

# Rotational spectroscopy and equilibrium structures of $S_3$ and $S_4$

S. Thorwirth,<sup>a)</sup> M. C. McCarthy, C. A. Gottlieb, and P. Thaddeus

Harvard-Smithsonian Center for Astrophysics, Cambridge, Massachusetts 02138

and Division of Engineering and Applied Sciences, Harvard University, Cambridge, Massachusetts 02138

H. Gupta and J. F. Stanton

Institute for Theoretical Chemistry, Department of Chemistry and Biochemistry,

University of Texas at Austin, Austin, Texas 78712

(Received 1 April 2005; accepted 4 May 2005; published online 10 August 2005)

The sulfur molecules thiozone  $S_3$  and tetrasulfur  $S_4$  have been observed in a supersonic molecular beam in the centimeter-wave band by Fourier transform microwave spectroscopy, and in the millimeter- and submillimeter-wave bands in a low-pressure glow discharge. For  $S_3$  over 150 rotational transitions between 10 and 458 GHz were measured, and for  $S_4$  a comparable number between 6 and 271 GHz. The spectrum of  $S_3$  is reproduced to within the measurement uncertainties by an asymmetric top Hamiltonian with three rotational and 12 centrifugal distortion constants; ten distortion constants, but an additional term to account for very small level shifts caused by interchange tunneling, are required to reproduce to comparable accuracy the spectrum of  $S_4$ . Empirical equilibrium ( $r_e^{\text{emp}}$ ) structures of  $S_3$  and  $S_4$  were derived from experimental rotational constants of the normal and sulfur-34 species and vibrational corrections from coupled-cluster theory calculations. Quantum chemical calculations show that interchange tunneling occurs because  $S_4$  automerizes through a transition state with  $D_{2h}$  symmetry which lies about 500  $\text{cm}^{-1}$  above the two equivalent  $C_{2v}$  minima on the potential energy surface. © 2005 American Institute of Physics. [DOI: 10.1063/1.1942495]

## I. INTRODUCTION

The large number of allotropic forms of sulfur—exceeding that of any other element—has long stimulated the curiosity of chemists. Besides a number of chain configurations, cyclic sulfur allotropes from  $S_6$  to  $S_{20}$  have been isolated and characterized (mostly in the solid phase) by various techniques, including x-ray diffraction and Raman and infrared spectroscopy,<sup>1</sup> but very little structural information has been available for small clusters except  $S_2$ . The most stable forms of thiozone ( $S_3$ ) and tetrasulfur ( $S_4$ ) were recently detected here by Fourier transform microwave (FTM) spectroscopy.<sup>2,3</sup> The present investigations (this work and Refs. 2 and 3) mark the first detection of  $S_3$  and  $S_4$  by means of high-resolution spectroscopy and the determination of precise molecular structures for both molecules.

Sulfur clusters pose challenging problems to quantum chemistry, and a large number of theoretical papers on them have appeared.<sup>4</sup> For  $S_3$  the structure with  $C_{2v}$  symmetry shown in Fig. 1 is predicted to be more stable than a cyclic  $D_{3h}$  structure by about 5–10 kcal/mol. For  $S_4$  there is some ambiguity, since the structures and relative energies of the isomers are quite sensitive to the level of theory, and several different isomers have been calculated to be the global minimum on the potential energy surface. Recent calculations suggest that the  $C_{2v}$  structure in Fig. 1 is the most stable,<sup>5,6</sup> but a rectangular  $D_{2h}$  isomer which has been found to be a transition state for interchange tunneling is only a few kcal/mol higher in energy. There is clear evidence for interchange

tunneling in  $S_4$  in the FTM data discussed here, but it is not possible to determine the barrier height from this data alone.<sup>3</sup>

Here a joint experimental and theoretical investigation of  $S_3$  and  $S_4$  is presented. The new laboratory and computational work amplifies our preliminary experimental results and extends the quantum calculations to coupled clusters methods. The present paper consists of (i) a comprehensive study of the rotational spectra of  $S_3$  and  $S_4$  in the centimeter and submillimeter bands to frequencies as high as 458 GHz; (ii) a study of the rotational spectra of the single sulfur-34

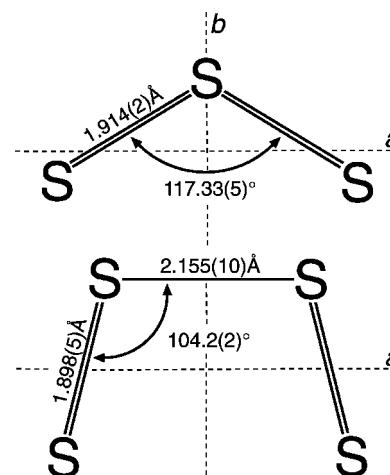


FIG. 1. Empirical equilibrium structures ( $r_e^{\text{emp}}$ ) of  $S_3$  and  $S_4$ . Bond lengths and angles were derived from experimental rotational constants which were converted to equilibrium constants with theoretical vibration-rotation interaction constants (see Sec. V). Estimated  $1\sigma$  uncertainties are given in parentheses. The bond orders are approximate.

<sup>a)</sup>Electronic mail: sthorwirth@cfa.harvard.edu

isotopic species; (iii) a study of the geometrical structures, dipole moments, and vibrational energies of ground-state  $S_3$  and  $S_4$ ; (iv) detection of vibrational satellites belonging to  $S_3$  and their assignment on the basis of a calculated cubic force field; and finally, (v) a theoretical investigation of the double minimum potential of  $S_4$ .

## II. EXPERIMENT

### A. Centimeter-wave measurements

The FTM spectrometer used for the present work operates from 5 to 43 GHz.<sup>7,8</sup> Reactive molecules are produced in the throat of a small supersonic nozzle by a low-current dc discharge through various precursor gases heavily diluted in a noble buffer gas. Free expansion from the nozzle into the large vacuum chamber of the spectrometer then forms a Mach 2 supersonic beam with a rotational temperature of only 1–3 K. Following excitation of a rotational transition by a short microwave pulse, radiation from the coherently rotating molecules is detected with a sensitive heterodyne receiver.

The initial detection of  $S_3$  was in a discharge of hydrogen sulfide ( $H_2S$ ) heavily diluted to about 0.2% in neon, chosen because  $H_2S$  is well known to be a good source of compounds such as  $HS_2$ ,  $H_2S_2$ , and  $H_2S_3$ .<sup>9</sup> Initially, a search for the  $2_{1,1}-2_{0,2}$  transition of  $S_3$  was undertaken in a fairly wide band between 20 and 25 GHz. Assays performed on candidate lines established that a line near 21.7 GHz was only seen as a discharge product and exhibited no detectable Zeeman effect when a permanent magnet was brought near the molecular beam—as expected for a closed-shell molecule such as  $S_3$ . This line was also present when  $D_2S$  replaced  $H_2S$  and argon replaced neon, ruling out hydrogen containing discharge products or a van der Waals complex with the buffer gas. Owing to the high signal-to-noise ratio, the corresponding lines of the two singly substituted sulfur-34 isotopic species were then found in spite of the low natural abundance of  $^{34}S$  (4.2%). The close agreement (to better than 1%) between the observed frequency shifts for the rare isotopic species and those predicted from the structure in Fig. 1 is compelling evidence that the 21.7-GHz line is from  $S_3$  and no other pure sulfur cluster.

Following spectroscopic identification, the production of  $S_3$  from various precursors was investigated. Five were tested, yielding the relative line intensities and abundances in Table I. Among those which are either gases at room temperature (OCS,  $SO_2$ , and  $H_2S$ ) or a highly volatile liquid ( $CS_2$ ),  $H_2S$  was found to be the best source of  $S_3$ , presumably because its primary decomposition products are  $SH$  ( $^2\Pi$ ) and  $S$  in its  $^3P$  ground state, formed by sequential loss of hydrogen;<sup>10</sup> subsequent reactions of either can form small sulfur clusters and  $S_nH_m$  species, including the closely related sulfur chains  $HS_2$  and  $HS_3H$ .<sup>9</sup> Although decomposition of the other precursors tested is also known to produce atomic sulfur,<sup>11</sup> each yields much weaker lines of  $S_3$  (by at least a factor of 10) relative to  $H_2S$ . The lower production of  $S_3$  with these precursors may correlate with the amount of ground-state  $S(^3P)$  in the discharge. In the flash photolysis of OCS, for example, nearly three-fourths of the S atoms are

TABLE I. Relative line intensities and abundances of  $S_3$  produced from different precursors.

