

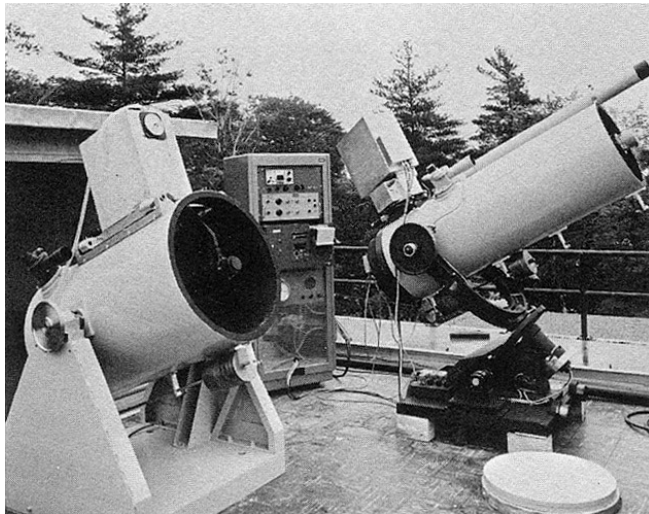
45 Years of Infrared Astronomy at the Air Force Laboratory



Dr. Stephan Price
Space Vehicles Directorate

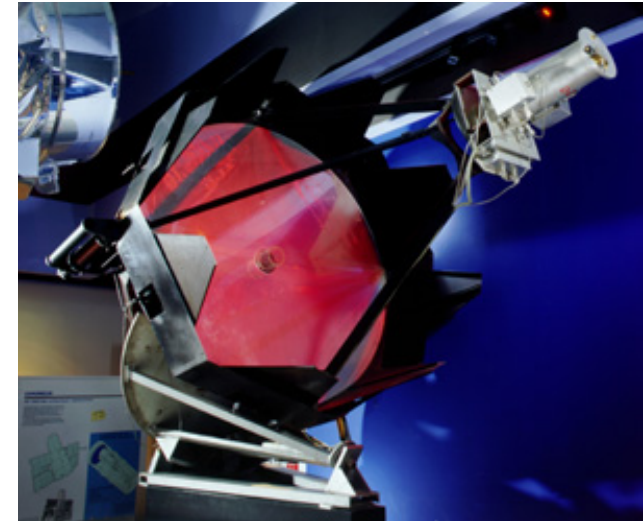


First Near IR Surveys



← ITT telescope

TMSS →



ITT southern survey sensor



- **Freeman Hall conducted the first IR survey (1962)**
- **TMSS (1965 – 1967) by Neugebauer and Leighton**
- **Southern sky (1966 – 1967) by Price**



Infrared Celestial Backgrounds Program



- **Define the nature & detailed character of the infrared celestial background**
- **Probe-rocket based experiments (1970 – 1985)**
 - **20 successes out of 23 attempts**
- **Satellites**
 - **Midcourse Space Experiment (1996 – 1997)**
 - **Observations with ISO and Spitzer**
- **2MASS**
 - **Supported proof of concept study**



First AFCRL Experiments



- **Sensor (top left)**
 - Double folded optics
 - 4" primary mirror
 - Linear array of 6 Mid-IR detectors

- **Two proof of concept flights in 1970**
 - Piggy-backed on an atmospheric experiment
 - Detected Orion Nebula
 - Lessons learned
- **ARPA & AFCRL also provided funds for**
 - Four Cornell rocket-based experiments (1970 – 1976)
 - Caltech & U. of Ariz. for 5 and 10 μm ground-based surveys.



