



Observations of High- z Galaxies

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SMA Follow-Up of AzTEC/COSMOS Sources

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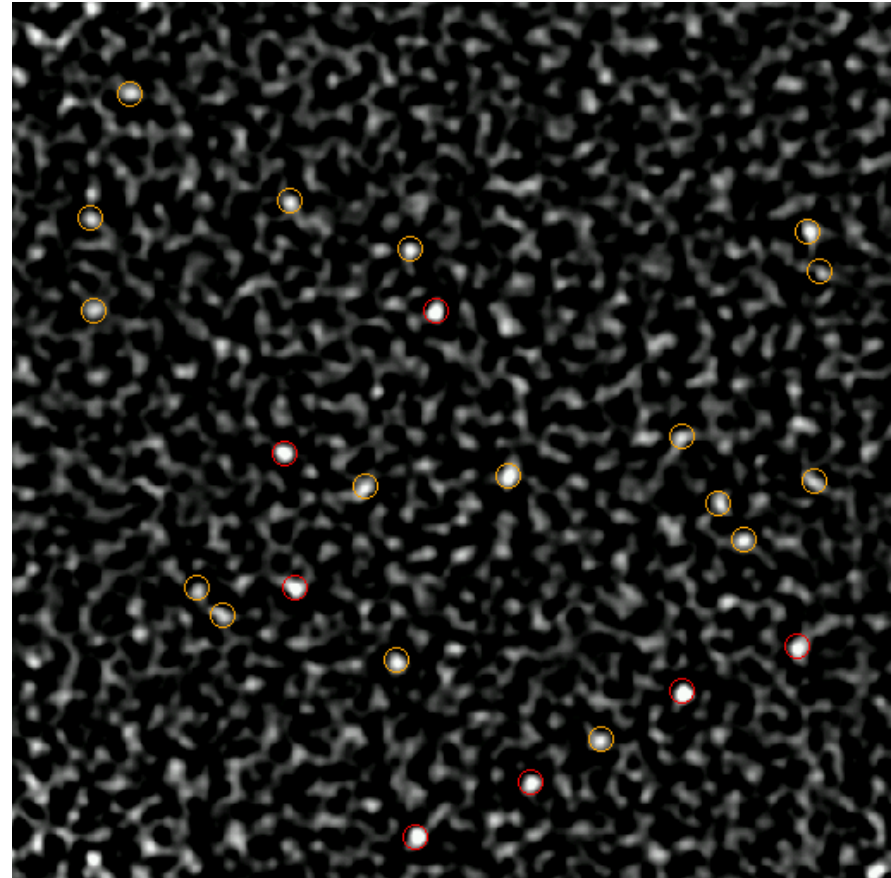
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OBSERVATIONS

- AzTEC camera observations (1.1 mm wavelength; 18 arcsec resolution) on the JCMT of the COSMOS field (0.15 deg²) detected 44 submm galaxies (SMGs) above 3.5 σ .
- SMA interferometric observations (890 μ m wavelength; 2 arcsec resolution) of the seven brightest AzTEC sources detected all seven SMGs and pinpointed their location to 0.2 arcsec.
- Follow-up observations by HST (ACS), SPITZER (IRAC and MIPS), and Very Large Array (VLA) revealed the detailed properties of these sources.

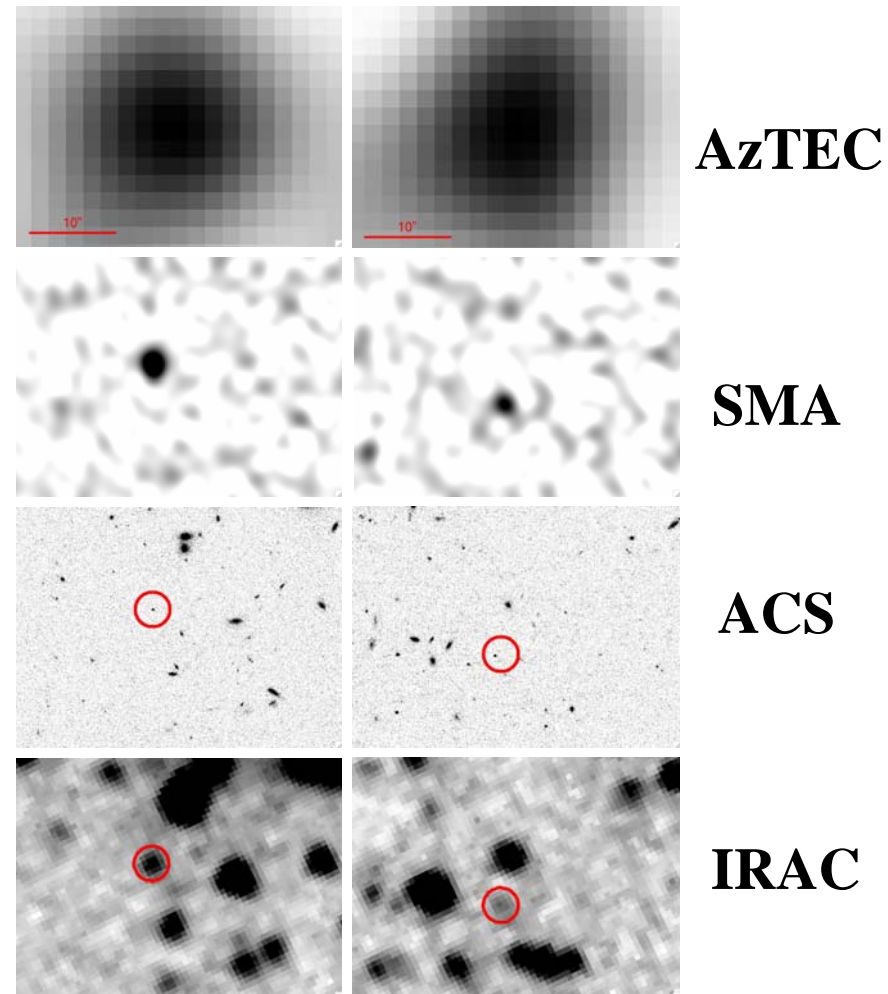
The AzTEC/COSMOS Survey
(Scott et al. 2007)





