A photograph of the Submillimeter Array (SMA) radio telescope array at night. The sky is dark blue with a dense field of stars and the Milky Way galaxy visible. Several large, white, parabolic radio telescope dishes are mounted on a dark structure, illuminated from below. The text is overlaid on the image.

# SMA [C II] Study of High Redshift Galaxies

Min S. Yun  
University of Massachusetts

*The Submillimeter Array: First Decade of Discovery*

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# Contributors

**UMASS:** M. Yun, G. Wilson, P. Schloerb, G. Narayanan, N. Erickson, R. Snell, K. Harrington

**CfA:** M. Gurwell, D. Wilner, G. Fazio, P. Ho

**INAOE:** D. Hughes, I. Aretxaga, M. Chavez, A. Montana, M. Zeballos

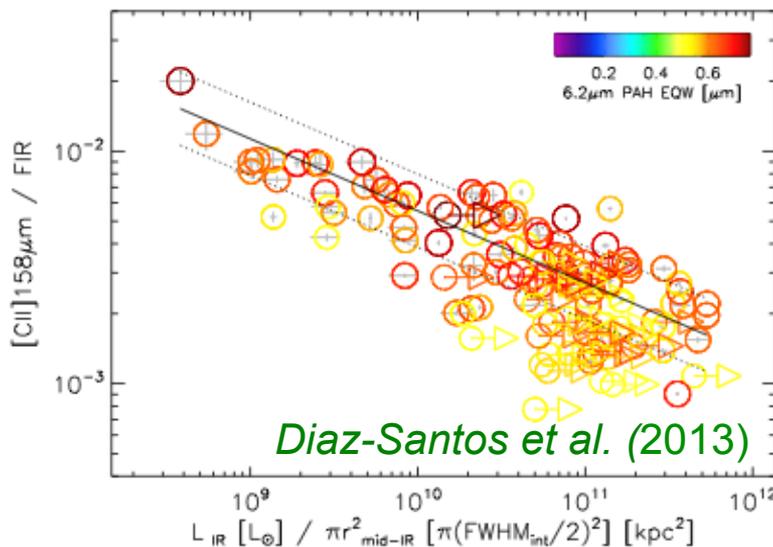
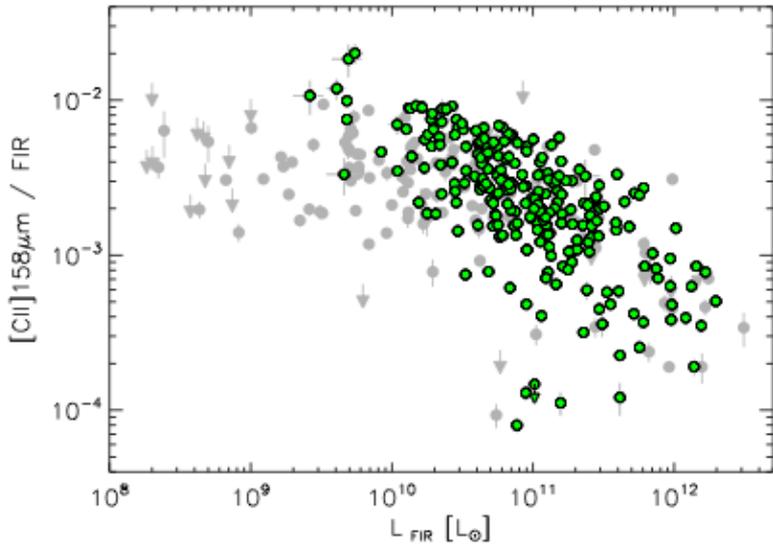
Others: D. Iono (ALMA-J), J. Younger



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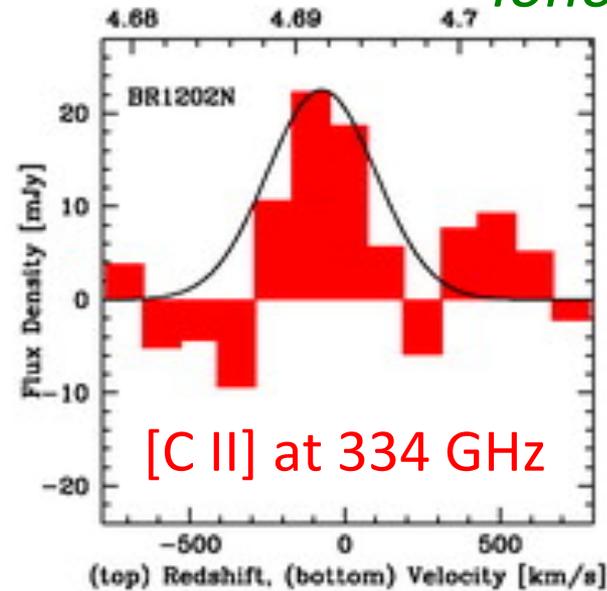
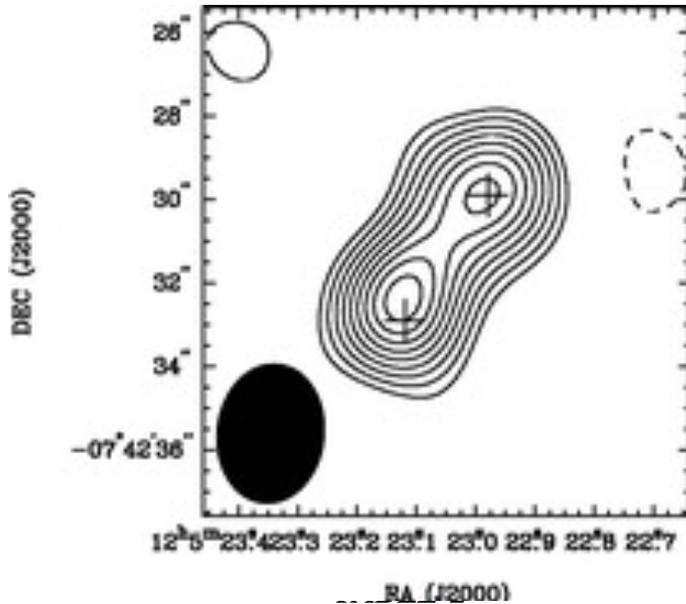
# [C II] as a Tracer of Gas and SF in Galaxies