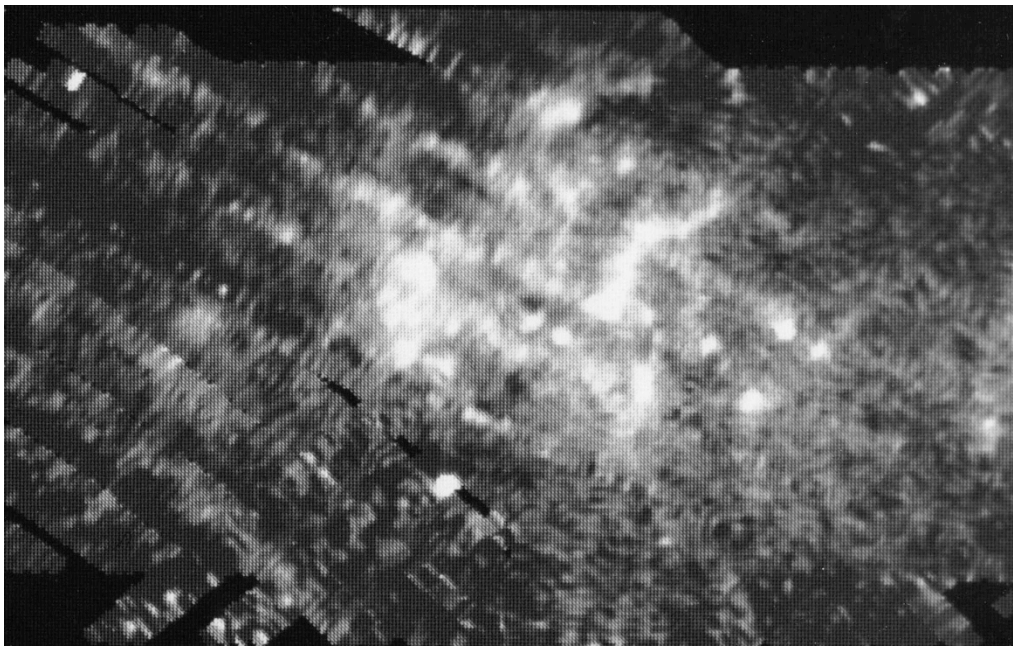
HISTAR & HIStar South



- **First successful mid-IR survey**
 - HISTAR from White Sands
 - April 1971 – Dec 1972 (7 flights in 20 months)
 - HI Star South (Woomera)
 - 1974
 - Southern sky survey
- **Results**
 - 4, 11, 21 and 27 μm point source catalog
 - First large scale maps of the diffuse IR emission from the galactic plane & the zodiacal background



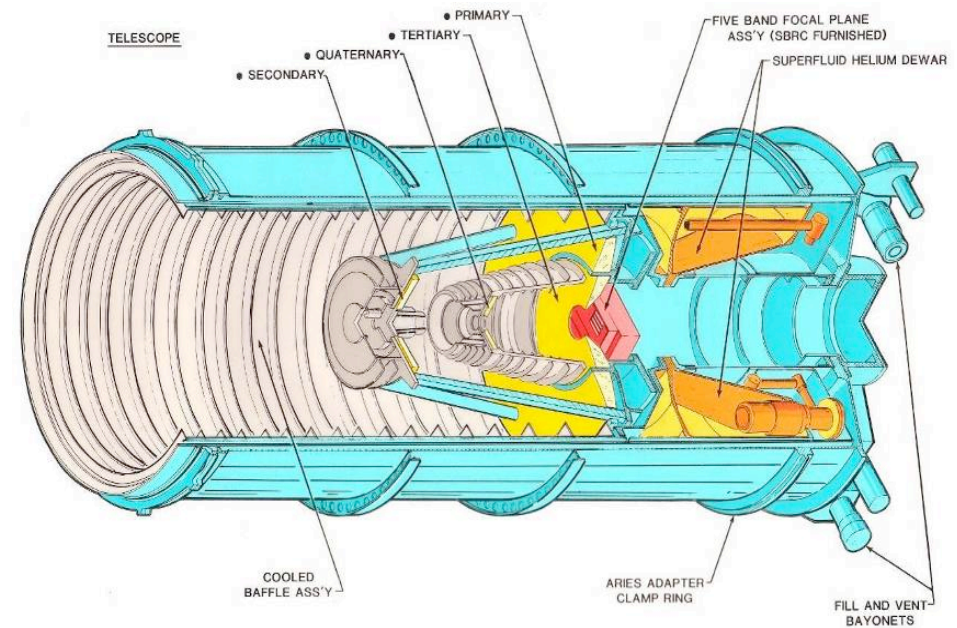
Cygnus X – HISTAR vs MSX



- **HISTAR had ac coupled electronics**
 - Extended emission extracted by digitally inverting the high frequency attenuation
 - A comparison of the HISTAR Cygnus X map to a higher resolution MSX image is shown