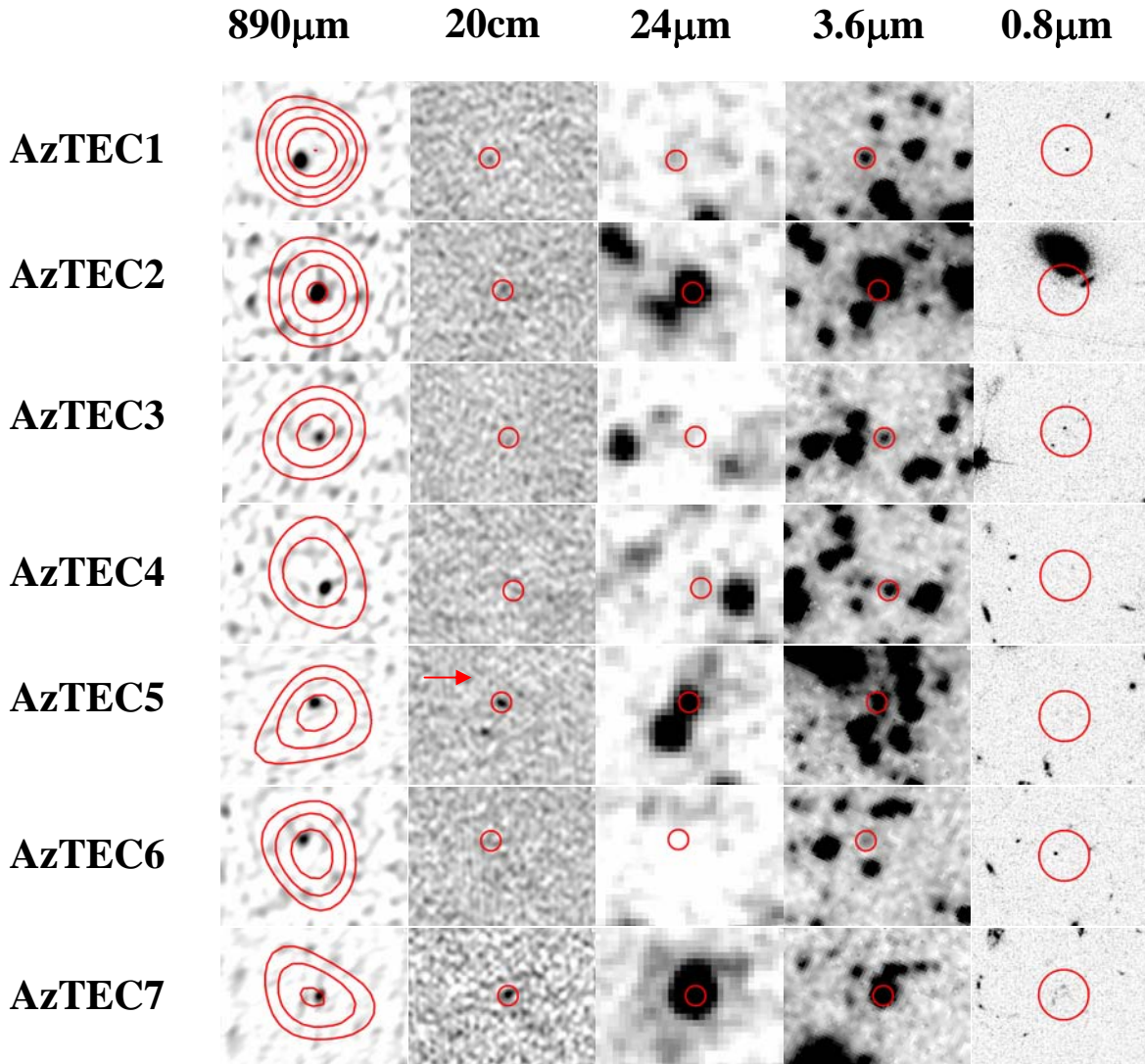
Extreme Starbursts in the Early Universe: SMA/AzTEC Observations of Submillimeter Galaxies

- Of the seven SMGs detected, five are radio-quiet.
- For the five radio-quiet sources ($< 60 \mu\text{Jy}$), the submm, infrared and optical properties of these counterparts suggest high redshift.
 - High submm/radio fluxes
 - Submm sizes < 1.2 arcsec
 - Very low IRAC fluxes
 - Lack of optical flux
 - No MIPS detection at $24 \mu\text{m}$
- These results are evidence for a population of extreme starburst galaxies when the universe was less than 16% of its current age ($z > 3$).