- [C II] 157 μ line is the most intense far-IR emission line in normal, SF galaxies – as much as 1% of total IR luminosity.
- ISO, ZEUS/CSO, and now Herschel-PACS have surveyed a large number of nearby (and some distant) galaxies.
- Promising tool for studying SF galaxies in high-z galaxies (a motivation for SMA)

# [C II] in $z=4.7$ QSO BR1202-0275

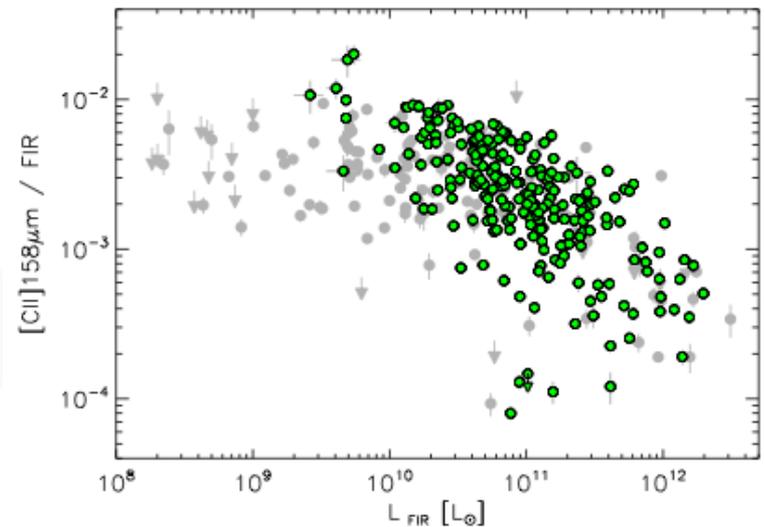
*Iono et al. (2006)*



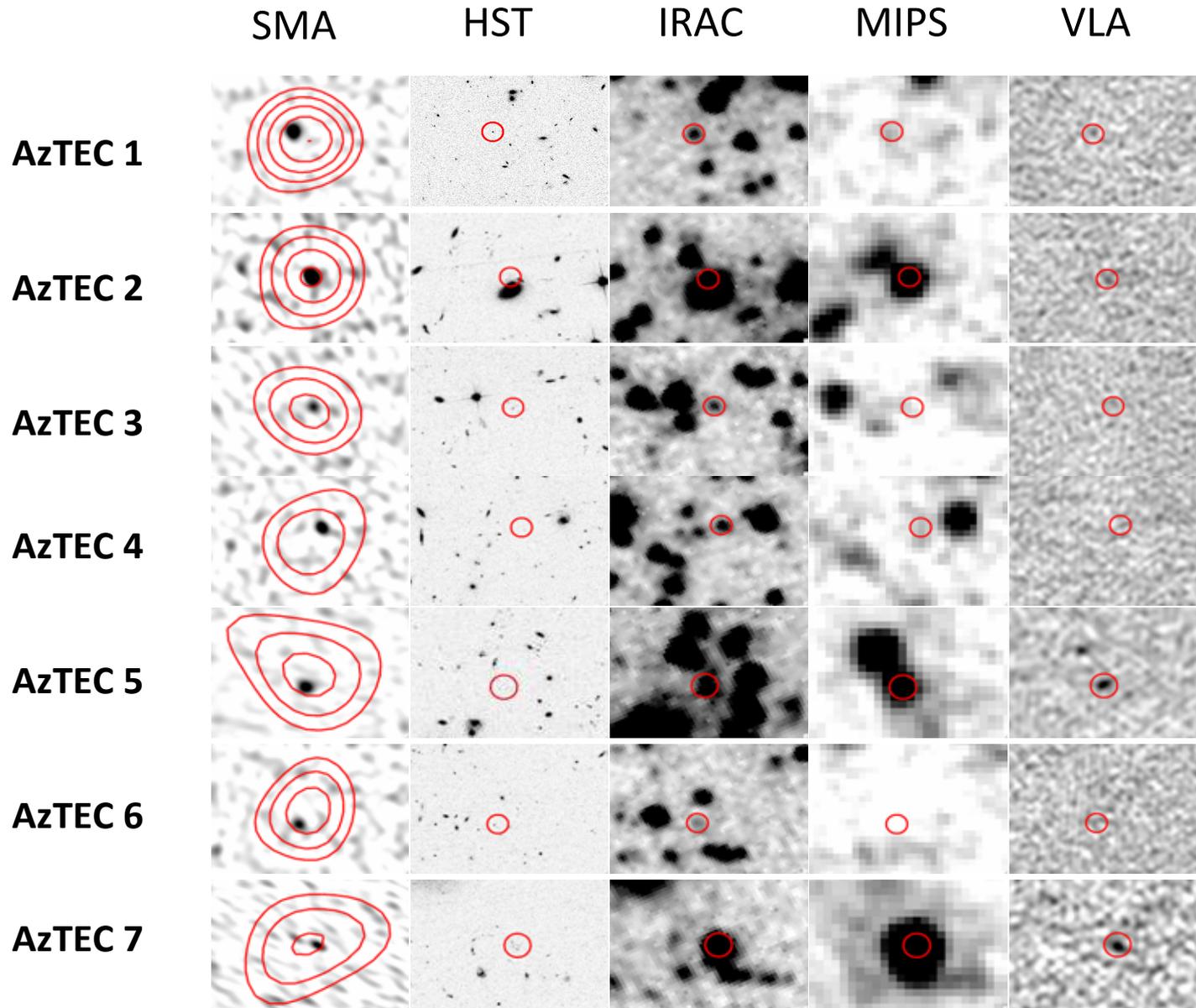
SUMMARY OF LUMINOSITY RATIOS

Source	$L_{\text{FIR}} (L_{\odot})$	$L_{[\text{C II}]} / L_{\text{FIR}} (\%)$	Ref.
Normal galaxies .....	$10^7 - 10^{11}$	0.1–1	1
ULIRGs .....	$\sim 10^{12}$	0.01–0.1	2
High- $z$ sources:			
BR 1202N ( $z = 4.7$ ) .....	$1.2 \times 10^{13}$	0.04	3
BR 1202S ( $z = 4.7$ ) .....	$2.6 \times 10^{13}$	$< 0.03$	3
SDSS J1148+5251 ( $z = 6.4$ ) .....	$1.2 \times 10^{13}$	0.02	4
CI 1358+62 ( $z = 4.9$ ) .....	$2.4 \times 10^{12}$	$< 0.4$	5

REFERENCES.—(1) Malhotra et al. 2001; (2) Luhman et al. 2003; (3) this work; (4) Maiolino et al. 2005; (5) Marsden et al. 2005.