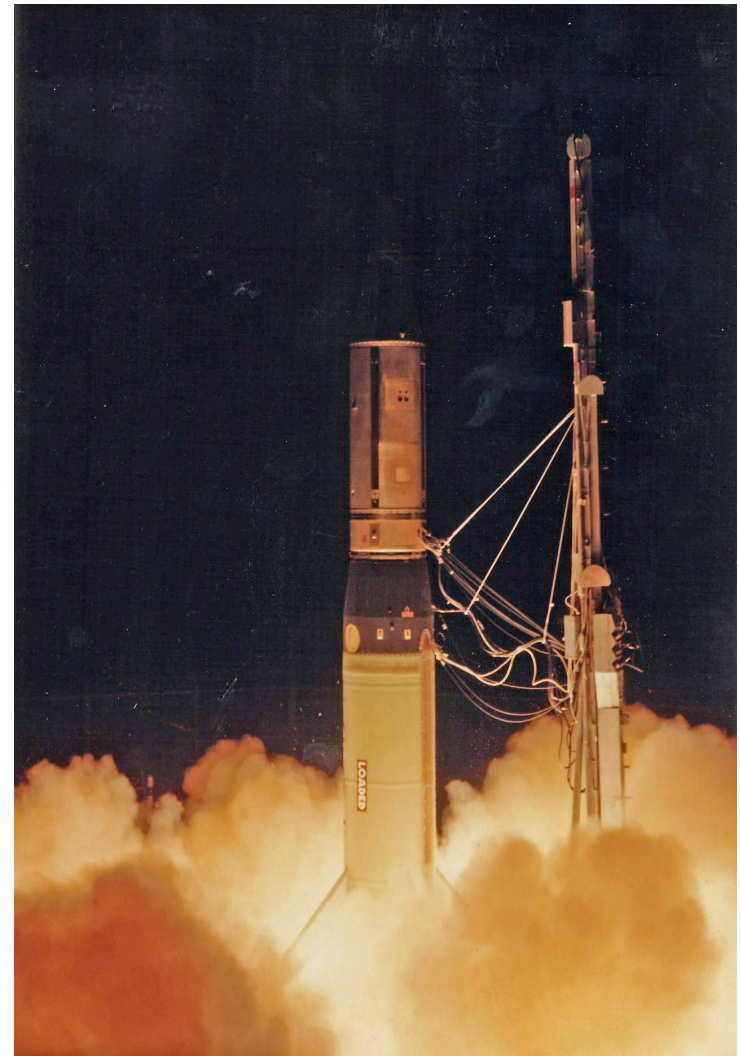
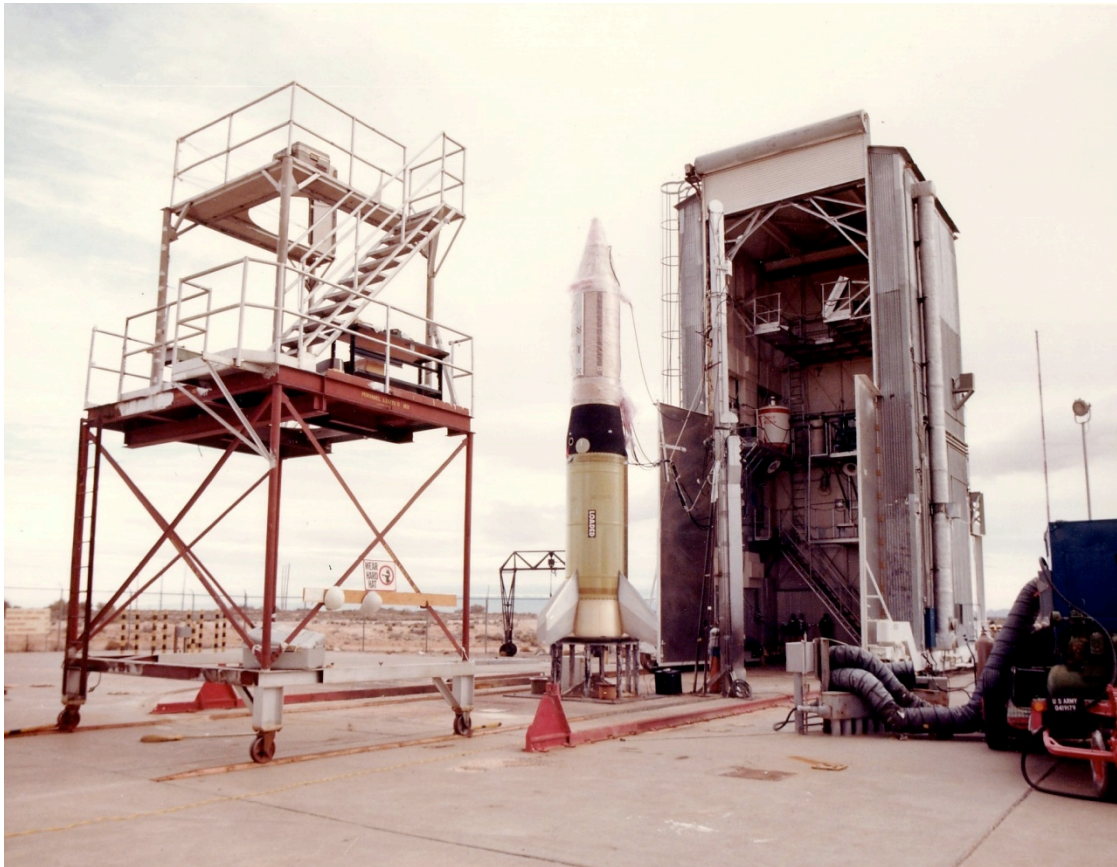
Background Measurements Program



- **35 cm diam. telescopes**
- **SPICE (left)**
 - 11-, 20- & 27 μm
- **FIRSSE (above)**
 - 20, 27, 50 & 90 μm
- **X10 HISTAR sensitivity,**