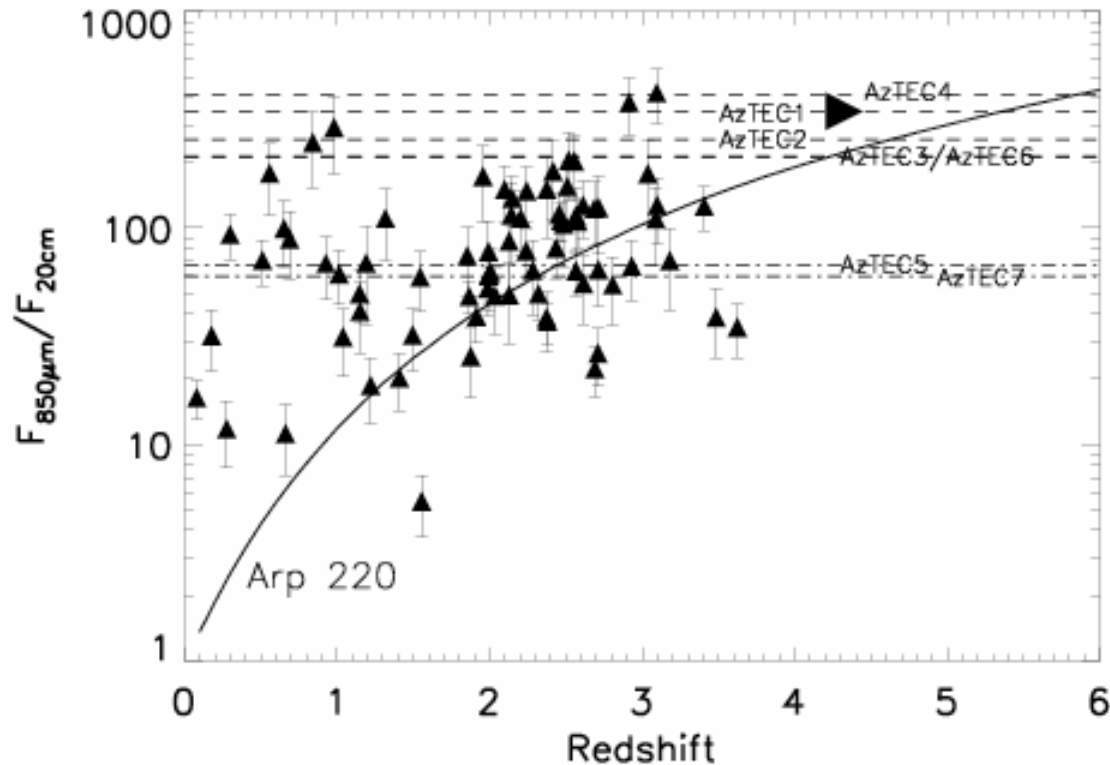


Multiwavelength Images





Radio/Submm Flux Ratios



Consistent with higher average/median redshift



IRAC Counterparts

QuickTime™ and a
TIFF (Uncompressed) decompressor
are needed to see this picture.

**Consistent with higher
average/median redshift**



CONCLUSIONS

- Detected a complete, flux limited sample of SMGs at high angular resolution
- Radio-quiet sources appear to be at higher redshift ($z > 3$) than radio-loud sources
 - Chapman redshift distribution is a lower-limit on z
- Brightest SMGs are compact; size scales analogous to local ULIRGs; $L > 5 - 10 \times 10^{12} L(\text{sun})$.
- Population of higher redshift SMGs that contribute significantly to observed number counts
 - Constraints on galaxy formation models; have to have massive starbursts at early times
 - Constraints on dust formation models; how to make enough dust given short age of the system
- The brightest SMGs may be at the highest z .



SMA 890 μm Images of High-z Galaxies

QuickTime™ and a
TIFF (Uncompressed) decompressor
are needed to see this picture.

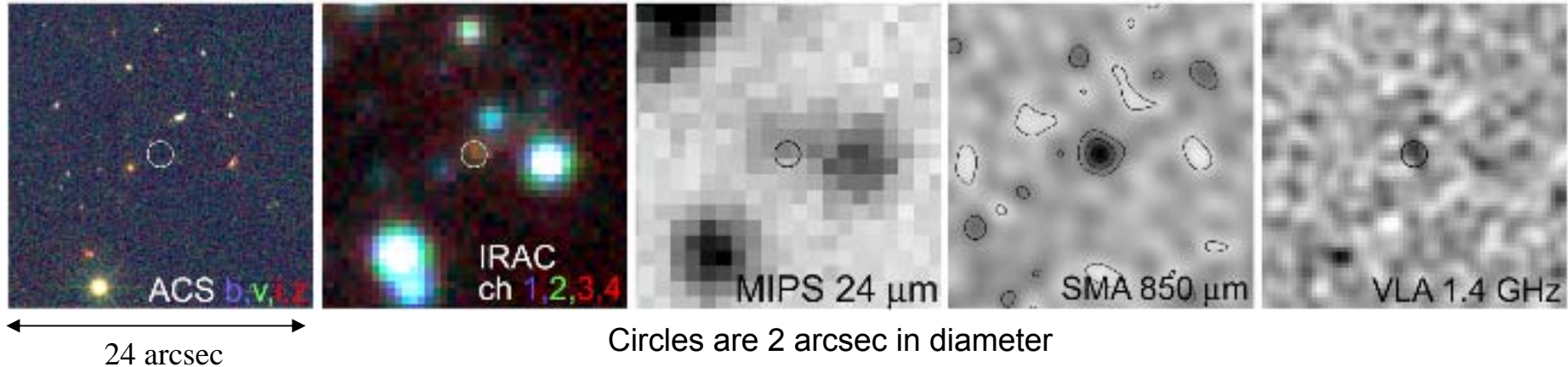
- GN20

- GOODS-N; S(850 μm) \sim 20 mJy
- High z: very weak radio; B-band dropout
- Position: 0.1 to 0.2 arcsec; size $<$ 1.2 arcsec
- Optical galaxy offset by 0.8 arcsec

D. Iono et al. 2006 ApJ, 640, L1



SMA Observations of GOODS 80-5 ($z > 4$)

