

High Accuracy Spectral Lines For Radiation Transport In Stellar Atmosphere

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10th HITRAN Conference, 2008

- Introduction and general interest in hydrocarbons.
- High accuracy spectroscopic calculations.
- Development of Potential Energy and Dipole Moment surfaces.
- Results: Ethyl Cation and Methane.
- Conclusion.

Radiation transport needs..

- PHOENIX: Synthetic spectra for young brown dwarfs and substellar objects.

Selected molecules considered in the EOS

NH	C ₂	CN	CO	MgH	CaH	SiH	TiO	H ₂ O	H ₂
N ₂	NO	CO ₂	O ₂	ZrO	VO	MgS	SiO	AlH	HCl
HF	HS	TiH	AlO	BO	CrO	LaO	MgO	ScO	YO
SiF	NaCl	CaOH	HCN	C ₂ H ₂	CH ₄	CH ₂	C ₂ H	HCO	NH ₂
LiOH	C ₂ O	AlOF	NaOH	MgOH	AlO ₂	Al ₂ O	AlOH	SiH ₂	SiO ₂
H ₂ S	OCS	KOH	TiO ₂	TiOCl	VO ₂	FeF ₂	YO ₂	ZrO ₂	BaOH
LaO ₂	C ₂ H ₄	C ₃	SiC ₂	CH ₃	C ₃ H	NH ₃	C ₂ N ₂	C ₂ N	CaF ₂
AlOCl	Si ₂ C	CS ₂	CaCl ₂	AlF	CaF	Si ₂	SiS	CS	AlCl
KCl	CaCl	TiS	TiCl	SiN	AlS	Al ₂	FeO	SiC	TiF ₂
FeH	LiCl	NS	NaH	SO	S ₂	AlBO ₂	AlClF	AlCl ₂	AlF ₂
AlOF ₂	AlO ₂ H	Al ₂ O ₂	BeBO ₂	ObF	HBO	HBO ₂	HBS	BH ₂	BO ₂ H ₂
BH ₃	H ₃ BO ₃	KBO ₂	LiBO ₂	NaBO ₂	BO ₂	BaCl ₂	BaF ₂	BaO ₂ H ₂	BaClF
BeCl ₂	BeF ₂	BeOH	BeH ₂	BeH ₂ O ₂	Be ₂ O	Be ₃ O ₃	CiCN	CHCl	CHF
CHP	CH ₃ Cl	KCN	NaCN	BeC ₂	C ₃ HCl	C ₂ HF	(NaCN) ₂	C ₄	C ₅
CaO ₂ H ₂	MgClF	SiH ₃ Cl	FeCl ₂	K ₂ Cl ₂	MgCl ₂	Na ₂ Cl ₂	TiOCl ₂	SrCl ₂	TiCl ₂
ZrCl ₂	TiCl ₃	ZrCl ₃	ZrCl ₄	CrO ₂	SiH ₃ F	OTiF	SiH ₂ F ₂	MgF ₂	SrF ₂
ZrF ₂	TiF ₃	ZrF ₄	FeO ₂ H ₂	SrOH	(KOH) ₂	(LiOH) ₂	MgO ₂ H ₂	(NaOH) ₂	SrO ₂ H ₂
PH ₂	PH ₃	SiH ₄	Si ₂ N	PO ₂	SO ₂	P ₄	Si ₃	NO ₂	NO ₃
C ₃ N	C ₂ H ₃	C ₄ H	HC ₃ N	C ₄ H ₂	CH ₃ CN	HC ₅ N	C ₆ H	C ₄ H ₄	C ₆ H ₂
HC ₇ N	C ₄ H ₄ S	C ₄ H ₄ O	C ₄ H ₆	C ₆ H ₄	HC ₉ N	C ₅ H ₃ N	C ₆ H ₅ O	C ₆ H ₆	C ₆ H ₆ O
HC ₁₁ N	OH ⁻	CH ⁻	C ₂ ⁻	OH	CH	CN ⁻	SiH ⁻	H ₂ ⁻	HS ⁻
CS ⁻	FeO ⁻	BO ⁻	AlCl ₂ ⁻	AlF ₂ ⁻	AlOF ₂ ⁻	AlOH ⁻	CO ₂ ⁻	NO ⁺	H ₂ ⁺
TiO ⁺	ZrO ⁺	AlOH ⁺	BaOH ⁺	HCO ⁺	CaOH ⁺	SrOH ⁺	H ₃ O ⁺	H ₃ ⁺	