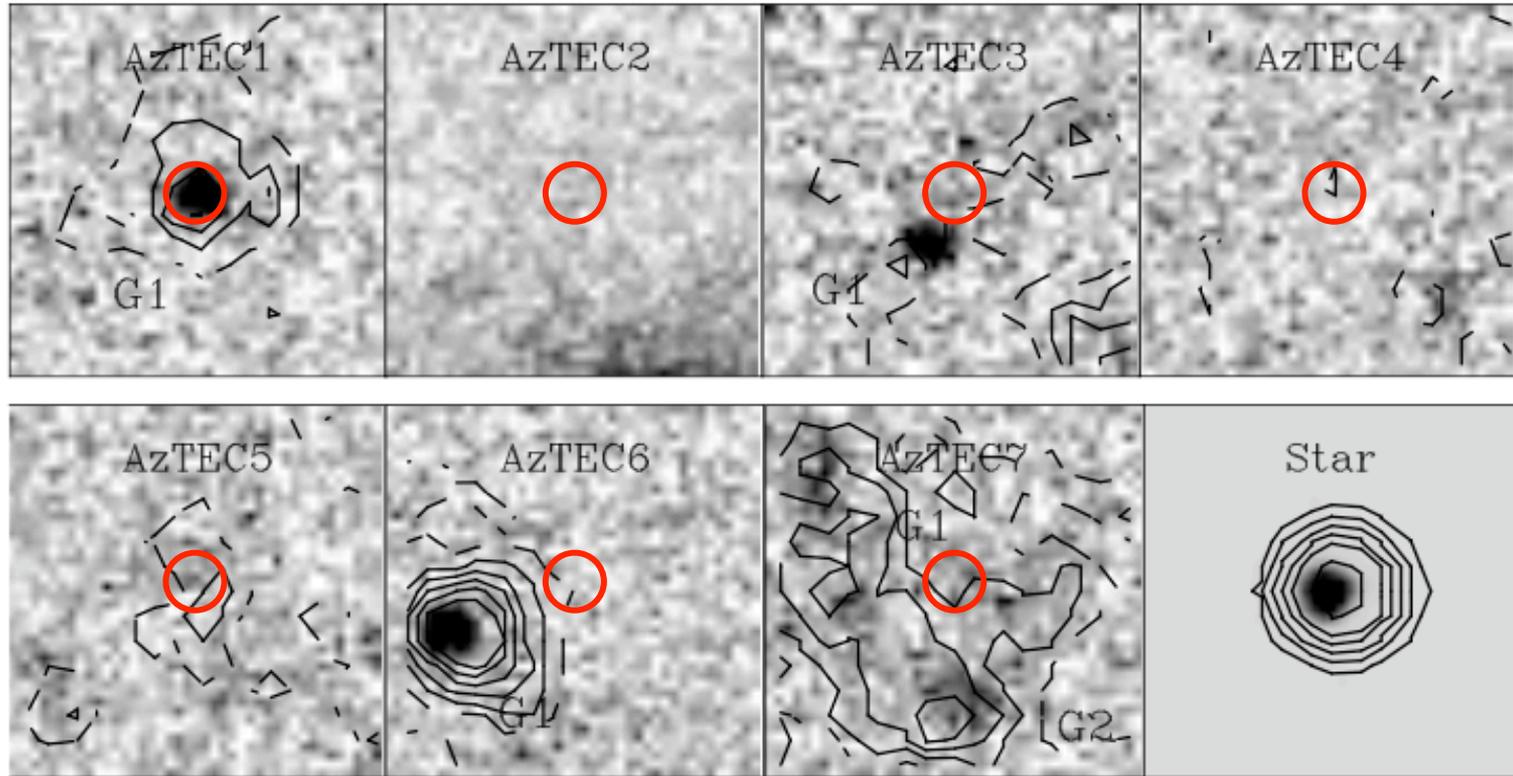


# SMA Follow-up of AzTEC 1.1mm Sources



*Younger et al. (2007,2009)*

# Young Massive SBs at high-z are Dark

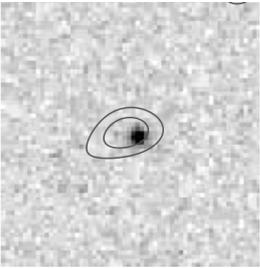


SMA

Greyscale: HST *i*-band, Contours: Subaru *r*-band

Many are exceedingly faint with  $i > 25$  and  $r > 27$

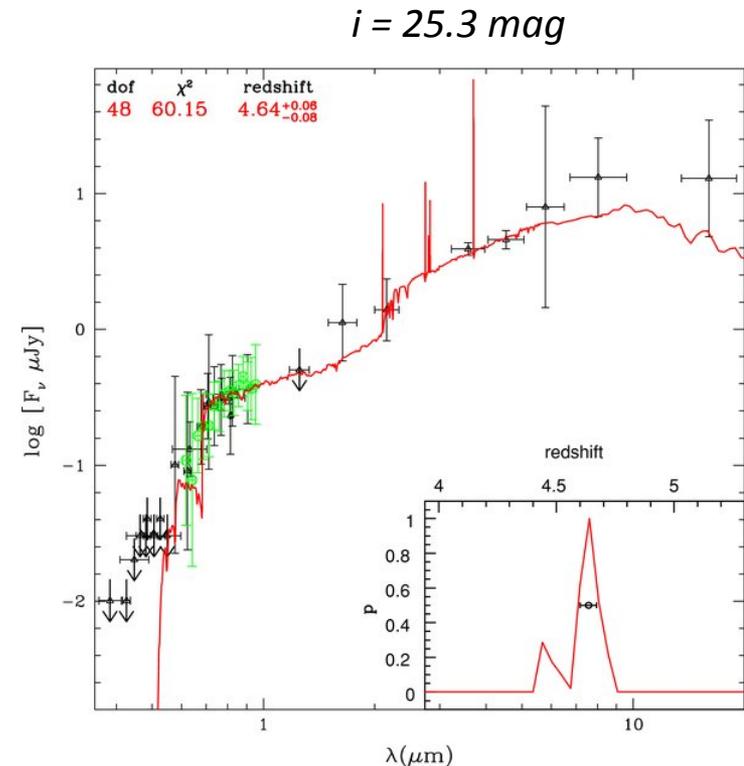
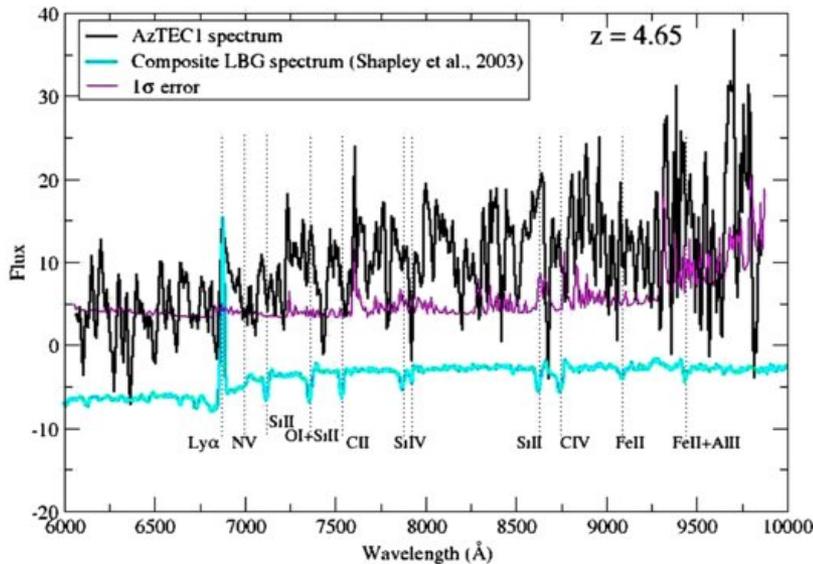
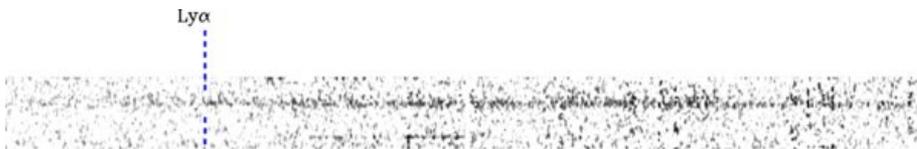