Precursor <sup>a</sup>	Relative intensity <sup>b</sup>	Relative abundance
$H_2S$	1.0	1.0
$SO_2$	0.1	0.1
$CS_2$	0.02	0.02
OCS	<0.02	<0.02
gaseous $S_8$	12	25 <sup>c</sup>

<sup>a</sup>Each gas was heavily diluted (99–99.9%) in neon.

<sup>b</sup>From measurements of the intensity of the  $2_{1,1}-2_{0,2}$  transition at 21 679.9 MHz.

<sup>c</sup>The relative abundance is about two times higher than the relative intensity, owing to the higher rotational temperature of  $S_3$  produced from gaseous  $S_8$  in the heated nozzle ( $T_{rot} \sim 4$  K) compared with that of the other precursors ( $T_{rot} \sim 2$  K).

produced in the  $^1D$  excited state<sup>12</sup> and many collisions (roughly 20 when  $CO_2$  is the buffer gas) are required to deactivate  $S(^1D)$  to the ground state.<sup>13</sup> If  $S_3$  is formed stepwise from  $S(^3P)$  and only 1/4 of the sulfur atoms are in this state, the mole fraction of  $S_3$  may be reduced by up to a factor of roughly  $(1/4)^3 = 1/64$  for OCS compared to  $H_2S$ , even if the mole fraction of S is the same.

By far the best source for  $S_3$  was found to be a discharge through sulfur vapor (0.1%) in neon produced by flowing the neon over sulfur powder resistively heated to 185 °C (see Appendix); with a discharge voltage of 600–700 V, this arrangement yielded lines stronger by about a factor of 12 relative to those from  $H_2S$  (Table I). As an example of the high abundance of  $S_3$  in our molecular beam, a spectrum of the  $1_{1,1}-0_{0,0}$  transition of the doubly substituted sulfur-34 species  $^{34}S\,^{34}SS$  was observed in natural abundance (0.3%) (Fig. 2). This large enhancement of  $S_3$  over that found with other sulfur precursors suggests that in addition to sequential formation from atomic and diatomic sulfur,  $S_3$  may be formed directly by fragmentation of  $S_8$ .

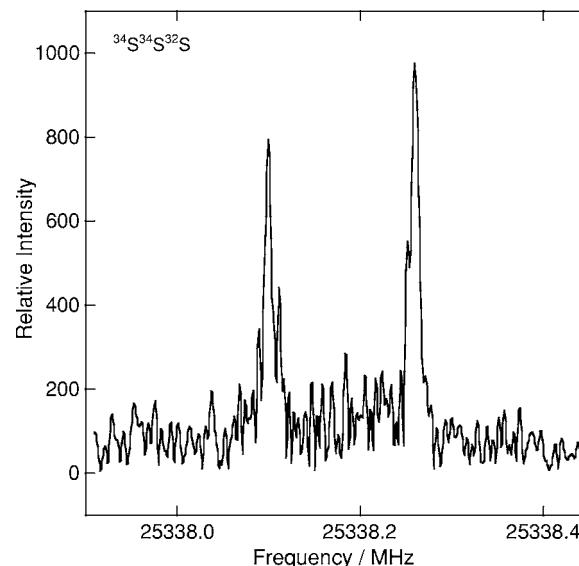


FIG. 2. The  $1_{1,1}-0_{0,0}$  transition of  $^{34}S\,^{34}SS$  at 25 338.2 MHz observed in natural abundance in a supersonic beam with an integration time of 14 min. The abundance of the doubly substituted sulfur-34 species is  $\sim 3 \times 10^{-3}$  that of the normal isotopic species.

Following detection and spectroscopic characterization of S<sub>3</sub>, a search for rotational lines of S<sub>4</sub> was undertaken in the sulfur vapor discharge. S<sub>4</sub> is predicted to be even more polar than S<sub>3</sub>, and to have a strong rotational transition (1<sub>1,1</sub>–0<sub>0,0</sub>) near 6.1 GHz. A strong line near that frequency was detected which was so intense that we were able to find the two singly substituted sulfur-34 counterparts in natural abundance at almost exactly the predicted isotope shifts.

## B. Millimeter-wave measurements

The millimeter-wave spectra of S<sub>3</sub> and S<sub>4</sub> were measured in absorption in a free-space glass discharge cell (1.5 m long, 7 cm in diameter) with a spectrometer described elsewhere.<sup>14</sup> Small sulfur clusters were produced by heating elemental sulfur *in situ* in a low-pressure dc discharge through a steady stream of argon. Lines of S<sub>3</sub> appear when the temperature of the discharge cell (50–60 °C) is well below the melting point of sulfur. At that temperature the vapor pressure of sulfur is very low (<1 mTorr), but the sublimation rate is fairly high.<sup>15</sup>

Spectra were recorded with the cell heated over its full length to 70 °C and carrying a discharge current of 20 mA, with an argon flow of 9 cm<sup>3</sup>/min at STP, and a total pressure (with the discharge on) of 45 mTorr. Lines of S<sub>3</sub> were then quite strong: for example, near 350 GHz, a signal-to-noise ratio S/N>100 could be achieved with an integration time of less than 5 min. Although the dipole moment (the dipole moments calculated at the CCSD(T)/cc-pVTZ level of theory are 0.51 D for S<sub>3</sub> and 1.15 D for S<sub>4</sub>) of S<sub>4</sub> is 2.3 times larger than that of S<sub>3</sub>, lines of S<sub>4</sub> observed in the same discharge were roughly ten times weaker than those of S<sub>3</sub> because the rotational and the vibrational partition functions are 3.5 times larger than that of S<sub>3</sub>, and the mole fraction of S<sub>4</sub> (2 × 10<sup>-4</sup>) is 4.5 times smaller than that of S<sub>3</sub> (9 × 10<sup>-4</sup>). Further increase in the intensity of the S<sub>3</sub> lines was observed when the temperature was raised still higher to 90–95 °C, but the discharge then became unstable, to the point that good spectra could not be obtained.

At 70 °C sulfur vapor consists almost entirely of S<sub>8</sub>, but Hassanzadeh and Andrews<sup>16</sup> showed that S and S<sub>2</sub> are produced at that temperature in a discharge very similar to ours (see Fig. 1 in Ref. 17). We were unable to observe atomic S because the lowest-frequency fine-structure transition (<sup>3</sup>P<sub>1</sub>–<sup>3</sup>P<sub>0</sub>) is in the far IR at 5.3 THz, but we have observed fairly intense spectra of three magnetic dipole transitions of S<sub>2</sub>: (N,J)=(13,13–13,12) at 163.6 GHz, (15,15–15,14) at 178.4 GHz, and (17,17–17,16) at 191.3 GHz. From these spectra we estimate a mole fraction of S<sub>2</sub> of ~0.01, implying that (i) roughly 10% of the S<sub>8</sub> in the vapor phase is converted to S<sub>2</sub> in the discharge, and (ii) the mole fraction of S<sub>2</sub> is ten times higher than that of S<sub>3</sub> and 50 times higher than that of S<sub>4</sub>.

Intense lines of S<sub>2</sub>O and SO<sub>2</sub> were also observed when the concentrations of S<sub>3</sub> and S<sub>4</sub> in the glow discharge were high. The mole fraction of S<sub>2</sub>O was only five times less than that of S<sub>3</sub>, probably because of residual sources of oxygen in the absorption cell. Rotational lines were readily identified

for SO<sub>2</sub> by consulting the standard spectral line catalog,<sup>18</sup> and for S<sub>2</sub>O from predictions based on published measurements.<sup>19</sup>

## III. QUANTUM CHEMICAL CALCULATIONS

The structural parameters of S<sub>3</sub> and S<sub>4</sub> required to interpret the present laboratory spectroscopic data were calculated with coupled-cluster theory,<sup>20</sup> and the correlation-consistent basis sets of Dunning.<sup>21</sup> Equilibrium geometries were obtained with the CCSD(T) treatment of correlation<sup>22</sup> and the cc-pVTZ basis set using analytical energy derivatives.<sup>23</sup> Anharmonic force fields were calculated by numerical differentiation of analytic second derivatives<sup>24</sup> evaluated at points displaced from equilibrium along the dimensionless normal coordinates defined by the quadratic force field. Details of this procedure are given in Ref. 25. The anharmonic force fields for S<sub>3</sub> and S<sub>4</sub> were then used in conjunction with standard rovibrational perturbation theory<sup>26</sup> to calculate the fundamental vibrational frequencies and vibration-rotation interaction constants  $\alpha_i^{A,B,C}$ .

