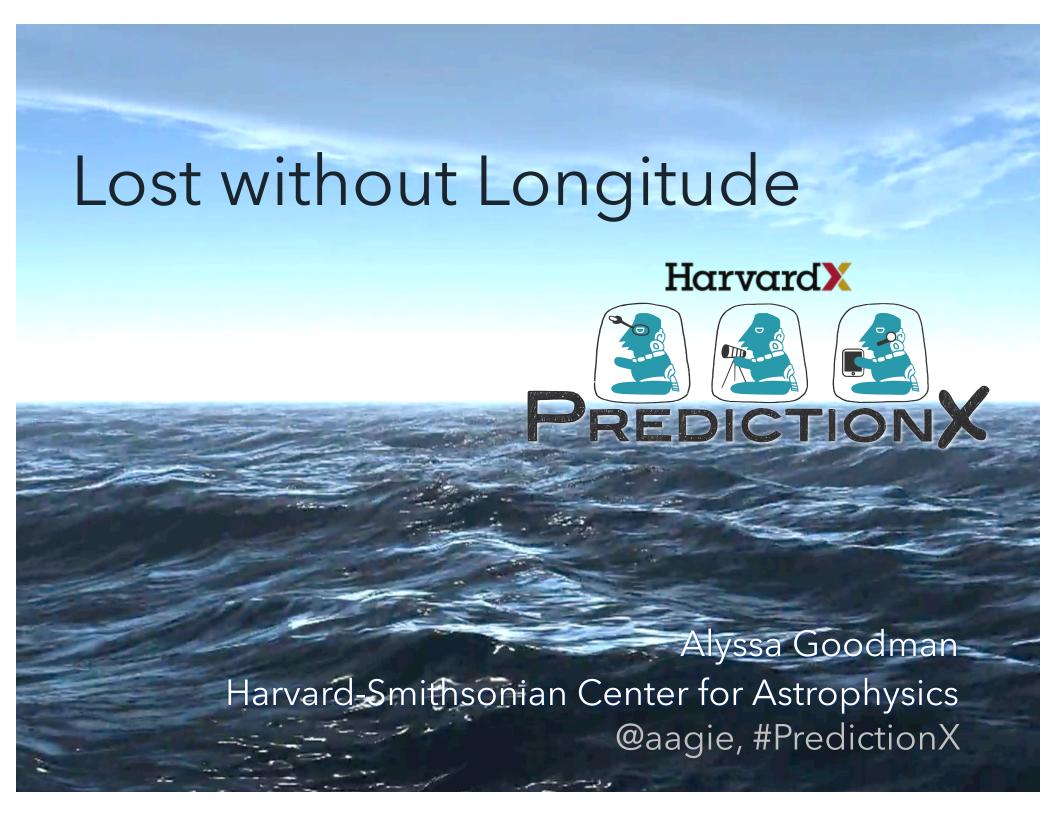


Open WorldWide Telescope

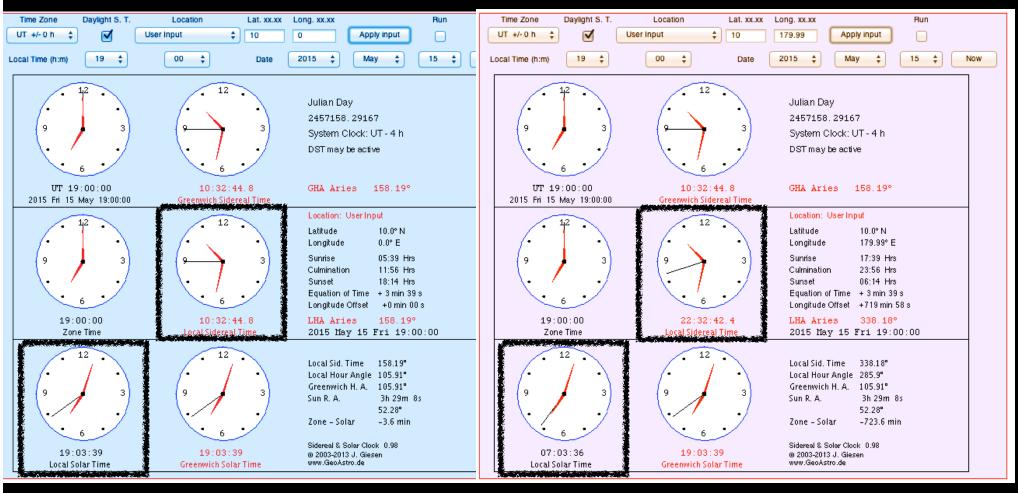
worldwidetelescope.org







solar-sidereal off by 2 min for 180 deg of longitude difference



Greenwich solar-sidereal= 8h 31m 54.2 s

local solar-sidereal=8 h 29 m 53.6s

Sidereal time is the **right ascension** of stars on your local meridian at any moment.