Younger et al. (2007)



# COSMOS AzTEC-1 at $z=4.6$ ?

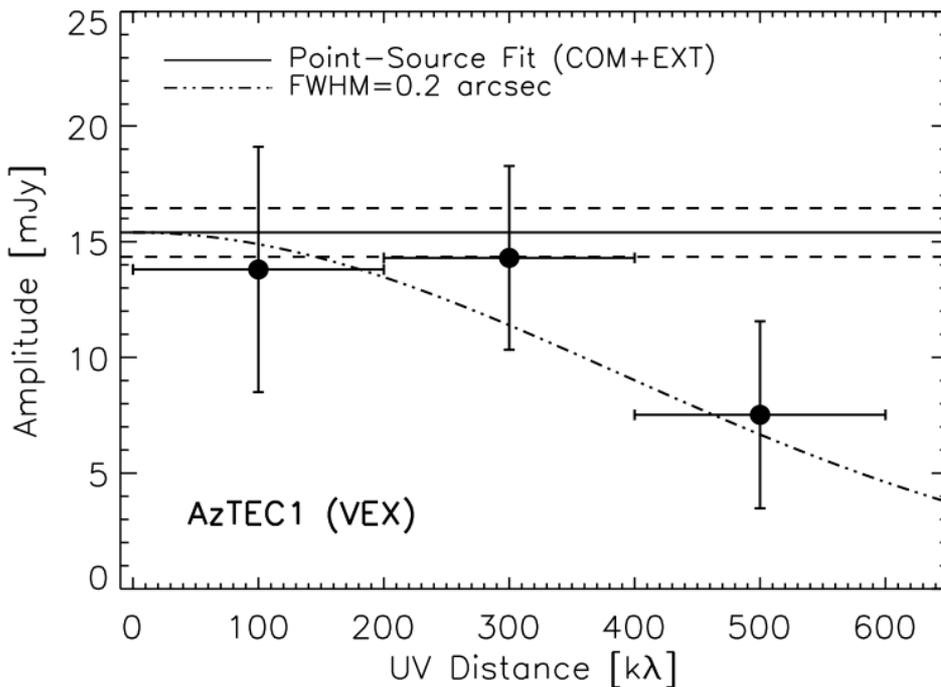
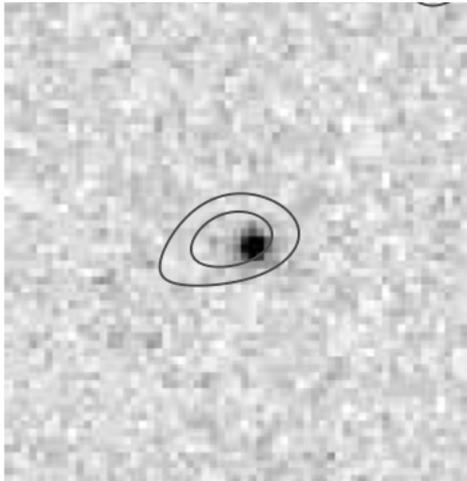
*Smolcic et al. (2011)*

- Keck DEIMOS Observations
  - No bright emission line
  - No clear Lyman break
- 20+ bands photometry
  - $z_{\text{PHOT}} = 4.64 \pm 0.08$



# SMA 890 $\mu$ Imaging of AzTEC-1

*Younger et al. (2008)*



Resolved with the VEX  
configuration

-- Source size  
 $\sim 0.2''$  (1.4 kpc)