To investigate the fluxional nature of S<sub>4</sub>, which automerizes through a transition state with D<sub>2h</sub> symmetry, further calculations were undertaken. First, the D<sub>2h</sub> structure was optimized with CCSD(T) and the cc-pVTZ basis set to estimate the barrier height. Additional calculations were carried out to study the pathway for automerization. In these, the S–S–S angle was varied over a grid of points and the remaining structural parameters optimized using analytic derivatives.

All calculations reported in this work were performed with a local version of the ACES II program system.<sup>27</sup> The treatment of correlation included all electrons in the evaluation of the molecular parameters (equilibrium structures, vibrational frequencies, and vibration-rotation interaction constants); only the valence electrons were correlated in the more qualitative exploration of the S<sub>4</sub> automerization potential.

## IV. RESULTS

Both S<sub>3</sub> and S<sub>4</sub> have C<sub>2v</sub> symmetry (Fig. 1) in the isomers calculated to be most stable at the highest level of theory studied.<sup>4</sup> The rotational spectra are those of a nearly prolate (S<sub>3</sub>) or somewhat prolate (S<sub>4</sub>) asymmetric top ( $\kappa$ =-0.97 and -0.54, respectively), with *b*-type transitions and Bose–Einstein statistics. As a result, the allowed rotational levels are those that are even upon nuclear permutations (i.e., with  $K_aK_c=ee$  or *oo*)—a restriction relaxed in the singly substituted isotopic species where the C<sub>2v</sub> symmetry is broken.

### A. Thiozone, S<sub>3</sub>

#### 1. Rotational spectrum

With the FTM spectrometer, 17 lines of the main isotopic species of S<sub>3</sub> with  $J\leq 12$  and  $K_a\leq 2$  were measured in the centimeter-wave band between 9 and 40 GHz [see Table I of supplementary material (SM), deposited in the Electronic Physics Auxiliary Publication Service (EPAPS) of the

TABLE II. Spectroscopic constants of thiozone,  $S_3$  (in MHz).

Constant <sup>a</sup>	$^{32}S_3$		$S\ ^{34}SS$	$^{34}SS_2$
	Measured <sup>b</sup>	Theoretical <sup>c</sup>	Measured <sup>d</sup>	Measured <sup>d</sup>
$A$	23 972.5807(4)	23 159.	23 033.906(2)	23 738.354(2)
$B$	2948.546 79(7)	2885.	2948.6856(6)	2861.0127(4)
$C$	2622.291 12(7)	2562.	2610.6946(5)	2550.1327(4)
$10^3 D_J$	0.799 67(3)	0.77	0.79(1)	0.754(8)
$10^3 D_{JK}$	-20.279(1)	-19.	-19.1(2)	-19.7(2)
$D_K$	0.503 44(1)	0.45	0.503 <sup>e</sup>	0.503 <sup>e</sup>
$10^3 d_1$	-0.163 171(7)	-0.16	-0.167(3)	-0.151(2)
$10^6 d_2$	-7.232(4)	-6.6	-7.23 <sup>e</sup>	-7.23 <sup>e</sup>
$10^9 H_J$	0.460(4)			
$10^9 H_{JK}$	-9.17(9)			
$10^6 H_{KJ}$	-1.59(1)			
$10^3 H_K$	0.034 02(8)			
$10^9 h_1$	0.216(1)			
$10^9 h_2$	0.0242(9)			
$10^{12} h_3$	6.1(2)			
$\Delta^f$	0.243		0.248	0.245

<sup>a</sup>The  $1\sigma$  uncertainties (in parentheses) are in the units of the last significant digits.

<sup>b</sup>Derived from a least-squares fit of Watson's  $S$ -reduced Hamiltonian to the data in Table I of SM (Ref. 28).

<sup>c</sup>Calculated at the CCSD(T)/cc-pVTZ level of theory.

<sup>d</sup>Derived from a least-squares fit of Watson's  $S$ -reduced Hamiltonian to the data in Table II of SM (Ref. 28).

<sup>e</sup>Constrained to value derived for the normal isotopic species.

<sup>f</sup>Inertial defect,  $\Delta = I_c - I_b - I_a$ , in units of amu  $\text{\AA}^2$ .

American Institute of Physics].<sup>28</sup> These were reproduced to within the measurement uncertainties (2–5 kHz) with eight spectroscopic constants in Watson's  $S$ -reduced Hamiltonian: three rotational and five fourth-order centrifugal distortion constants. Two singly substituted sulfur-34 species were also investigated: nine lines for  $S\ ^{34}SS$  and 16 lines for  $^{34}SS_2$ , including previously forbidden transitions now allowed by the broken symmetry [Table II of SM (Ref. 28)]. For both species, only three of the five fourth-order distortion constants ( $D_J$ ,  $D_{JK}$ , and  $d_1$ ) were determined ( $D_K$  and  $d_2$  were constrained to the values for the normal species).

To better determine the centrifugal distortion,  $S_3$  was then studied in the millimeter- and submillimeter-wave bands. The first lines observed were series of  $R$ -branch transitions with low  $K_a$  in the region of 170–191 GHz, found

within a few megahertz of predictions based on the FTM data. In all, 152 lines of  $S_3$  with  $J \leq 87$  and  $K_a \leq 12$  were measured at frequencies as high as 458 GHz [Table 1 of SM (Ref. 28)], allowing the determination of all seven sixth-order centrifugal distortion constants. The rotational and centrifugal distortion constants of  $S_3$  are given in Table II.

## 2. Vibrational excitation

Although rotational cooling to only a few kelvin is achieved in our molecular beam, vibrational cooling is much less efficient, and rotational lines in vibrationally excited states (vibrational satellites) are frequently observed in FTM

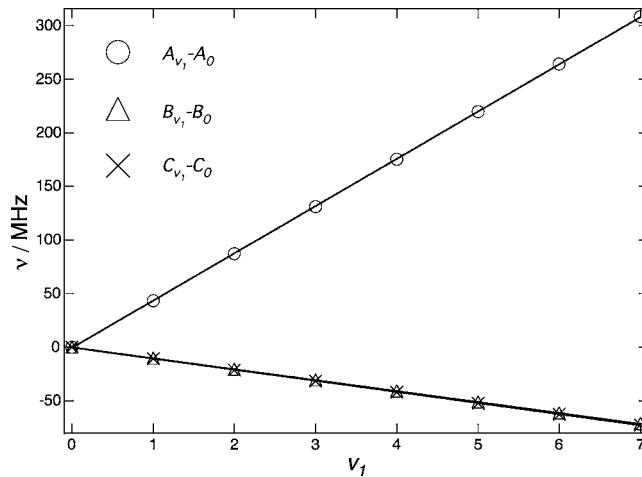


FIG. 3. Differences in rotational constants of  $S_3$  in the ground and the vibrationally excited levels of the  $\nu_1$  mode vs vibrational quantum  $\nu_1$ . Experimental values for  $\alpha_1^A$ ,  $\alpha_1^B$ , and  $\alpha_1^C$  determined from the slopes are in very good agreement with those calculated theoretically (see Sec. IV A 2).

TABLE III. Rotational constants of  $S_3$  in the  $\nu_1$  vibrational mode (in MHz).

Vibrational state	Constant <sup>a</sup>			
	$A$	$B$	$C$	$\Delta^b$
$1\nu_1$	24 015.954	2938.279	2611.552	0.475
$2\nu_1$	24 059.681	2928.031	2600.944	0.700
$3\nu_1$	24 103.706	2917.801	2590.461	0.920
$4\nu_1$	24 147.964	2907.595	2580.096	1.134
$5\nu_1$	24 192.353	2897.408	2569.844	1.343
$6\nu_1$	24 236.735	2887.243	2559.697	1.547
$7\nu_1$	24 280.914	2877.101	2549.650	1.746

<sup>a</sup>Rotational constants derived from a least-squares fit of Watson's  $S$ -reduced Hamiltonian to the data in Table III of SM (Ref. 28) with centrifugal distortion constants constrained to those in the ground vibrational state (see Table II).