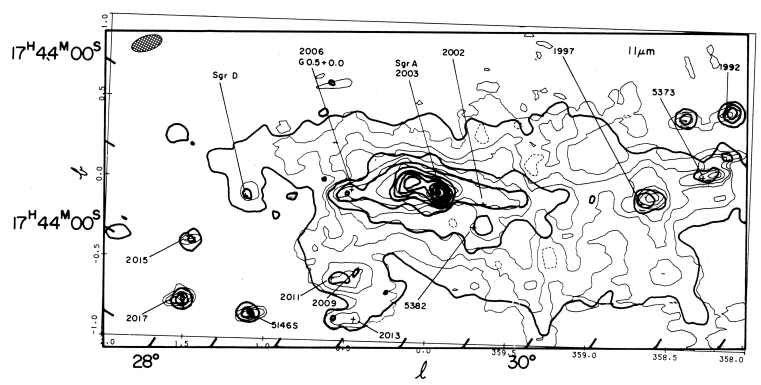
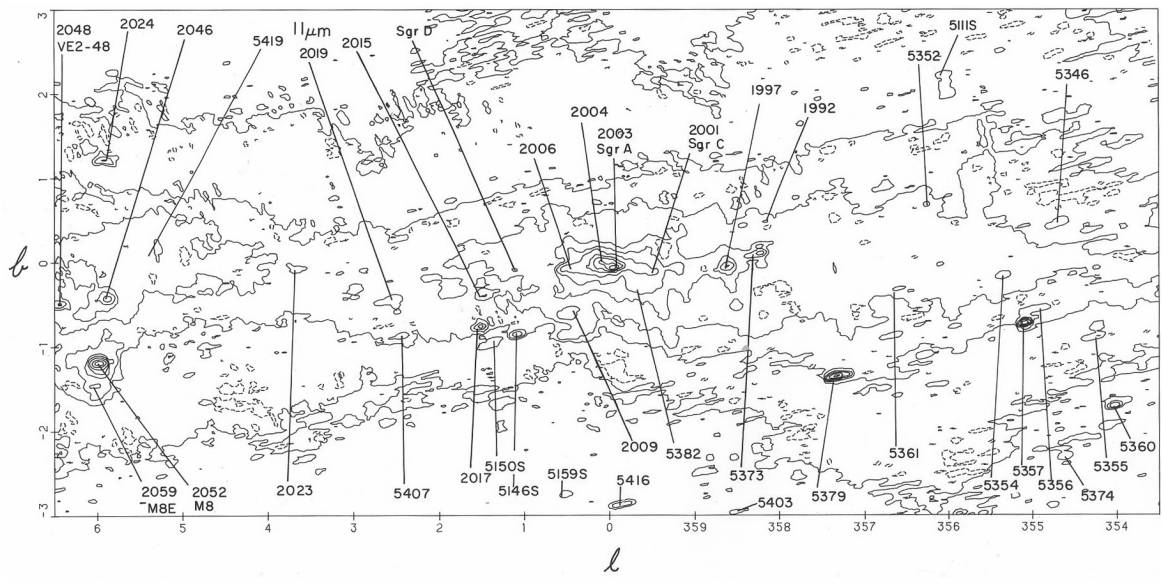
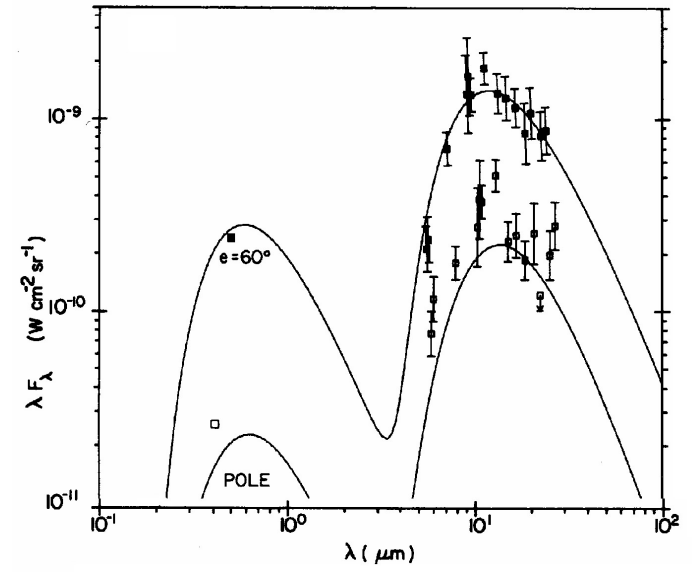
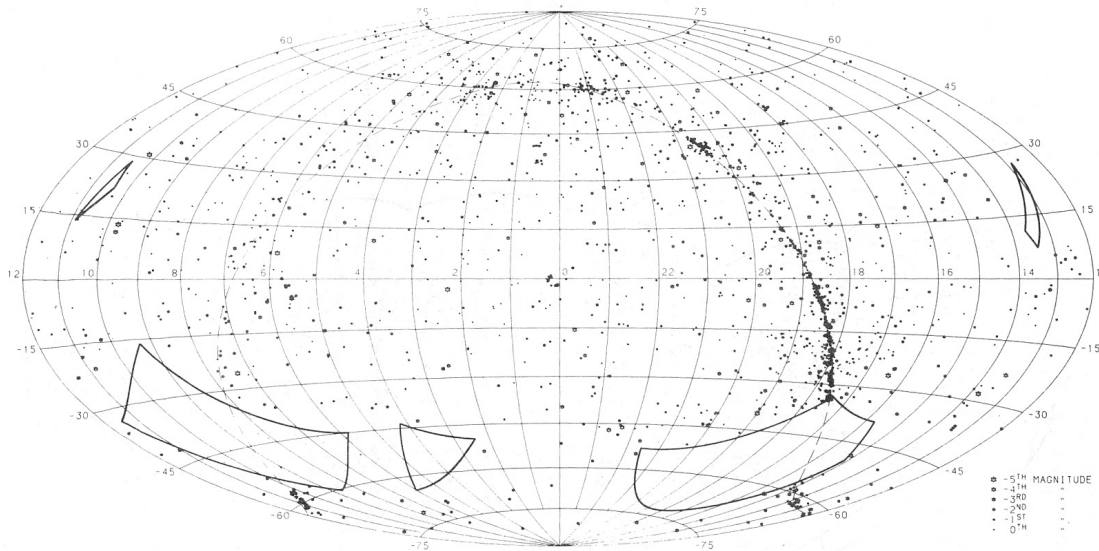
Launch of CB Experiment