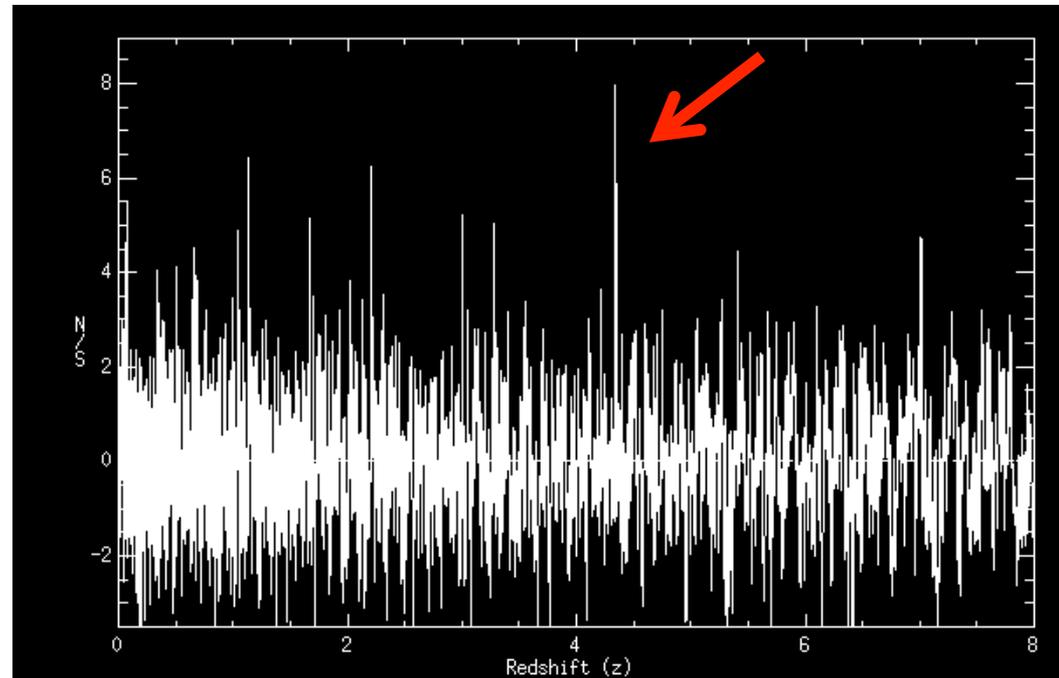
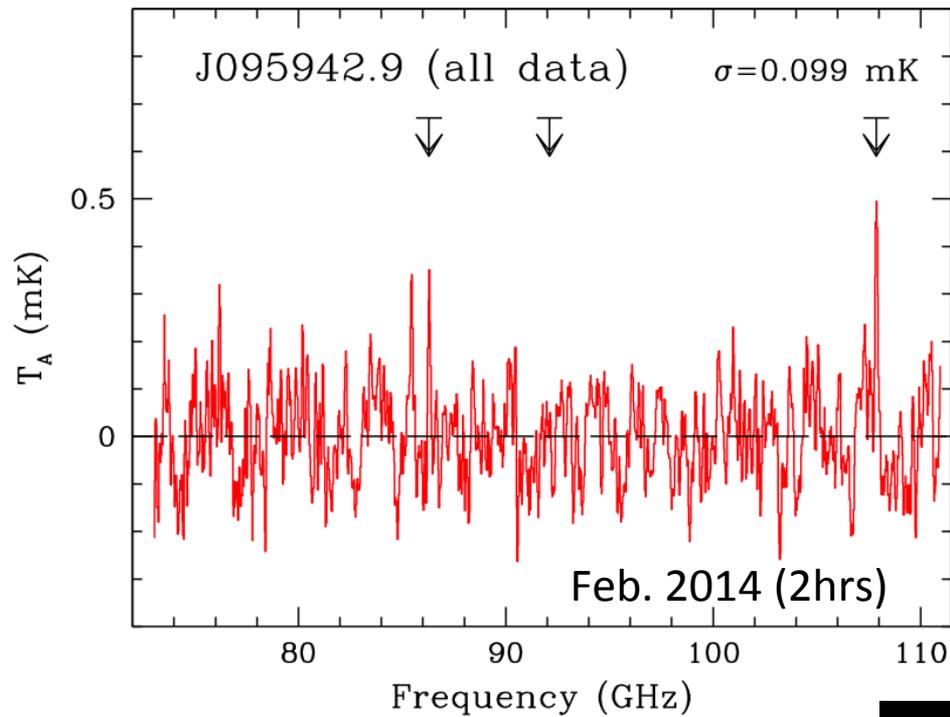
$$\rightarrow \Sigma_{\text{SFR}} = 880 M_{\odot}/\text{kpc}^2$$