<sup>b</sup>The inertial defect,  $\Delta = I_c - I_a - I_b$ , in units of amu  $\text{\AA}^2$ .

TABLE IV. Theoretical vibration-rotation interaction constants of S<sub>3</sub> (in MHz) calculated at the CCSD(T)/cc-pVTZ level of theory.

Vib. mode and $\Sigma_i \alpha_i / 2$	SSS		
	$\alpha_i^A$	$\alpha_i^B$	$\alpha_i^C$
$\nu_1$	-47.06	+9.93	+10.34
$\nu_2$	-242.73	-0.60	+2.42
$\nu_3$	+199.66	+7.59	+6.18
$\Sigma_i \alpha_i / 2$	-45.07	+8.46	+9.47
34SSS			
	$\alpha_i^A$	$\alpha_i^B$	$\alpha_i^C$
$\nu_1$	-48.87	+9.58	+10.00
$\nu_2$	-236.40	-0.59	+2.27
$\nu_3$	+197.18	+7.31	+5.98
$\Sigma_i \alpha_i / 2$	-44.04	+8.15	+9.12
S 34SS			
	$\alpha_i^A$	$\alpha_i^B$	$\alpha_i^C$
$\nu_1$	-38.90	+9.74	+10.09
$\nu_2$	-233.85	-0.57	+2.51
$\nu_3$	+186.51	+7.46	+6.05
$\Sigma_i \alpha_i / 2$	-43.12	+8.32	+9.32

experiments.<sup>29</sup> Vibrational satellites up to  $v=7$  were observed for several rotational transitions of the normal species of S<sub>3</sub> [Table III of SM (Ref. 28)], and from these sets of rotational constants were determined (Table III).

The vibrational dependence of the difference in the rotational constants in the vibrationally excited states (Table III) and those in the ground state satisfy the well-known relation

$$B_v - B_0 = -v\alpha_i^B, \quad (1)$$

with analogous equations for the A and C rotational constants. The experimental vibration-rotation interaction constants  $\alpha_i^{A,B,C}$  for a particular vibrational state were obtained from linear fits to the differences ( $A_v - A_0$ ,  $B_v - B_0$ , and  $C_v - C_0$ ) as a function of the vibrational quantum number  $v$  (Fig. 3), yielding the values  $\alpha^A = -44.1$  MHz,  $\alpha^B = 10.2$  MHz, and  $\alpha^C = 10.4$  MHz, with estimated  $1\sigma$  uncertainties of 0.1 MHz. A comparison of these experimental vibration-rotation interaction constants with those obtained theoretically (Table IV) shows that the observed vibrational

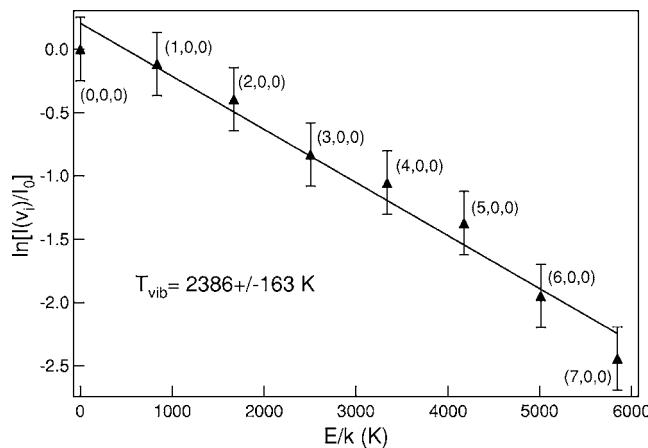


FIG. 4. Vibrational temperature diagram for the  $\nu_1$  mode of S<sub>3</sub> in our supersonic beam. Relative intensities in the vibrational levels  $[I(v_i)/I_0]$  are from measurements of the  $1_{1,1}-2_{0,2}$  transition observed here in a discharge through sulfur vapor with the FTM spectrometer.

satellites almost certainly arise from the symmetric stretching mode  $\nu_1$  calculated to lie at  $580\text{ cm}^{-1}$  (Table V). The close agreement of the theoretical interaction constants with the experimental values (to 7% or better) confirms that the theoretical potential surface provides a fairly good description of the vibrational dependence of the rotational constants (see Sec. V).

The vibrational temperature of the  $\nu_1$  mode of S<sub>3</sub> was determined from a fit of the relative intensities of the  $1_{1,1}-2_{0,2}$  line in different levels with respect to the ground state on the assumption of a Boltzmann distribution. From these measurements an effective vibrational temperature of  $2400 \pm 160$  K was found (Fig. 4). Detection of the  $\nu_1$  mode but not  $\nu_2$  near  $255\text{ cm}^{-1}$  is consistent with previous studies of vibrational excitation of polyatomic molecules in our discharge nozzle which showed that vibrational levels whose frequencies are greater than  $kT$  ( $\sim 450\text{ cm}^{-1}$ ) are little cooled in the supersonic expansion.<sup>29</sup> The  $\nu_3$  mode of S<sub>3</sub> near  $670\text{ cm}^{-1}$  (Table V) was not investigated here.

## B. Tetrasulfur, S<sub>4</sub>

Eighty-four lines of all sulfur-32 S<sub>4</sub> with  $J \leq 12$  and  $K_a \leq 6$  have been measured in the centimeter-wave band from 6 to 39 GHz [Table IV of SM (Ref. 28)]. Because the A rotational constant of S<sub>4</sub> is five times smaller than that of S<sub>3</sub> (4.7 GHz versus 24 GHz) and single isotopic substitution

TABLE V. Vibrational frequencies of S<sub>3</sub> (in  $\text{cm}^{-1}$ ).

Vib. Mode	Symmetry	Frequency <sup>a</sup>		Experiment	
		Harmonic	Fundamental <sup>b</sup>	Gas <sup>c</sup>	Matrix <sup>d</sup>
$\nu_1$	$a_1$	585	580	581	581–585
$\nu_2$	$a_1$	256	255	281	...
$\nu_3$	$b_1$	681	671	...	676–680

<sup>a</sup>Calculated at the CCSD(T)/cc-pVTZ level of theory (see Sec. III).

<sup>b</sup>Includes anharmonic contributions.

<sup>c</sup>From gas-phase Raman spectroscopy (Ref. 30).

<sup>d</sup>From infrared matrix isolation spectroscopy (Ref. 17).

TABLE VI. Spectroscopic constants of tetrasulfur,  $S_4$  (in MHz).

Constant <sup>a</sup>	$^{32}S_4$		$^{34}SS_3$	$S^{34}SS_2$
	Measured <sup>b</sup>	Theoretical <sup>c</sup>	Measured <sup>d</sup>	Measured <sup>e</sup>
<i>A</i>	4655.333 65(6)	4525.	4587.4305(2)	4586.5206(2)
<i>B</i>	2221.538 96(5)	2213.	2176.3209(2)	2198.8733(2)
<i>C</i>	1502.378 80(3)	1485.	1474.59812(7)	1484.83185(7)
$10^3D_J$	0.8742(2)	0.88	0.862(1)	0.833(1)
$10^3D_{JK}$	-1.882(1)	-1.9	-1.894(6)	-1.751(6)
$10^3D_K$	3.047(2)	2.8	3.00(1)	2.90(1)
$10^3d_1$	-0.3413(1)	-0.35	-0.3357(9)	-0.3263(10)
$10^3d_2$	-0.035 34(4)	-0.036	-0.0345(3)	-0.0343(4)
$10^9H_J$	-1.13(4)			
$10^9H_{JK}$	9.3(4)			
$10^6H_{KJ}$	-0.027(1)			
$10^6H_K$	0.028(2)			
$10^9h_1$	-0.48(2)			
$10^3\Delta E^f$	14.1(2)			
$\Delta^g$	0.336		0.340	0.338

<sup>a</sup>The  $1\sigma$  uncertainties (in parentheses) are in the units of the last significant digits.