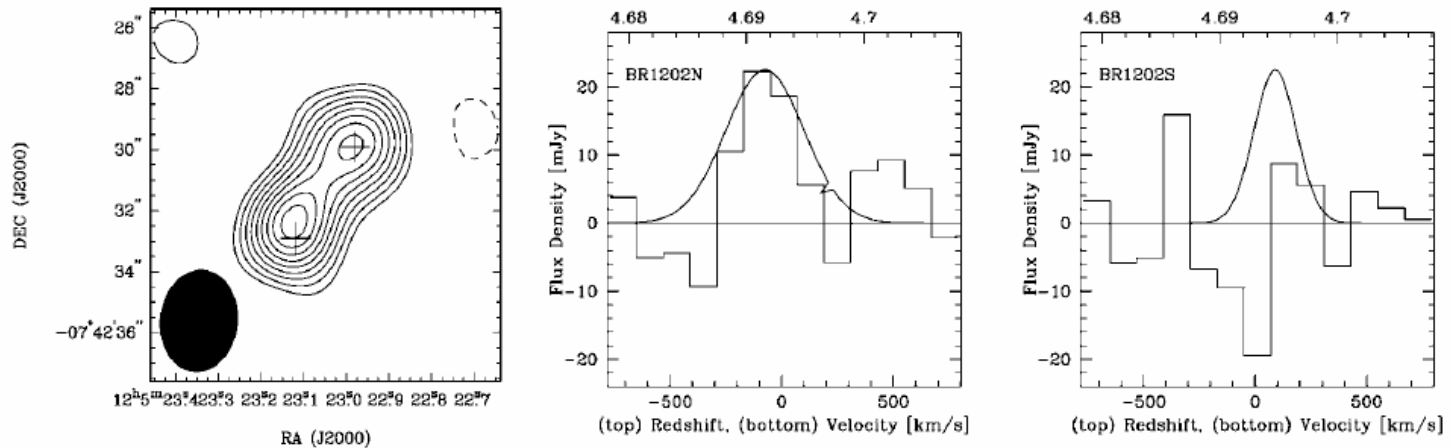


- Submm: 12 mJy source at 850 μ m
- Radio: $< 20 \mu$ Jy at 1.4 GHz, indicating high z
- Spitzer: 1 - 5 μ Jy at 3.6 - 8 μ m; 46 mJy at 24 μ m
- May be the first $z > 4$ source identified in the submm

W-H. Wang et al., submitted to ApJL



[CII] Line Emission in a $z = 4.7$ Quasar



- SMA imaging (~ 3 arcsec) of quasar QSO BR 1202-0725 at $z = 4.7$ yields two components
- [CII] detection with SMA from northern source [$4.5 \times 10^9 L_{(\text{sun})}$]
- [CII]/FIR luminosity = 0.04%
- X-ray emission: Pair of interacting galaxies (both AGN)

D. Iono et al., 2006 ApJ, 645, L97



A Biased mm-Wavelength AzTEC Survey Towards 4C41.17 at $z \sim 3.8$

Tracing accelerated galaxy formation in
a high- z (proto-)cluster

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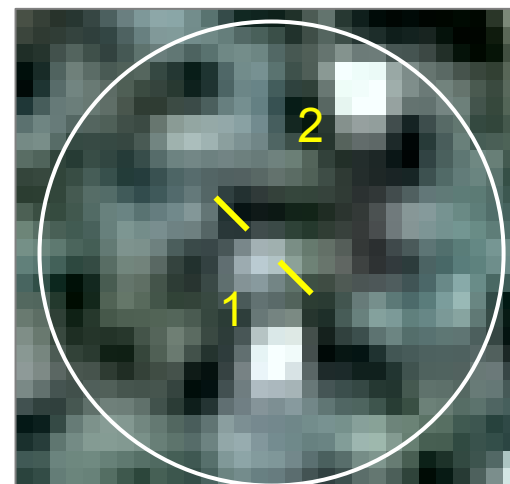
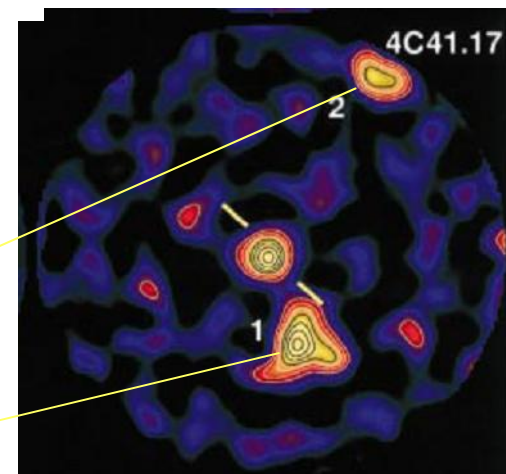
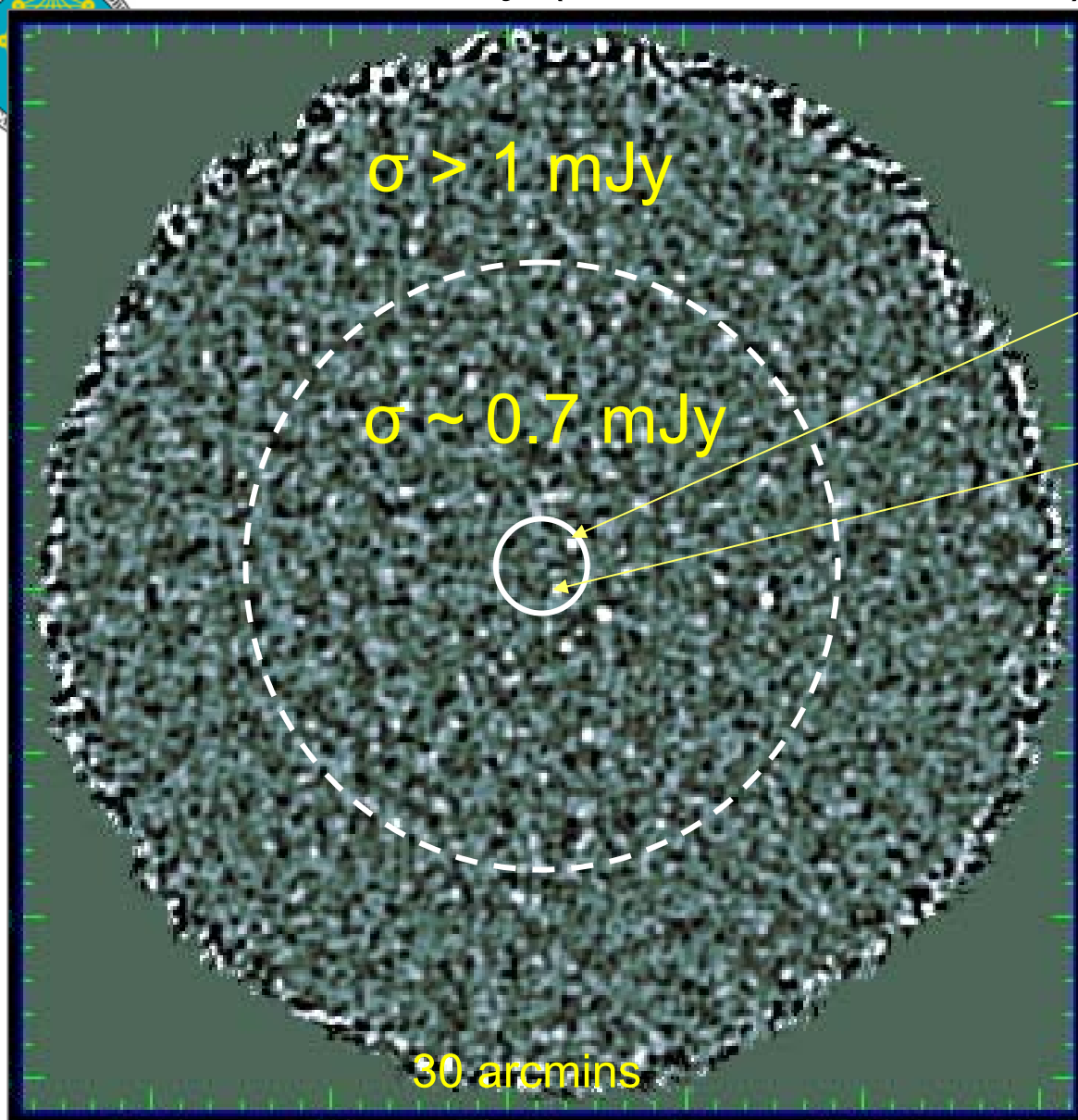