$$(\Sigma_{\text{IR}} = 1.1 \times 10^{13} L_{\odot}/\text{kpc}^2)$$

# LMT-RSR Observations of AzTEC-1

*Yun et al. (in prep)*

$$Z(\text{CO}) = 4.3421 \pm 0.0006$$



# AzTEC-1: SMA [C II] Observations

*Yun et al. (in prep)*

Observed on March 1, 2014  
[C II] at 355.8 GHz and 345  
GHz continuum

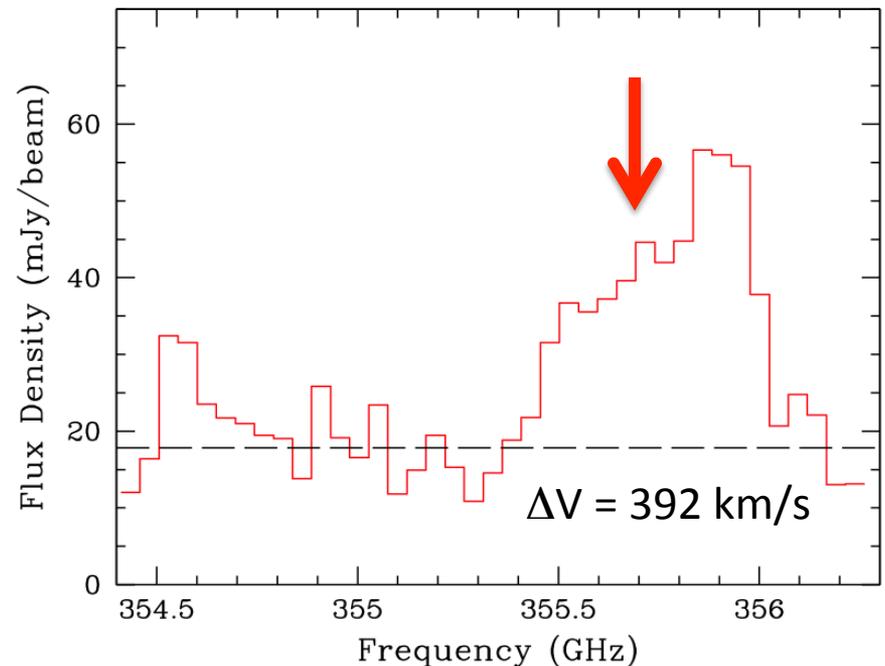
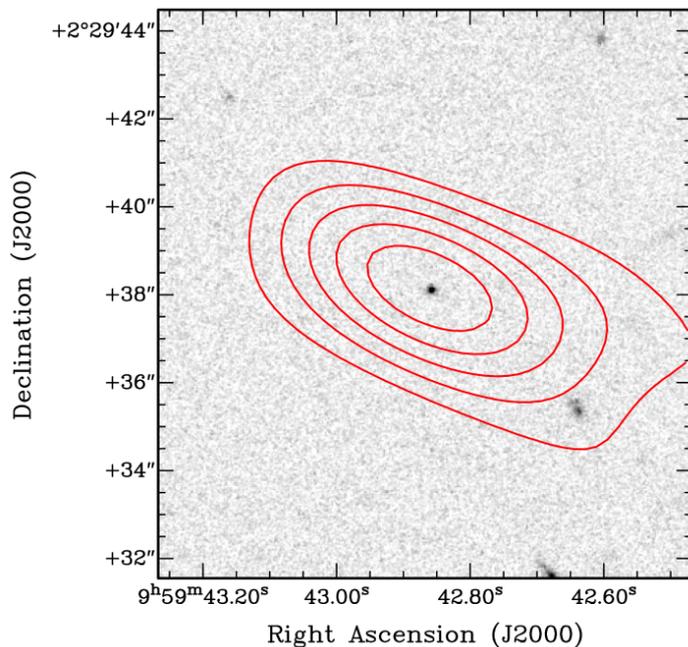
→  $z(\text{C II}) = 4.3415 \pm 0.0003$

$S(\text{CII})\Delta V = 13.1 \pm 0.9 \text{ Jy km/s}$

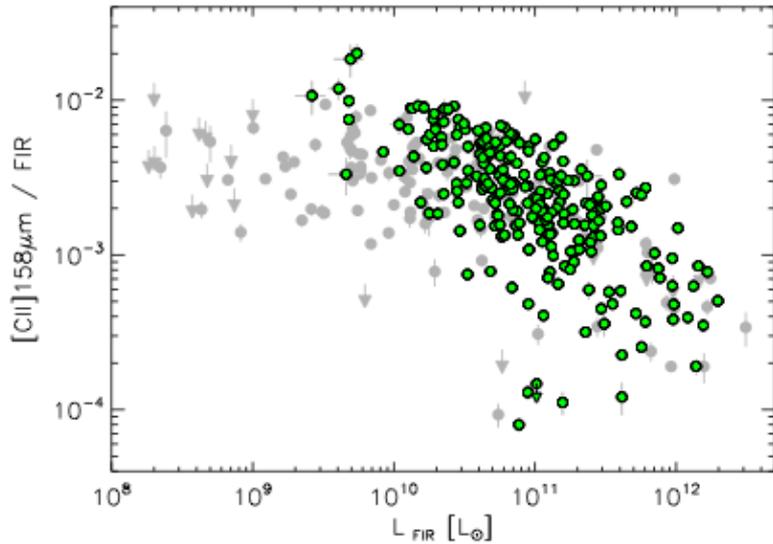
$L(\text{CII}) = 7.8 \times 10^9 L_{\odot}$

$L(\text{IR}) = 1.7 \times 10^{13} L_{\odot}$

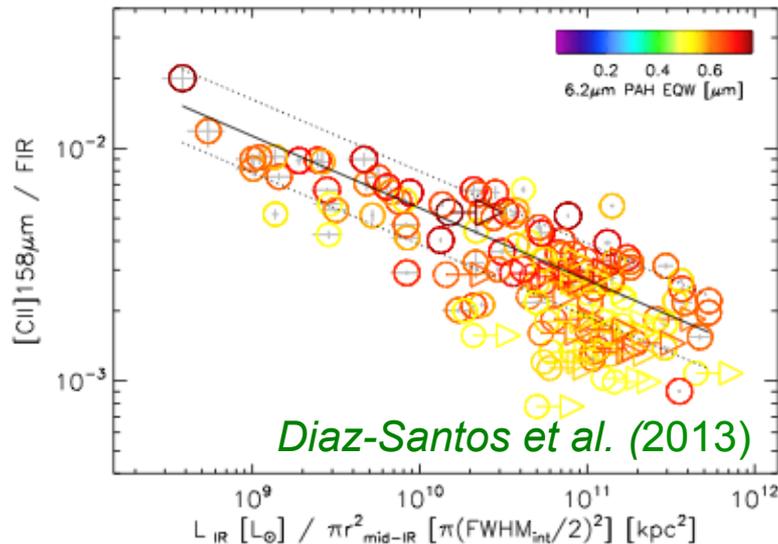
→  $L(\text{CII})/L(\text{IR}) = 0.00046$