- T-3 day rehearsal above
- Launch at left

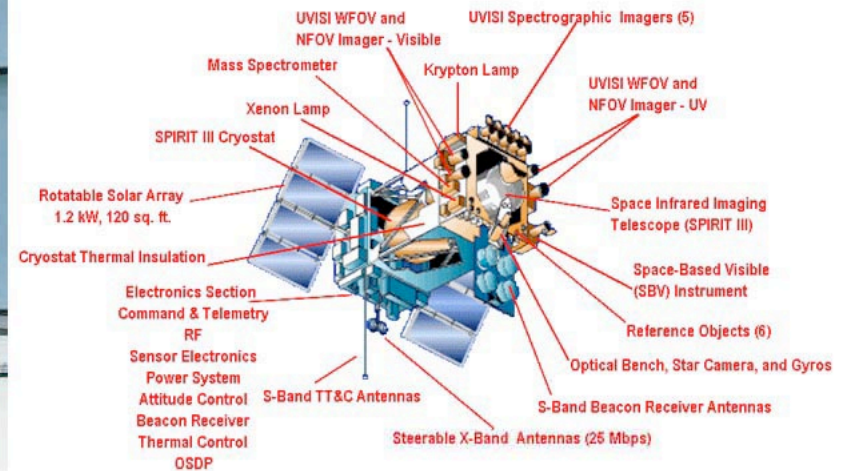
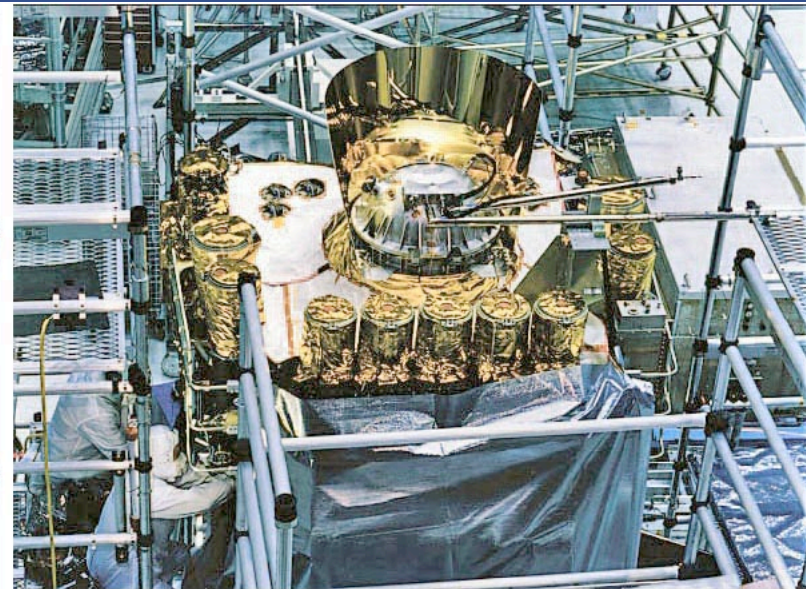
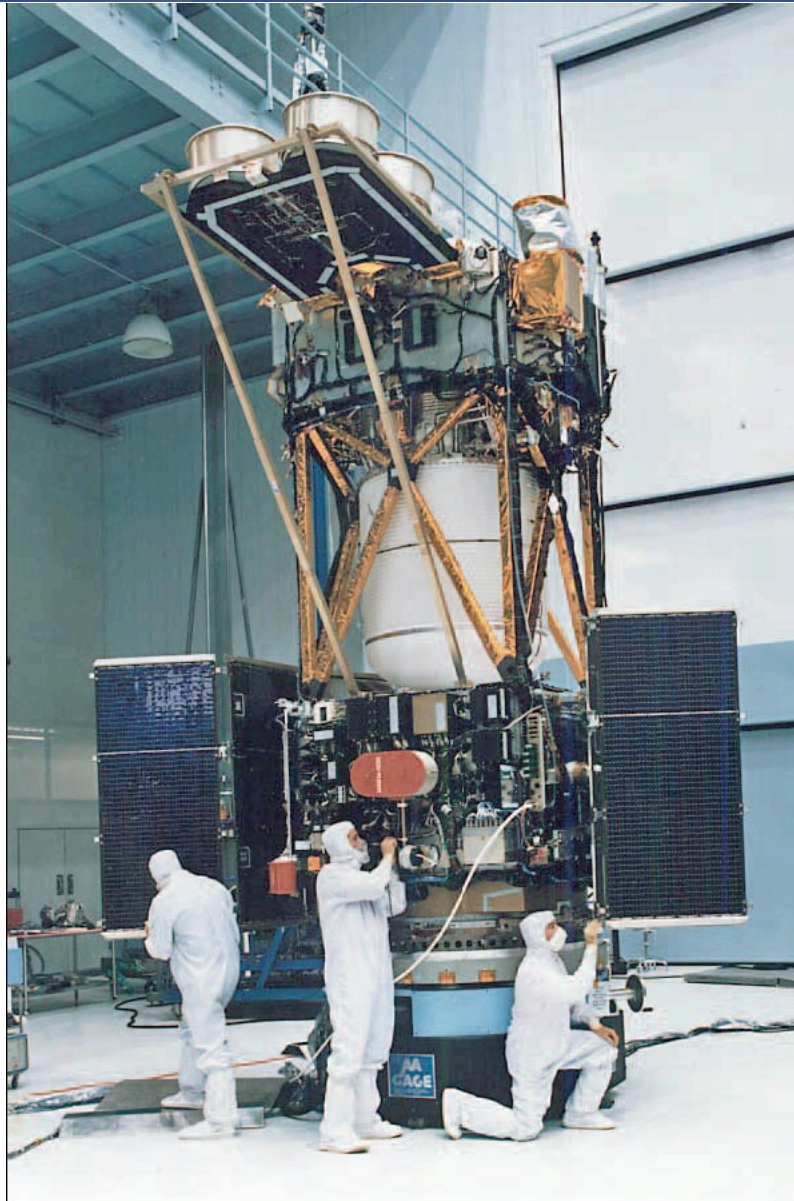