Survey Goals

- Is the AzTEC source over-density associated with the environment of 4C41.37?
- Determine the multiplicity of AzTEC sources.
- Is star formation or AGN activity powering the FIR luminosity?

AzTEC 1.1mm survey (50 hours on JCMT)

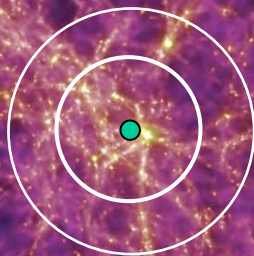
SCUBA/JCMT 850 μ m
Stevens et al. 2003



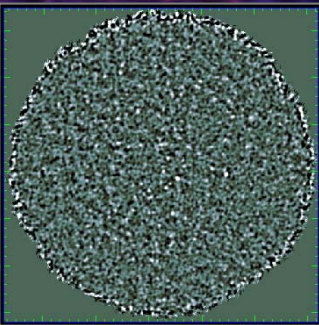
AzTEC 1.1mm

$z \sim 5.7$

31.25 Mpc/h



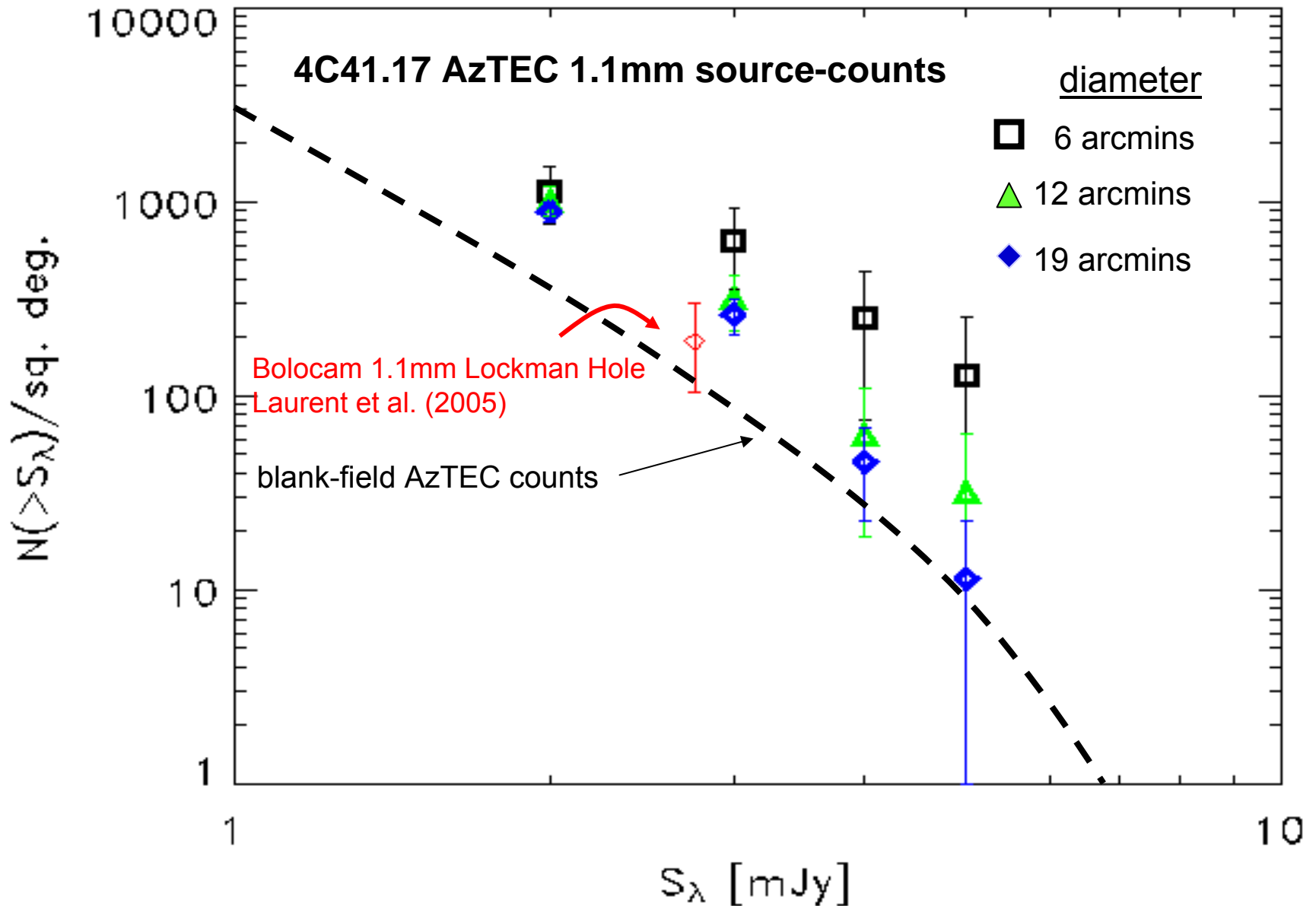
AzTEC 1.1mm



30 arcmin

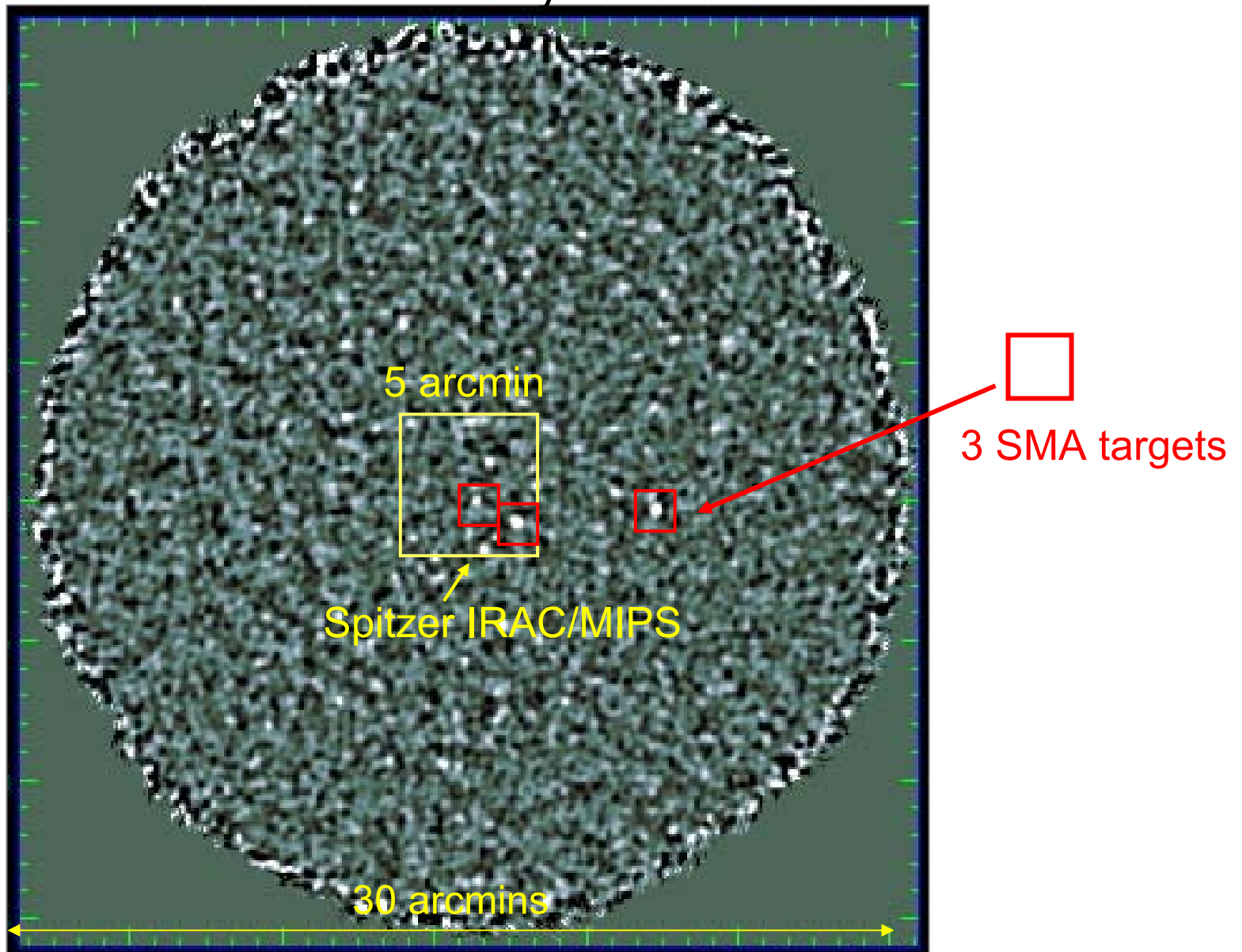
Millennium simulation - Springel et al 2005.

Radially-increasing (factor x 5-10) over-density in $N(>S)$ towards 4C41.17 compared to luminous SMGs in unbiased blank-fields