<sup>b</sup>Derived from a least-squares fit of Watson's *S*-reduced Hamiltonian to the data in Table IV of SM (Ref. 28).

<sup>c</sup>Calculated at the CCSD(T)/cc-pVTZ level of theory.

<sup>d</sup>Derived from a least-squares fit of Watson's *S*-reduced Hamiltonian to the data in Table V of SM (Ref. 28).

<sup>e</sup>Derived from a least-squares fit of Watson's *S*-reduced Hamiltonian to the data in Table VI of SM (Ref. 28).

<sup>f</sup>Energy term that accounts for interchange tunneling.

<sup>g</sup>Inertial defect,  $\Delta=I_c-I_b-I_a$ , in units of amu  $\text{\AA}^2$ .

breaks the  $C_{2v}$  symmetry, a fairly large number of lines of the singly substituted sulfur-34 species were also observed: 69 for  $^{34}SS_3$  [Table V of SM (Ref. 28)] and 70 for  $S^{34}SS_2$  [Table VI of SM (Ref. 28)]. When the transition frequencies of the rare isotopic species were analyzed, again with Watson's *S*-reduced Hamiltonian, including the full set of quartic centrifugal distortion constants, the rms of the fits was comparable to the measurement uncertainties (Table VI). The same fitting procedure was less successful when applied to the parent species, where a rms residual of about 14 kHz was found.<sup>3</sup> Further analysis showed that the residuals exhibited a bimodal distribution: the  $K_aK_c=ee\rightarrow oo$  transitions were systematically displaced from the  $oo\rightarrow ee$  transitions by almost 15 kHz (see Fig. 5)—much more than the measurement uncertainty of a few kilohertz. Inclusion of a term in the Hamiltonian to account for interchange tunneling ( $\Delta E=O^+-O^-$ ),

however, reduced the rms error of the fit to about 1 kHz (Fig. 5), comparable to that found for the singly substituted sulfur-34 species, and yielding a precise value of the interchange tunneling frequency:  $14.1\pm 0.2$  kHz. A similar effect is observed in the rotational spectrum of  $\text{Ar}-\text{SO}_2$ , where the tunneling frequency is more than  $10^4$  times larger.<sup>31</sup> The tunneling frequency observed here for  $S_4$  is roughly seven orders of magnitude smaller than the ammonia inversion frequency, and is near the limit of what is now currently resolvable with our spectrometer. No smaller tunneling splitting to our knowledge has yet been observed.

The origin of the interchange tunneling in  $S_4$  is apparent from quantum chemical calculations, which show that there are two equivalent accessible minima on the potential-energy surface that differ only in the identity of the *inner* and *outer* sulfur atoms (see Fig. 3 in Ref. 3). Between these two trap-

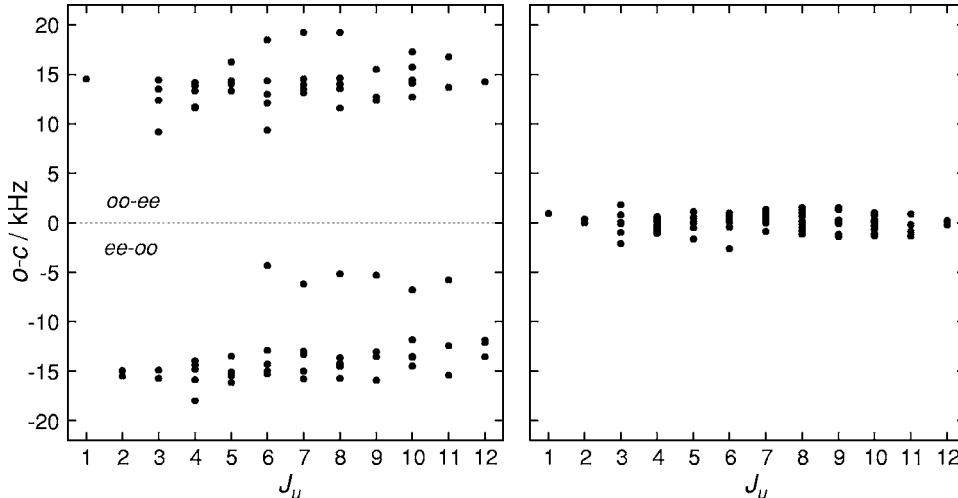


FIG. 5. Observed minus calculated frequencies ( $O-C$ ) vs  $J_{\text{upper}}$  from a least-squares fit to the centimeter-wave measurements of  $S_4$ . Left: The residuals fall into two distinct groups depending on whether the transition corresponds to  $K_aK_c=oo\rightarrow ee$  or  $ee\rightarrow oo$ . Right: The rms of the fit (1 kHz) is comparable to the measurement uncertainties when a term that accounts for interchange tunneling ( $\Delta E_{O^+-O^-}$ ) is included in the Hamiltonian.

ezoidal C<sub>2v</sub> structures is a rectangular D<sub>2h</sub> transition state located about 500 cm<sup>-1</sup> above the two equivalent minima. Using three-dimensional models of the potential-energy surface governing the tunneling motion that differ only slightly in the details of the parametrization, ground-state tunneling frequencies of about 1–100 kHz were obtained that are comparable to the observed frequency.

Following the FTM study of S<sub>4</sub>, measurements were extended in the millimeter-wave band up to 271 GHz. Initially, R-branch transitions with low K<sub>a</sub> in the range between 145 and 188 GHz were observed within a few megahertz of predictions from the centimeter-wave data. Assignment of the millimeter-wave spectrum of S<sub>4</sub> was more difficult than that of S<sub>3</sub> because lines were weaker, and there was occasional confusion with background lines, possibly the result of vibrationally excited S<sub>3</sub> and S<sub>4</sub>. In all, 147 lines of S<sub>4</sub> were measured [Table IV of SM (Ref. 28)], yielding the spectroscopic constants in Table VI. The full set of fourth-order and five sixth-order centrifugal distortion constants were needed to adequately reproduce the measurements.

## V. STRUCTURES OF S<sub>3</sub> AND S<sub>4</sub>

Experimental (r<sub>0</sub>) structures of S<sub>3</sub> and S<sub>4</sub> obtained by least-squares adjustment of bond lengths and angles to the observed rotational constants (A<sub>0</sub>, B<sub>0</sub>, and C<sub>0</sub>) for three isotopic species were reported in preliminary accounts of this work.<sup>2,3</sup> In the present investigation, the structural determination is extended to the equilibrium (r<sub>e</sub>) structure that corresponds to the minimum on the potential-energy surface. To this end, the equilibrium rotational constants B<sub>e</sub> are estimated from the expression

$$B_e = B_0 + \frac{1}{2} \sum_i \alpha_i^B, \quad (2)$$

where B<sub>0</sub> is the measured rotational constant in the ground vibrational state; similar equations hold for A<sub>e</sub> and C<sub>e</sub>, and the sum over vibration-rotation interaction constants is taken over each of the normal modes of vibration. The experimental determination of all the α<sub>i</sub><sup>A,B,C</sup> is increasingly difficult for large molecules, but these terms can be calculated theoretically to adequate accuracy. While the vibration-rotation interaction constants (α<sub>i</sub>) are often complicated by Coriolis and Fermi resonances, the difference between the equilibrium rotational constants and those of the ground vibrational state is determined by the sum (Σ<sub>i</sub>α<sub>i</sub>/2) which is not affected by these resonance interactions within second-order perturbation theory (see Ref. 32 and references therein). This procedure of combining experimental rotational constants with theoretical vibrational corrections is generally an efficient way to determine accurate equilibrium geometries of polyatomic molecules (for a review see Ref. 33). The equilibrium structures are derived by the same least-squares fitting procedure used for the r<sub>0</sub> geometries, except that the equilibrium constants obtained from Eq. (2) are used instead of the experimental constants. The vibration-rotation interaction constants and the sum (Σ<sub>i</sub>α<sub>i</sub>/2) for S<sub>3</sub> and S<sub>4</sub> calculated at the CCSD(T)/cc-pVTZ level of theory are given in Tables IV and VII.

TABLE VII. Theoretical vibration-rotation interaction constants of S<sub>4</sub> (in MHz) calculated at the CCSD(T)/cc-pVTZ level of theory.