BMP Results



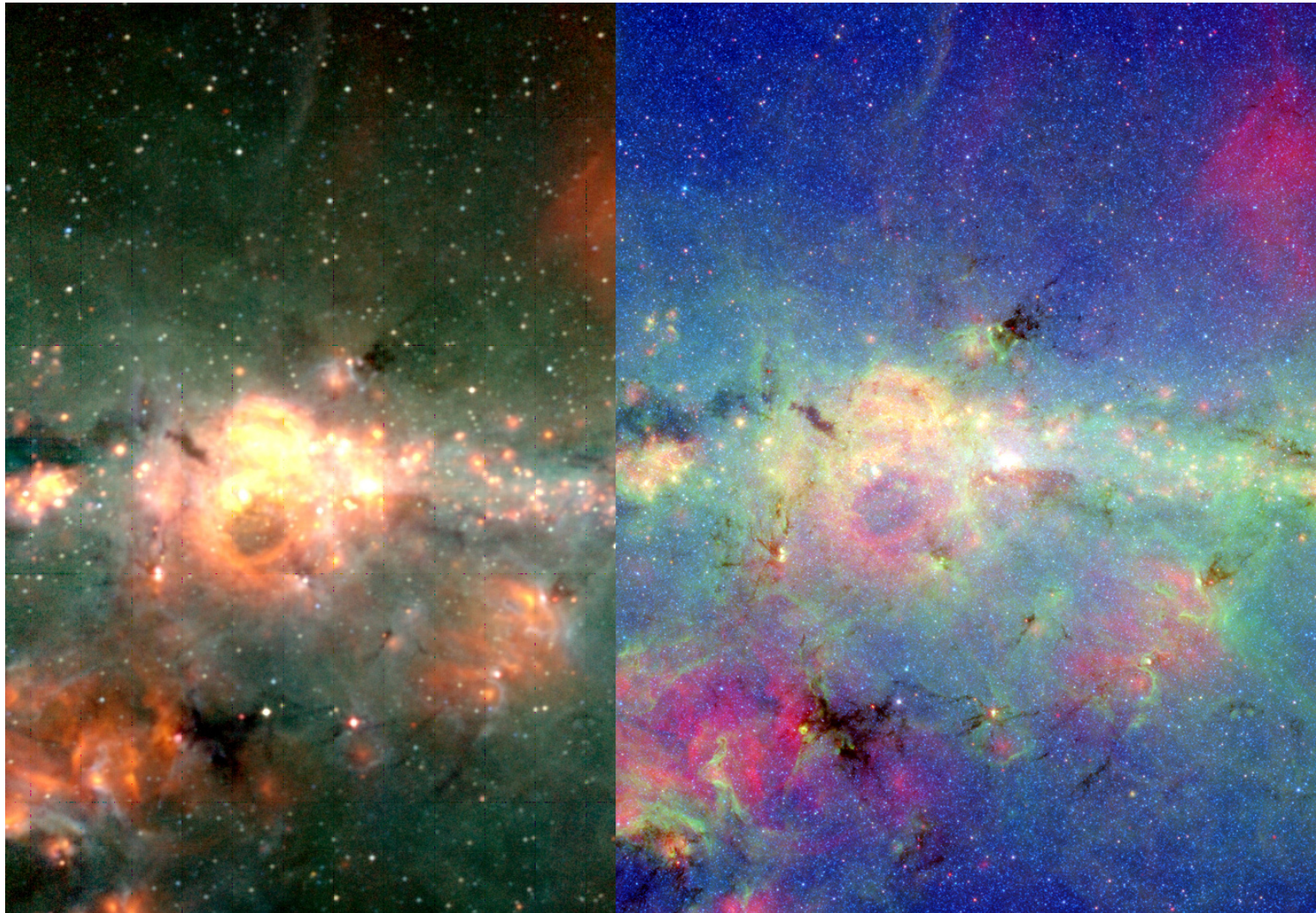


Midcourse Space Experiment (MSX)