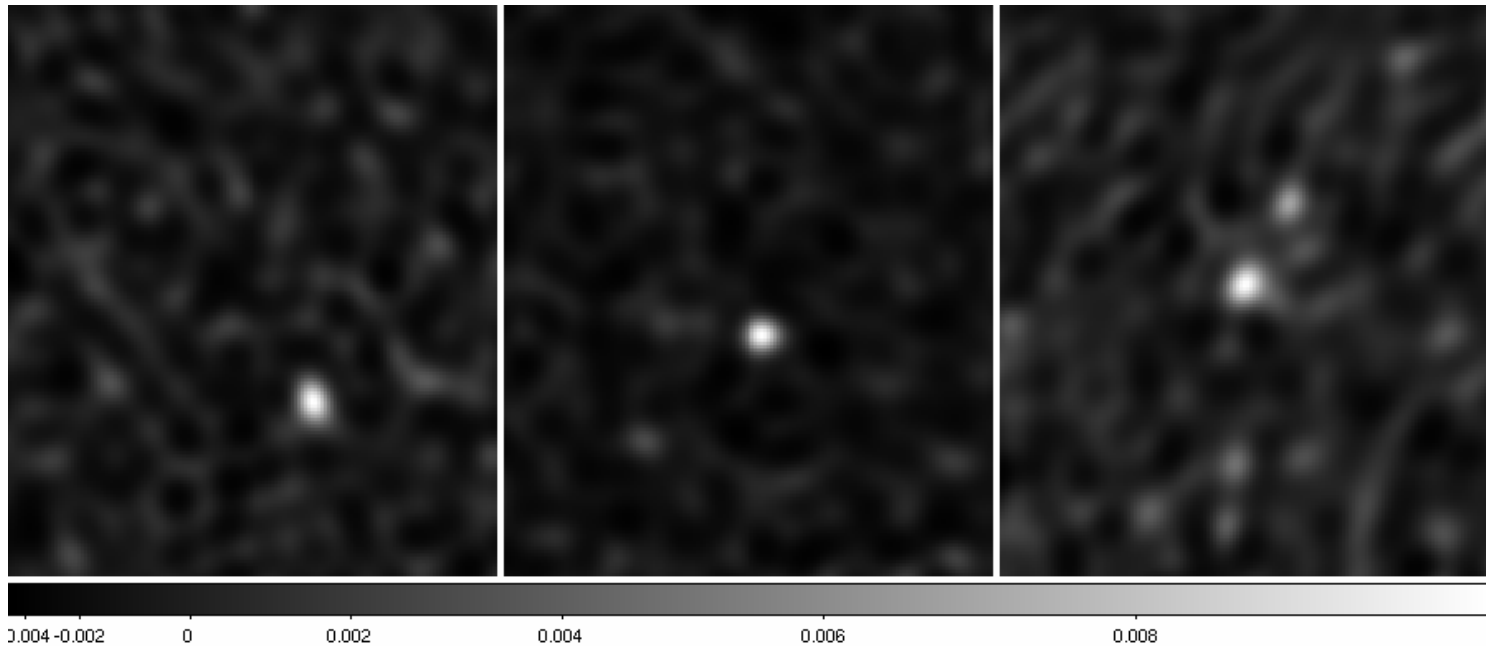


AzTEC 1.1mm Survey





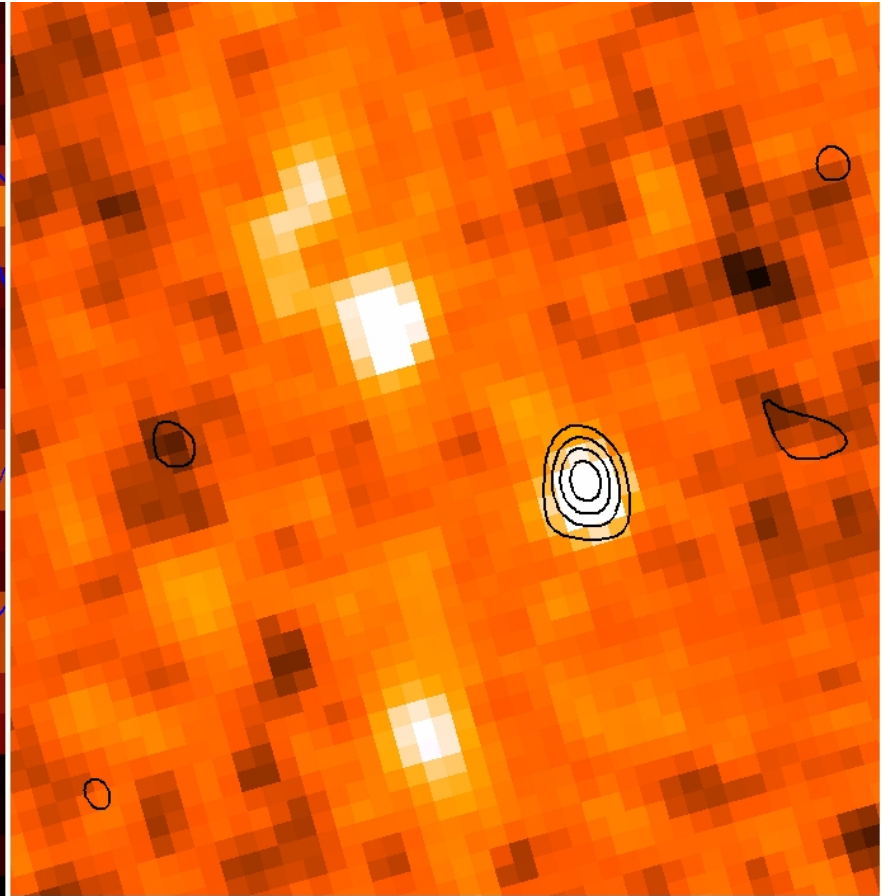
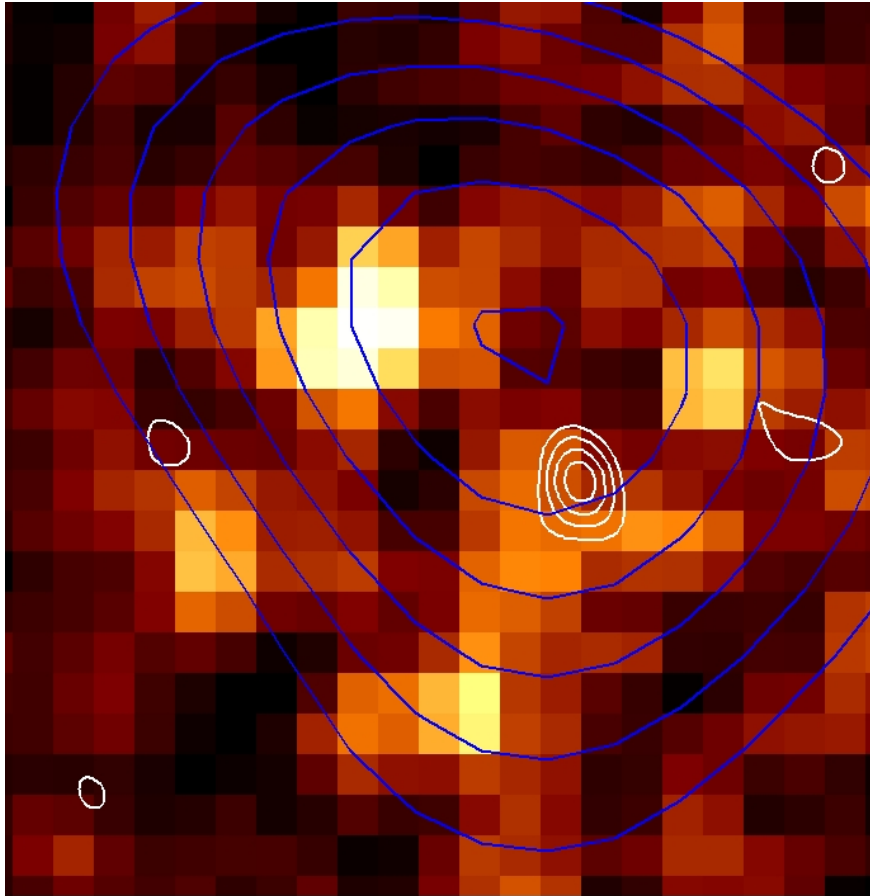
SMA follow-up of bright AzTEC sources in 4C41.17