Vib. Mode and Σ <sub>i</sub> α <sub>i</sub> /2	SSSS		
	α <sub>i</sub> <sup>A</sup>	α <sub>i</sub> <sup>B</sup>	α <sub>i</sub> <sup>C</sup>
ν <sub>1</sub>	+8.25	+5.50	+3.21
ν <sub>2</sub>	+18.27	-7.44	-38.05
ν <sub>3</sub>	-8.88	+8.73	+3.12
ν <sub>4</sub>	+0.01	+5.79	+2.35
ν <sub>5</sub>	+11.79	+3.07	+2.66
ν <sub>6</sub>	-6.28	+2.25	+40.43
Σ <sub>i</sub> α <sub>i</sub> /2	+11.58	+8.96	+6.86
	<sup>34</sup> SSSS		
	α <sub>i</sub> <sup>A</sup>	α <sub>i</sub> <sup>B</sup>	α <sub>i</sub> <sup>C</sup>
	+8.15	+5.43	+3.18
ν <sub>1</sub>	+16.94	-7.33	-18.53
ν <sub>2</sub>	-8.67	+8.19	+2.98
ν <sub>3</sub>	0.00	+5.90	+2.34
ν <sub>4</sub>	+11.46	+3.03	+2.59
ν <sub>5</sub>	-4.80	+1.60	+20.57
ν <sub>6</sub>	+11.53	+8.41	+6.57
	S <sup>34</sup> SSS		
	α <sub>i</sub> <sup>A</sup>	α <sub>i</sub> <sup>B</sup>	α <sub>i</sub> <sup>C</sup>
	+8.05	+5.26	+3.08
ν <sub>1</sub>	+6.99	-2.76	-13.24
ν <sub>2</sub>	-8.67	+8.87	+3.11
ν <sub>3</sub>	+0.02	+5.41	+2.25
ν <sub>4</sub>	+11.55	+3.04	+2.62
ν <sub>5</sub>	+4.30	-1.64	+15.84
Σ <sub>i</sub> α <sub>i</sub> /2	+11.12	+9.09	+6.83

## A. S<sub>3</sub>

For S<sub>3</sub>, the bond length and angle are nearly the same for the experimental (r<sub>0</sub>) and the empirical equilibrium (r<sub>e</sub><sup>emp</sup>) structures (Table VIII), probably owing to little vibrational anharmonicity (see Table V for a comparison of the calculated fundamental and harmonic vibrational frequencies). Although the structure remains essentially unchanged when the rotational constants are corrected for vibrational effects, the inertial defects (Δ=I<sub>c</sub>-I<sub>a</sub>-I<sub>b</sub>) derived from the measured ro-

TABLE VIII. Structures of S<sub>3</sub> and S<sub>4</sub>.

Structure <sup>a</sup>	S <sub>3</sub>		S <sub>4</sub>		
	r(S=S) (Å)	α (deg)	r(S=S) outer (Å)	r(S=S) inner (Å)	α (deg)
r <sub>0</sub>	1.917(1)	117.36(6)	1.899(7)	2.173(32)	103.9(3)
r <sub>e</sub> <sup>calc</sup>	1.938	117.0	1.920	2.177	103.6
Recommended structure:					
r <sub>e</sub> <sup>emp</sup>	1.914(2)	117.33(5)	1.898(5)	2.155(10)	104.2(2)

<sup>a</sup>Experimental (r<sub>0</sub>) structures are from Refs. 2 and 3. Theoretical equilibrium (r<sub>e</sub><sup>calc</sup>) structures were calculated at the CCSD(T)/cc-pVTZ level of theory. Empirical equilibrium (r<sub>e</sub><sup>emp</sup>) structures were determined by the method described in Sec. V.

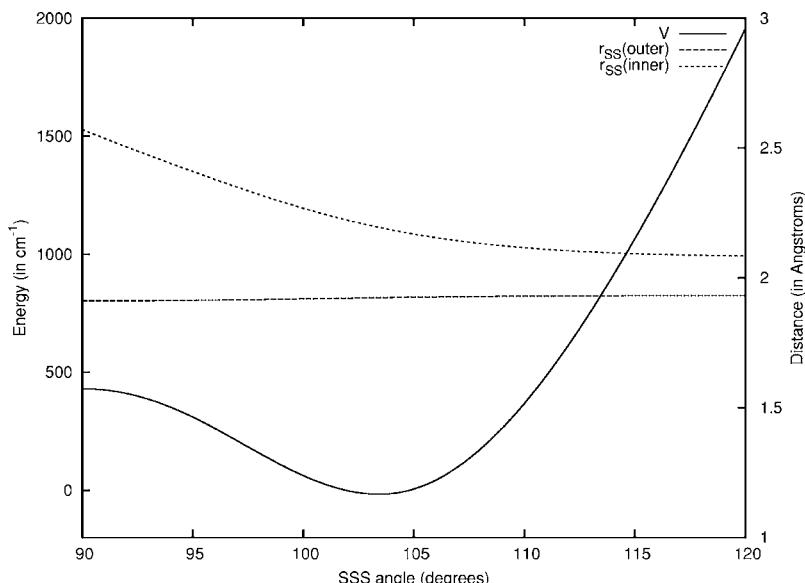


FIG. 6. Potential energy and geometrical parameters associated with the interconversion of  ${}^4\text{SSSS}$  and  $\text{S}{}^4\text{SSS}$  via a  $D_{2h}$  transition state. The ordinate is the SSS bond angle, along which the other two geometrical parameters (the SS bond lengths) are optimized at each point. Calculations were performed in the frozen-core approximation at the CCSD(T)/cc-pVTZ level of theory.

tational constants for the three isotopic species, all  $\sim 0.24$  amu  $\text{\AA}^2$ , are roughly two orders of magnitude smaller for the empirical equilibrium structure ( $-0.005$  to  $-0.001$  amu  $\text{\AA}^2$ ). This large reduction in  $\Delta$ , which vanishes for a rigid planar molecule, indicates that the calculated quadratic and cubic force constants for  $\text{S}_3$  are fairly accurate.

The bond length in the theoretical equilibrium ( $r_e^{\text{calc}}$ ) structure calculated at the CCSD(T) level with the cc-pVTZ basis set (1.938  $\text{\AA}$ ) is about 0.02  $\text{\AA}$  longer than the experimental value (Table VIII). Experience has shown that basis set expansion tends to reduce internuclear distances while correlation tends to increase them. In the example of  $\text{S}_3$ , the CCSD(T) method provides an adequate treatment of correlation; the major shortcoming of the calculation is the cc-pVTZ basis set. One would expect that the equilibrium bond length calculated with the cc-pVTZ basis set would be slightly on the high side, and that is what is observed for  $\text{S}_3$ . In a similar calculation but with a more complete basis set (cc-pCVQZ), the calculated bond length (1.913  $\text{\AA}$ ) is in nearly perfect agreement with that of the  $r_e^{\text{emp}}$  structure,<sup>34</sup> confirming that the calculation of the theoretical ( $r_e^{\text{calc}}$ ) structure is limited by the cc-pVTZ basis set.

## B. $\text{S}_4$

The empirical equilibrium geometry ( $r_e^{\text{emp}}$ ) of  $\text{S}_4$  was obtained as for  $\text{S}_3$ . The central (inner) SS distance in the  $r_0$  and  $r_e^{\text{emp}}$  structures differ by 0.02  $\text{\AA}$ , while the outer distance and the SSS bond angle are the same to within the estimated uncertainties (Table VIII). Discrepancies between the  $r_e^{\text{emp}}$  and  $r_e^{\text{calc}}$  structures for  $\text{S}_4$  are qualitatively similar to those noted for  $\text{S}_3$  (see Table VIII). On the assumption that the  $r_0$  structure approximates the vibrationally averaged structure in the ground vibrational state, this difference in the inner distance in the  $r_0$  and  $r_e^{\text{emp}}$  structures is attributed to the coupling of the stretching motion of the inner S atoms to the tunneling coordinate (here taken as the SSS bond angle  $\theta$ ), where  $\theta$  is 90° at the transition state and  $\theta_e$  is roughly 105° at the equilibrium geometry. The shape of the potential for the bending motion (Fig. 6) implies that the ground state wave

function is greater at angles less than  $\theta_e$  where the S-S distance is longer, than at  $\theta > \theta_e$  where the distance is shorter than  $r_e$ .