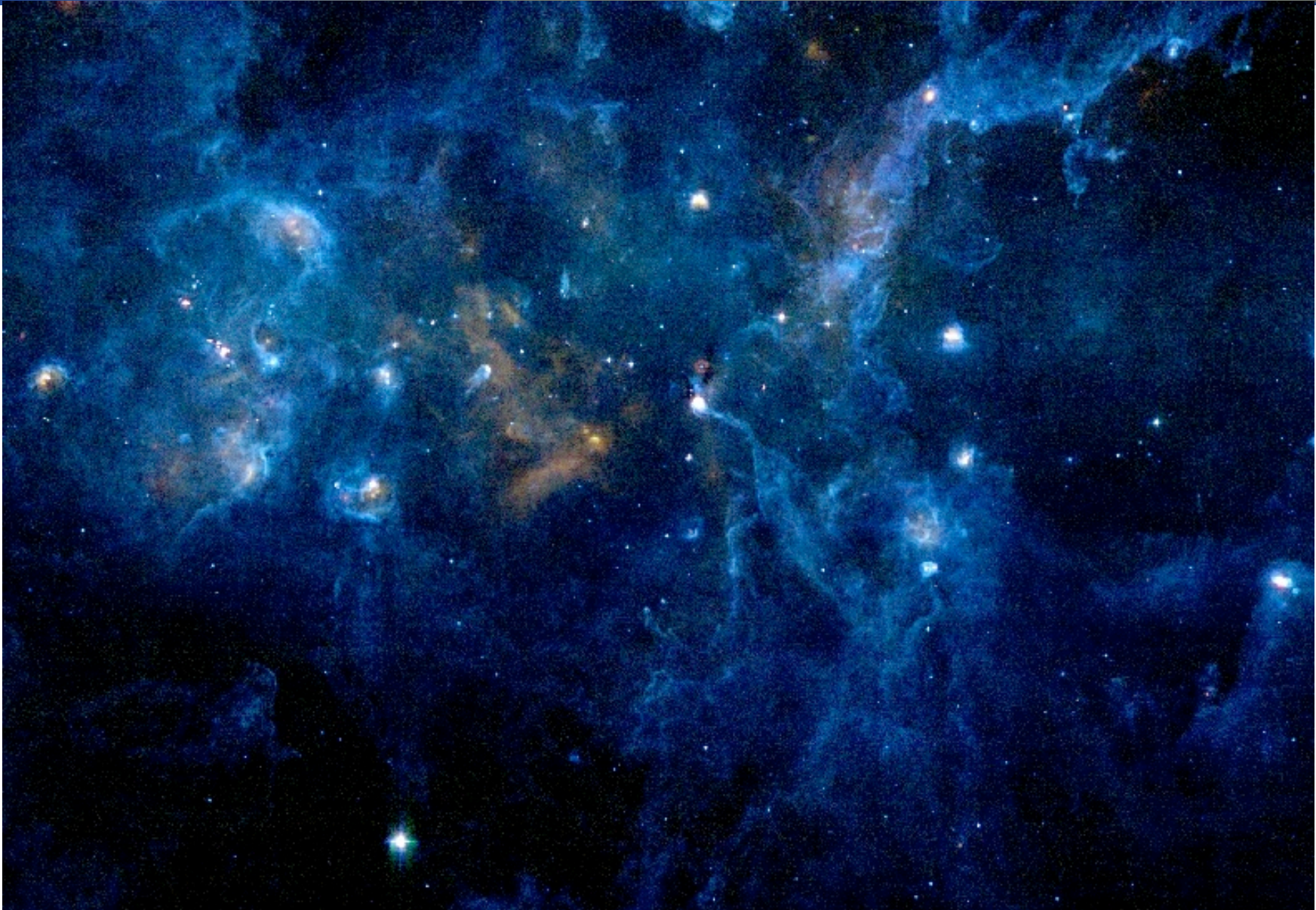
The IR Galactic Center



MSX (left) and Spitzer (right): IR 3 color images of the Galactic center



MSX Cygnus X 3-color Image





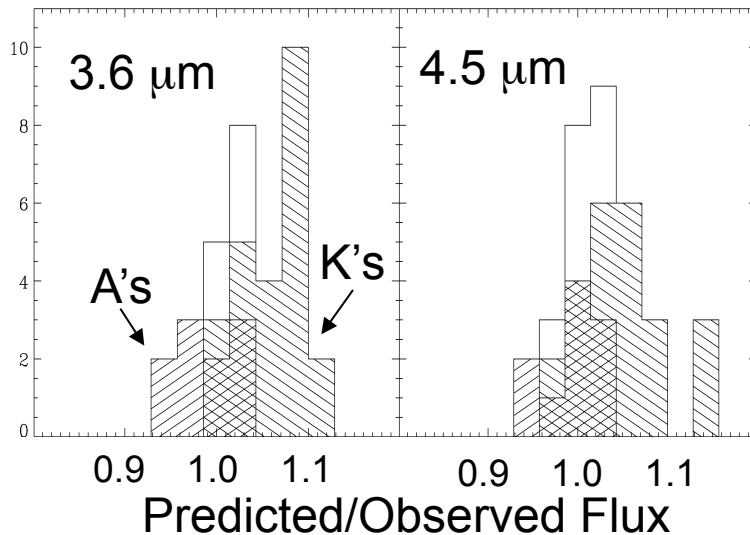
Current Activities



- **Extend the absolute spectral fluxes of the calibration network stars into SWIR & visible to support system calibration at these wavelengths**
- **Upgrade entire calibration network**
 - Create 0.4 – 30 μm spectral templates
 - Apply templates to all tertiary standards
 - Include additional spectral types
 - Add best characterized stars from the Bright Star Atlas to calibration network
- **Thermo-physical lunar model**

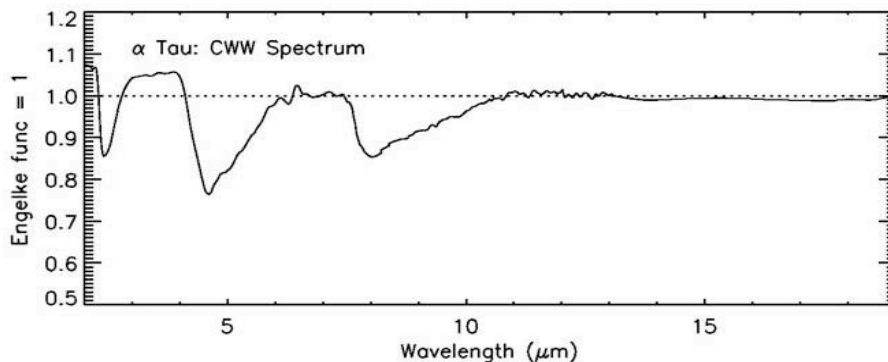


Application: *Spitzer* Space Telescope



***Spitzer* Infrared Array Camera calibration paper*:
Systematic bias between K star calibrators and A star calibrators in the 3.6 and 4.5 μm bands – K stars rejected**

**Original CWW spectrum of α Tau:
bias at wavelengths $<5 \mu\text{m}$**



New SWS+MSX K star spectra remove the bias allowing use of K star calibrators

*Reach et al. 2005



The Moon for Calibration



- **USGS products**

36 narrow spectral bands $0.35 \mu\text{m} < \lambda < 2.39 \mu\text{m}$

- **Extended $\sim 0.5^\circ$ source**
 - Only celestial object beside the Sun for radiance calibration
 - Radiance comparable to that from the Earth – within dynamic range of Earth looking sensors
- **Irradiance calibration for low resolution and low sensitivity instruments**
- **Complicated model needed**
 - Complex viewing geometry due to lunar orbit
 - Complex albedo distribution
 - Thermo-physical model