- SMA target selection: $AzTEC > 4$ mJy at 1.1mm (equivalent to > 10 mJy at 850um)
- SMA requires 4-8 hours for 6-12 sigma detections
- Easily detected all 3 AzTEC sources



SMA/AzTEC/ Spitzer Observations

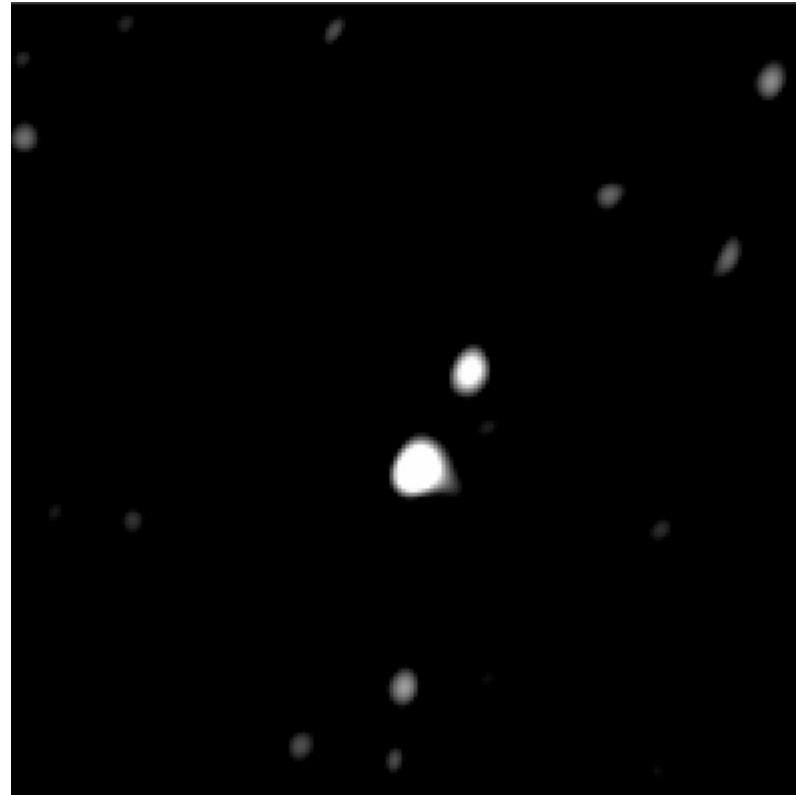
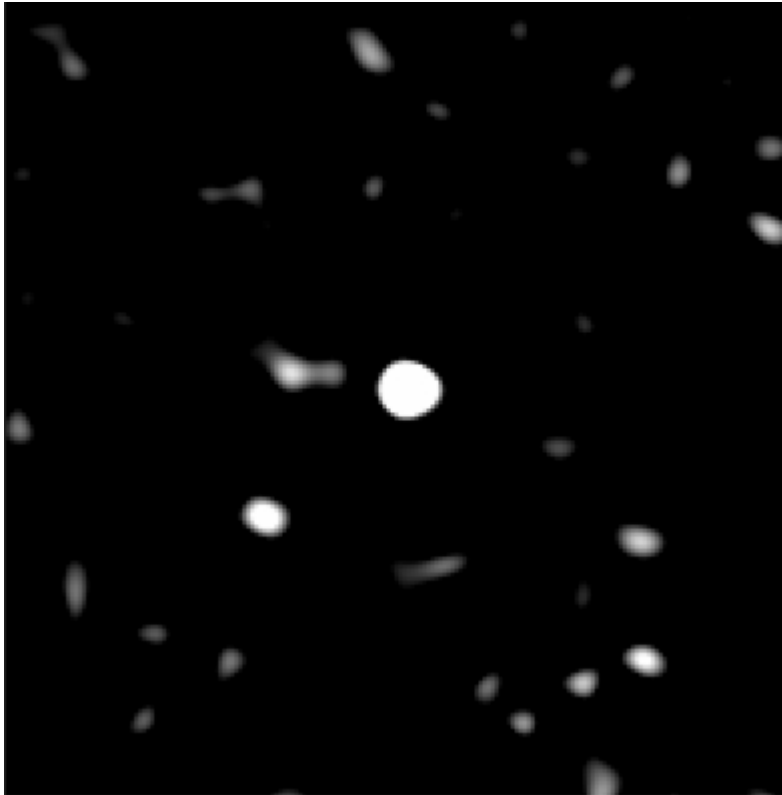


24 microns AzTEC SMA

8 microns SMA



Multiplicity of sub-mm galaxies?



SMA observations of two bright $> 4\text{mJy}$ AzTEC sources toward 4C41.17



Summary

- Significant improvements in the sensitivity and stability of the SMA now permit the observation at 890 microns of galaxies at $z > 3$.
- The SMA is essential in locating the precise position of these high- z galaxies, measuring their size and multiplicity; when combined with multiwavelength observations permits the determination of SEDs/photo- z .
- Brightest submm galaxies detected by SMA may be at high redshift ($z > 3$); important for determining SFR in early Universe.



Future Work

- Very deep Spitzer follow-up (approved)
 - Proposal accepted to image detected radio-quiet sources with IRAC, MIPS, and IRS to get better handle on the redshift
- Ground-based observations with Keck to get redshifts
- Expand the high- z SMG sample
 - Program accepted to do SMA follow-up of AzTEC sources in MS0451; Are SMGs in the field like SMGs in a biased environment?
 - SMA follow-up of next 21 AzTEC/COSMOS bright sources (11 of these sources are radio quiet)
 - SMA follow-up of 10 brightest LABOCA/APEX sources in the ECDFS (with Ian Smail, Fabian Walter and Axel Weiss)
- Resolve SMGs with large SMA baselines
 - Get a better constraint on size scales
 - Compare to merger models
- Theoretical numerical simulations of the formation of SMGs at high z (with Lars Hernquist).