# [C II] in AzTEC-1

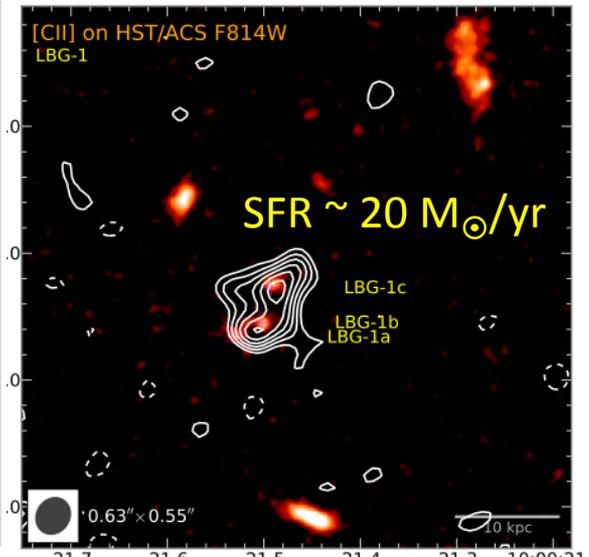
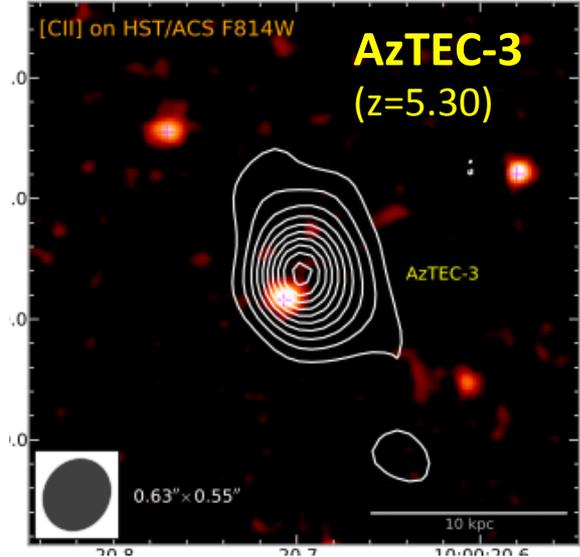
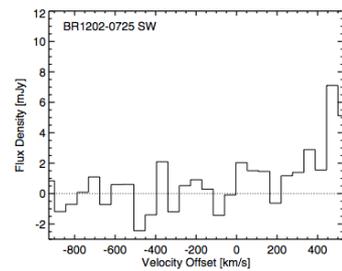
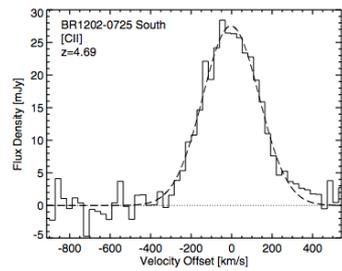
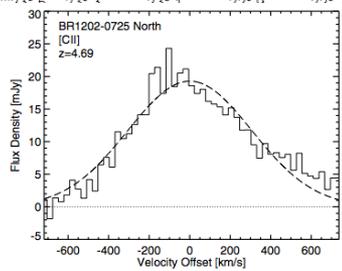
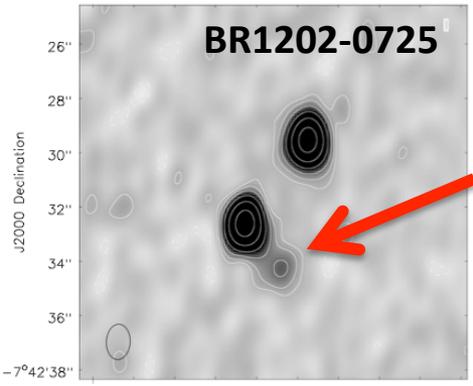


- SMA observations of [C II] line in AzTEC-1 (and BR1202) extend the observed local  $L(CII)/L(FIR)$  trend to  $z=4$  galaxies
- These observations also extend the  $L(CII)/L(FIR)$  relation to higher IR surface brightness density by another order of mag



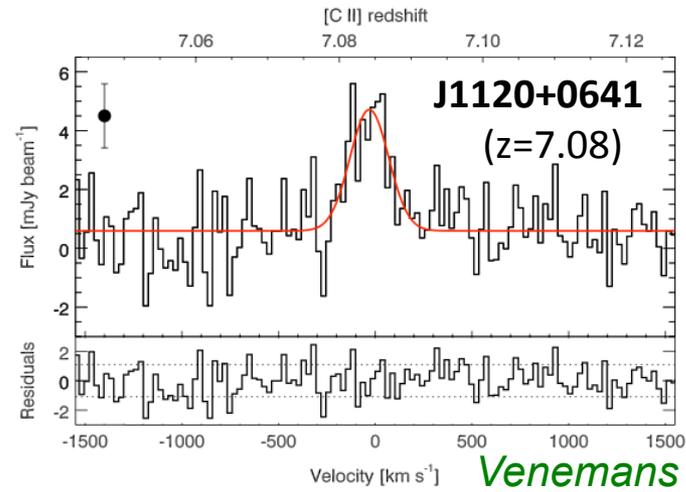
→ the deficit in [C II] due to an increase in ionization parameter  $\langle U \rangle$ ?