The inertial defect derived from the empirical equilibrium structure of  $\text{S}_4$  ( $\leq -0.01$  amu  $\text{\AA}^2$ ) is more than an order of magnitude smaller than that derived from the measured rotational constants ( $\sim 0.34$  amu  $\text{\AA}^2$ , Table VI). Although the magnitude of the residual in  $\Delta$  for  $\text{S}_4$  is roughly an order of magnitude larger than that for  $\text{S}_3$ , it is comparable to that found for the empirical equilibrium structures of other planar molecules of similar size, such as  $\text{SiC}_3$ .<sup>35</sup>

Owing to the difficulties in properly treating electron correlation, previous theoretical studies predicted geometrical parameters of  $\text{S}_4$  that are in poor agreement with those derived here. Qualitatively,  $\text{S}_4$  can be thought of as two  $\text{S}_2$  triplet states in which one of the unpaired electrons of an  $\text{S}_2$  fragment is used to make a bond to the other fragment,<sup>36</sup> resulting in a substantial amount of biradical character. Inspection of the coupled-cluster wave function for  $\text{S}_4$  reveals relatively large double-excitation amplitudes, which is a hallmark of biradical character in species such as ozone and the benzenes. Thus the rather powerful treatment of correlation afforded by the CCSD(T) model provides a good description of the electronic structure of this molecule.

## VI. DISCUSSION

Further study of interchange tunneling in  $\text{S}_4$  is worthwhile, including a more sophisticated theoretical treatment of the dynamical motion, and observations of rotational spectra in vibrationally excited states, particularly in  $\nu_2$  and  $\nu_3$  (Fig. 7) which play principal roles in the tunneling coordinate. While the vibrational state dependence of the tunneling splitting in the normal isotopic species of  $\text{S}_4$  is of interest, this is even more so for the isotopic forms. The present laboratory study has succeeded in determining the ground-state rotational constants for  ${}^{34}\text{SS}_3$  and  $\text{S}{}^{34}\text{SSS}$ , but the distinction between these two structures will diminish as the level of vibrational excitation increases, because the two are linked by the low-lying transition state of  $D_{2h}$  symmetry. Therefore

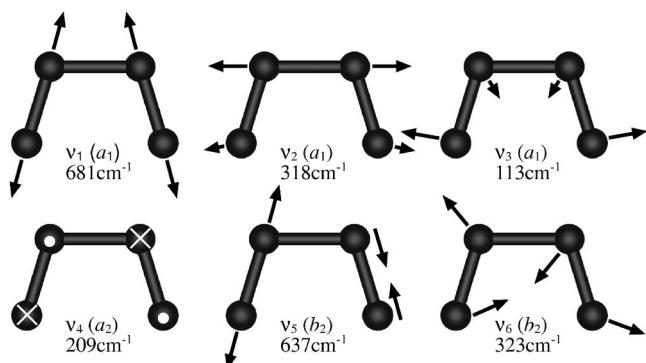


FIG. 7. Normal vibrational modes of  $S_4$ . Fundamental vibrational frequencies calculated at the CCSD(T)/cc-pVTZ level of theory. Length and direction of displacement vectors are approximate.

an investigation of the rotational spectra of isotopically substituted species of  $S_4$  in vibrationally excited states might provide additional insights into the dynamics of this curious molecule. Interchange tunneling could be observed directly in the fully substituted sulfur-33 species because this species obeys Fermi-Dirac statistics and all rotational levels exist; with the present sensitivity this would require an isotopically enriched sample.

The substantial abundances of  $S_3$  and  $S_4$  in our molecular beam suggest that less stable isomers (such as the  $C_s$  isomer of  $S_4$  in Fig. 8), and even larger sulfur clusters with more than four atoms may be within reach. Very little is known about the next member in the series  $S_5$ . Its most stable isomer is predicted to be a half-boat structure with  $C_s$  symmetry that is moderately polar [0.39 D (Ref. 37)], but fairly floppy, with a low barrier to pseudorotation. For  $S_6$ , the most stable form is predicted to be a nonpolar chair configuration similar to that observed in the solid phase. Less stable isomers of  $S_6$ , and even larger pure sulfur clusters may also be present in our beam.

In addition to pure sulfur clusters, exotic sulfur oxides are of considerable chemical interest. A number of small oxides of sulfur such as  $S_3O$  and higher sulfur monoxides  $S_nO$  have not been studied by high-resolution spectroscopy, but may be detectable by the same technique used here to characterize the pure sulfur clusters. Both the planar *cis*- and *trans*-isomers of  $S_3O$  are polar, unlike isovalent  $S_4$  and  $S_2O_2$ , for which only the *cis*-isomers are polar. Owing to the high reactivity of heated sulfur with oxygen and to the intense rotational lines of  $S_2O$  observed in our discharge sources,

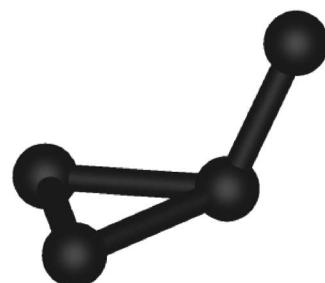


FIG. 8. Molecular structure of a polar  $C_s$  isomer of  $S_4$  predicted to lie about 14 kcal/mol higher in energy than the  $C_{2v}$  isomer (Ref. 6) shown in Fig. 1.

several of these sulfur oxides are good candidates for detection.

Both  $S_3$  and  $S_4$  are of astronomical interest because sulfur is one of the most cosmically abundant second-row elements, and these are the smallest polar sulfur clusters. Sulfur chemistry in space is not very well understood, primarily because the major reservoir of sulfur has not been identified. Thus detection of  $S_3$  and  $S_4$  might provide important quantitative information on the abundance and form of elemental sulfur. A detailed account of the prospects for observing  $S_3$  and  $S_4$  in Galactic molecular sources has been discussed elsewhere.<sup>38</sup> In the solar system, both  $S_3$  and  $S_4$  may be detectable in the Jovian moon Io and in comets, because S and  $S_2$  have been observed at UV wavelengths in these objects. The abundances of many of the 14 sulfur-containing molecules detected in Galactic molecular sources are high in hot dense cores, so  $S_3$  and  $S_4$  are plausible candidates for detection in such objects. The laboratory measurements in the centimeter- and millimeter-wave bands discussed here provide the basic spectroscopic information needed for deep radio astronomical searches of these two fundamental molecules.

## ACKNOWLEDGMENTS

We thank W. Klemperer, J. K. G. Watson, H. M. Pickett, H. S. P. Müller, J. Gauss, and J. Hahn for helpful discussions; F. J. Lovas for providing rotational data on the  $H_2S$  dimer prior to publication; and E. S. Palmer for laboratory assistance. The work in Cambridge was supported by NASA Grant No. NAG5-9379 and NSF Grant No. AST-9820722. One of the authors (J.F.S.) acknowledges the support of the Robert A. Welch Foundation and the NSF. Another author

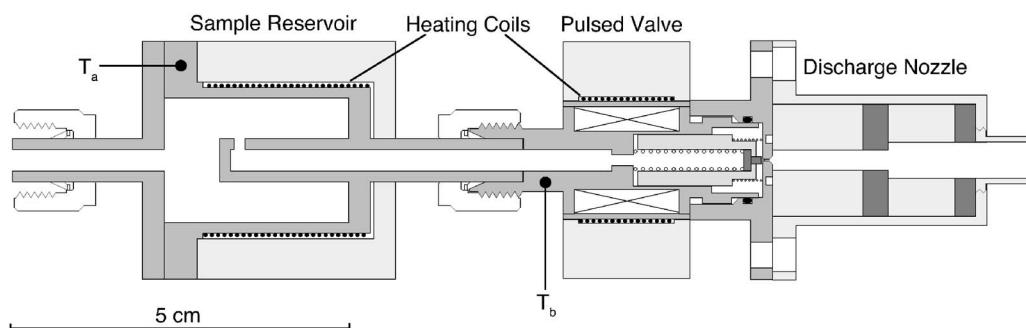


FIG. 9. Cross section of the heated nozzle.

(S.T.) is grateful to the Alexander von Humboldt Foundation for a Feodor Lynen research fellowship.

## APPENDIX: THE HEATED NOZZLE

A heated nozzle for low volatile compounds was constructed during the course of the present investigation (Fig. 9). It consists of a sample reservoir which can be resistively heated, a pulsed valve (General valve series 9), and a discharge nozzle similar to that described in Ref. 8. Heating coils (80/20 Nichrome) are wrapped around the sample reservoir and pulsed valve. The coils are enclosed in jackets of Torlon 4203 shown as light shaded areas in Fig. 9. Torlon is a strong machinable thermoplastic that is usable to 260 °C and provides excellent electrical insulation. The sample reservoir and pulsed valve can be heated separately, but the maximum temperature of the entire assembly (200 °C) is limited by the valve.

- <sup>1</sup>B. Eckert and R. Steudel, *Top. Curr. Chem.* **231**, 31 (2003), and references therein.
- <sup>2</sup>M. C. McCarthy, S. Thorwirth, C. A. Gottlieb, and P. Thaddeus, *J. Am. Chem. Soc.* **126**, 4096 (2004).
- <sup>3</sup>M. C. McCarthy, S. Thorwirth, C. A. Gottlieb, and P. Thaddeus, *J. Chem. Phys.* **121**, 632 (2004).
- <sup>4</sup>M. W. Wong, *Top. Curr. Chem.* **231**, 1 (2003).
- <sup>5</sup>G. E. Quelch, H. F. Schaefer III, and C. J. Mardsen, *J. Am. Chem. Soc.* **112**, 8719 (1990).
- <sup>6</sup>M. W. Wong and R. Steudel, *Chem. Phys. Lett.* **379**, 162 (2003), and references therein.
- <sup>7</sup>M. C. McCarthy, M. J. Travers, A. Kovács, C. A. Gottlieb, and P. Thaddeus, *Astrophys. J., Suppl. Ser.* **113**, 105 (1997).
- <sup>8</sup>M. C. McCarthy, W. Chen, M. J. Travers, and P. Thaddeus, *Astrophys. J., Suppl. Ser.* **129**, 611 (2000).
- <sup>9</sup>S. Yamamoto and S. Saito, *Can. J. Phys.* **72**, 954 (1994); G. Winnewisser, M. Winnewisser, and W. J. Gordy, *J. Chem. Phys.* **49**, 3465 (1968); M. Behnke, dissertation, Universität zu Köln (2001).
- <sup>10</sup>R. L. Brown, *J. Chem. Phys.* **44**, 2827 (1965); R. B. Langford and G. A. Oldershaw, *J. Chem. Soc., Faraday Trans.* **68**, 1550 (1972).
- <sup>11</sup>B. Meyer, *Chem. Rev. (Washington, D.C.)* **76**, 367 (1976).
- <sup>12</sup>R. B. Langford and G. A. Oldershaw, *J. Chem. Soc., Faraday Trans.* **69**, 1389 (1973).
- <sup>13</sup>The reaction  $S(^1D) + OCS \rightarrow S_2 + CO$  yields a significant fraction of  $S_2$  in the first electronically excited state [ $a^1\Delta$ ; A. J. Hynes, R. C. Richter, A. R. Rosendahl, and C. D. Clark, *Chem. Phys. Lett.* **295**, 25 (1998)], where it cannot readily react to form  $S_3$  (with  $^1A_1$  symmetry), owing to spin and orbital angular momentum restrictions.
- <sup>14</sup>C. A. Gottlieb, P. C. Myers, and P. Thaddeus, *Astrophys. J.* **588**, 655 (2003).
- <sup>15</sup>D. B. Nash, *Icarus* **72**, 1 (1987).
- <sup>16</sup>P. Hassanzadeh and L. Andrews, *J. Phys. Chem.* **96**, 6579 (1992).
- <sup>17</sup>G. D. Brabson, Z. Mielke, and L. Andrews, *J. Phys. Chem.* **95**, 79 (1991).
- <sup>18</sup>Cologne Database for Molecular Spectroscopy, <http://www.cdms.de>; H. S. P. Müller, S. Thorwirth, D. A. Roth, and G. Winnewisser, *Astron. Astrophys.* **370**, L49 (2001).
- <sup>19</sup>R. L. Cook, G. Winnewisser, and D. C. Lindsey, *J. Mol. Spectrosc.* **46**, 276 (1973).
- <sup>20</sup>For reviews of coupled-cluster theory, see: T. J. Lee and G. E. Scuseria, in *Quantum Mechanical Electronic Structure Calculations with Chemical Accuracy*, edited by S. R. Langhoff (Kluwer, Dordrecht, 1995); R. J. Bartlett, in *Modern Electronic Structure Theory*, edited by D. R. Yarkony (World Scientific, Singapore, 1995), Pt. II; J. Gauss, in *Encyclopedia of Computational Chemistry*, edited by P. v. R. Schleyer (Wiley, New York, 1998).
- <sup>21</sup>T. H. Dunning, *J. Chem. Phys.* **90**, 1007 (1989).
- <sup>22</sup>K. Raghavachari, G. W. Trucks, J. A. Pople, and M. Head-Gordon, *Chem. Phys. Lett.* **157**, 479 (1989).
- <sup>23</sup>G. E. Scuseria, *J. Chem. Phys.* **94**, 442 (1991).
- <sup>24</sup>J. Gauss and J. F. Stanton, *Chem. Phys. Lett.* **276**, 70 (1997).
- <sup>25</sup>J. F. Stanton, C. L. Lopreore, and J. Gauss, *J. Chem. Phys.* **108**, 7190 (1998).
- <sup>26</sup>I. M. Mills, in *Modern Spectroscopy: Modern Research*, edited by K. N. Rao and C. W. Matthews (Academic, New York, 1972), pp. 115–140.
- <sup>27</sup>J. F. Stanton, J. Gauss, J. D. Watts, W. J. Lauderdale, and R. J. Bartlett, *Int. J. Quantum Chem.* **44**, 879 (1992).
- <sup>28</sup>See EPAPS Document No. E-JCPA6-123-009528 for complete lists of experimental transition frequencies. This document can be reached via a direct link in the online article's HTML reference section or via the EPAPS homepage (<http://www.aip.org/pubservs/epaps.html>).
- <sup>29</sup>M. E. Sanz, M. C. McCarthy, and P. Thaddeus, *J. Chem. Phys.* **122**, 194319 (2005).
- <sup>30</sup>E. Piquenard, O. El Jaroudi, and J. Corset, *J. Raman Spectrosc.* **24**, 11 (1993).
- <sup>31</sup>L. H. Coudert, K. Matsumura, and F. J. Lovas, *J. Mol. Spectrosc.* **147**, 46 (1991).
- <sup>32</sup>F. Pawłowski, P. Jørgensen, J. Olsen, F. Hegelund, T. Helgaker, J. Gauss, K. L. Bak, and J. F. Stanton, *J. Chem. Phys.* **116**, 6482 (2002).
- <sup>33</sup>J. F. Stanton and J. Gauss, *Int. Rev. Phys. Chem.* **19**, 61 (2000).
- <sup>34</sup>The structural optimization at the CCSD(T)/cc-pCVQZ level of theory yields a bond length of 1.913 Å and an angle of 117.54° (J. Gauss, private communication).
- <sup>35</sup>J. F. Stanton, J. Gauss, and O. Christiansen, *J. Chem. Phys.* **114**, 2993 (2001).
- <sup>36</sup>The binding energy of  $S_4$  (including zero-point vibrational corrections) was calculated at the CCSD(T) level with the cc-pVTZ basis set. The calculation gives a  $S_2$  dimerization enthalpy of -15.3 kcal/mol, implying that the exact value lies within the range of 10–20 kcal/mol. Thus it is not appropriate to view this molecule as a van der Waals complex in spite of the low-energy interchange tunneling mechanism.
- <sup>37</sup>S. Millefiori and A. Alparone, *J. Phys. Chem. A* **105**, 9489 (2001).
- <sup>38</sup>C. A. Gottlieb, S. Thorwirth, M. C. McCarthy, and P. Thaddeus, *Astrophys. J.* **619**, 939 (2005).