# More Intriguing [C II] Observations

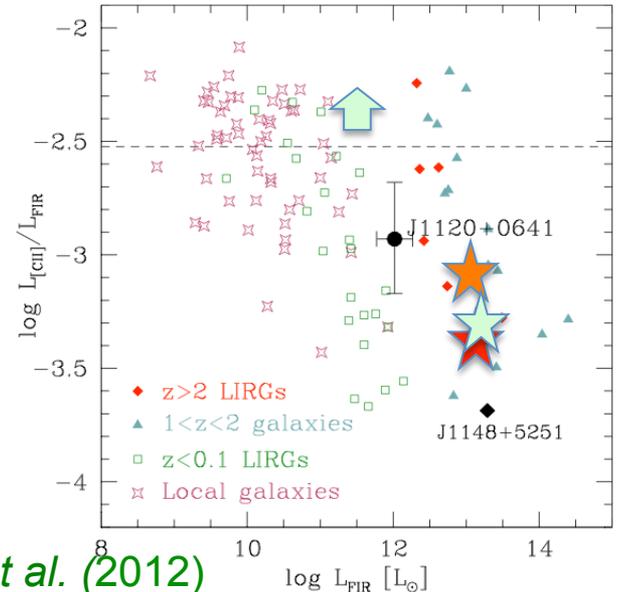


*Riechers et al. (2014)*

*Wagg et al. (2012)*



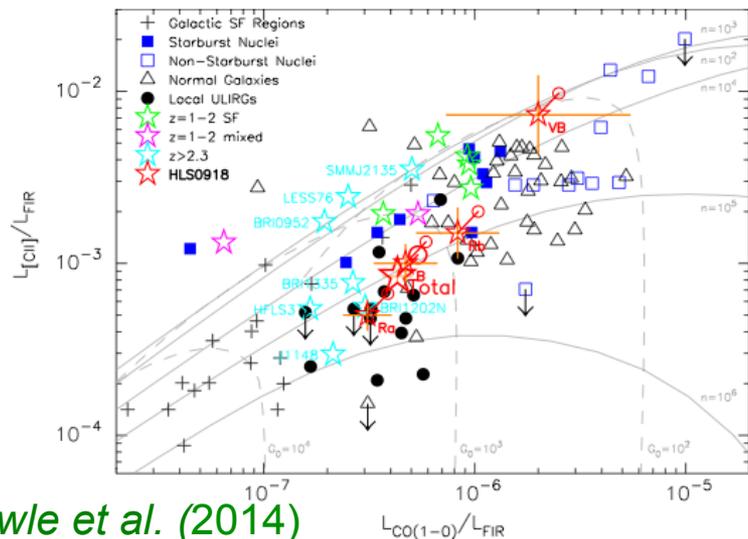
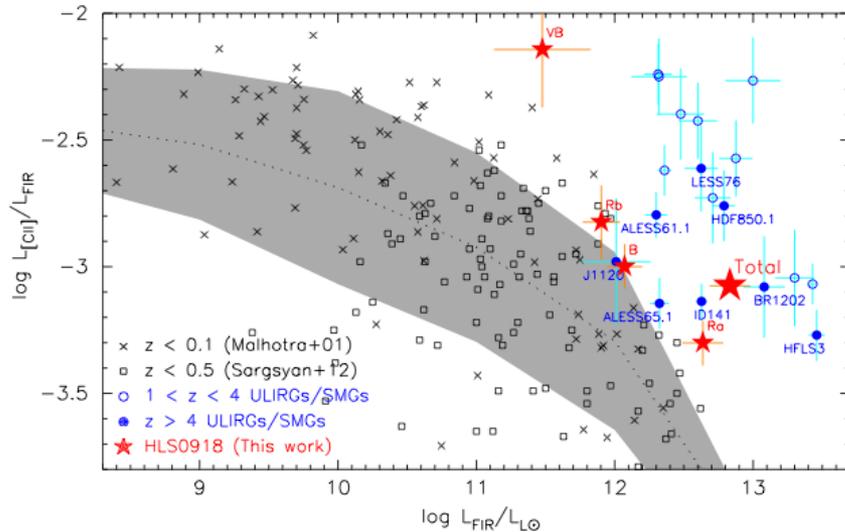
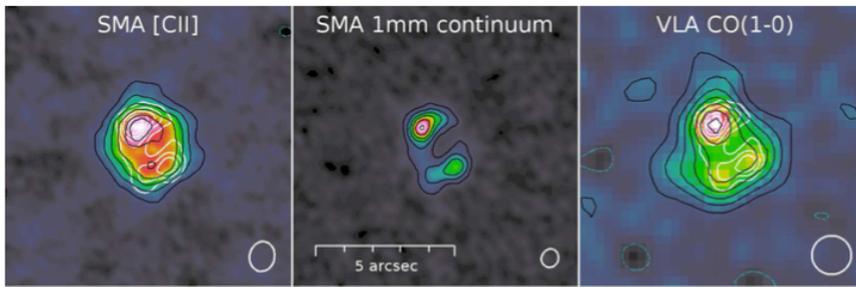
*Venemans et al. (2012)*



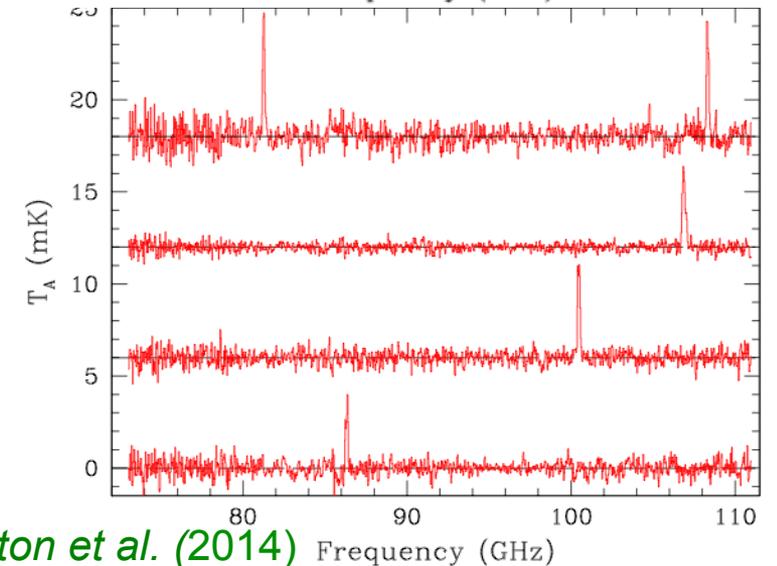
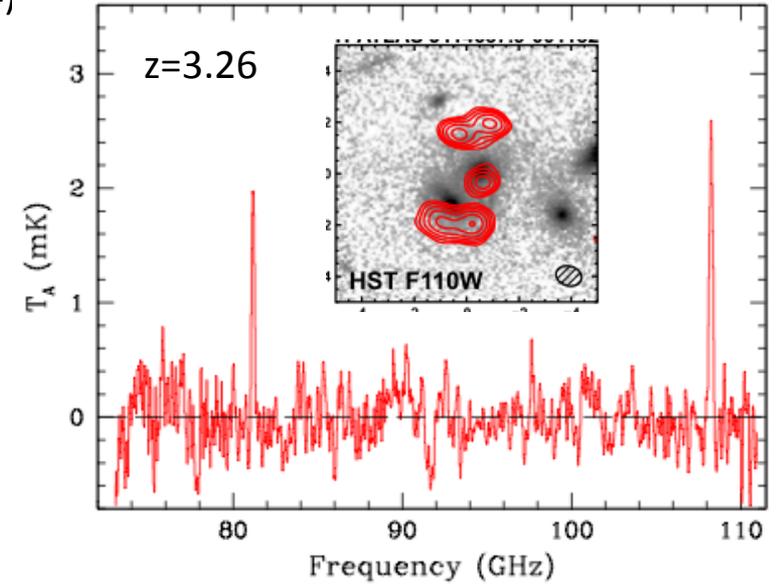
# Strongly Lensed

# Sources *Bussmann+ (2013)*

HLSJ091828  
( $z=5.24$ )



*Rawle et al. (2014)*



*Harrington et al. (2014)*

# Summary

- [C II] 157 $\mu$  line provides a highly sensitive tracer of gas and SF distribution in early Universe.
- SMA Observations of BR1202, AzTEC-1, and HLSJ091829 extended the local [C II]-FIR relations to  $z > 4$  and supports the “maximum SB” scenario (and only a minor contribution from a buried QSO) with high  $\langle U \rangle$ .
- SMA [C II] studies of highly lensed sources can potentially yield a valuable new insight on the “Epoch of Galaxy Growth”, including “normal” galaxies.

# HCG 57: An Interesting Local Example

*Alatalo et al. (2014)*