- Hydrocarbons; Ethyl cation ($C_2H_5^+$) and $C_2H_3^+$ (1) are of interest in the field of Plasma chemistry, combustion and other fields.
- Recent observations on Titan measured ion density.
- Bridged global minimum and Y-Shape first order saddle point (local min for $C_2H_3^+$).
- Energy difference 7.14 kcal/mol ($C_2H_5^+$) and 3.414 kcal/mol ($C_2H_3^+$).
- Located accurately on PES (not explicitly added).
- We aim at performing high accuracy spectroscopic calculations using ab-initio based PES.

(1) A. R. Sharma, *et.al.* J. Chem. Phys., 2006; 125 (22) 224306

Watson Hamiltonian for a polyatomic molecule in normal coordinates:

$$\hat{H} = -\frac{\hbar^2}{2} \sum_{i=1}^N \frac{\partial^2}{\partial Q_i^2} + \frac{\hbar^2}{2} \sum_{\alpha=1}^3 \sum_{\beta=1}^3 (J_\alpha - \pi_\alpha) \mu_{\alpha\beta} (J_\beta - \pi_\beta) - \frac{\hbar^2}{8} \sum_{\alpha=1}^3 \mu_{\alpha\alpha} + V(Q_1, \dots, Q_N)$$

J_α and π_α components of the total and vibrational angular momentum operator

μ effective reciprocal inertia tensor

Q_i mass-weighted normal coordinate for mode i

N number of vibrational degrees of freedom

S. Carter, J. M. Bowman and N. Handy, Theor. Chem. Acc. (1998)100,191-198

Non-rotating ($J=0$)

Watson hamiltonian(in atomic units):

$$\hat{H} = \frac{1}{2} \sum_{\alpha\beta} \hat{\pi}_\alpha \mu_{\alpha\beta} \hat{\pi}_\beta - \frac{1}{8} \sum_{\alpha} \mu_{\alpha\alpha} - \frac{1}{2} \sum_i \frac{\partial^2}{\partial Q_i^2} + V(Q_1, \dots, Q_N)$$

the VSCF method approximates the vibrational wave function as a Hartree product of single-mode wavefunctions called *modals*.

$$\Psi^n(Q_1, \dots, Q_N) = \prod_i^N \phi_i^{n_i}(Q_i)$$

VSCF method is a variational procedure for obtaining the modals, and the optimized wave function of the form is obtained by minimizing the total energy with respect to all the modals subject to the constraint $\langle \phi_i | \phi_j \rangle = \delta_{ij}$;enforced by the Lagrange multipliers.

J. M. Bowman, J. Chem. Phys. 68, 608 (1978)

- VSCF eigenstates used for further accurate VCI calculations.☰



Rotational energy calculated as function of nuclear coordinates and added to nonrotating system.

$$\begin{aligned}H_{K_a, K_c}^J &= H^{J=0} + E_{K_a, K_c}^J(\mathbf{Q}) \\ &= H^{J=0} + [A\hat{J}_z^2 + B\hat{J}_y^2 + C\hat{J}_x^2]\end{aligned}$$

Adiabatic Rotation Approximation (ARA): excellent agreement with exact calculations for HO₂ (near prolate symmetric top); good agreement H₂O (strongly asymmetric top)

S. Carter and J. M. Bowman, J. Chem. Phys. 108, 4397 (1998); S. Carter, J. M. Bowman and N. Handy, Theor.

Chem. Acc. (1998)100,191-198

- Ro-vibrational calculations do require accurate PES description!!

Potential Energy Surface

- Electronic structure energy calculated at the highest level of theory.
- Nuclear configuration space sampled using MD and DMC.
- PES represented as many-body expansion.
- Translational and rotational invariance: PES is function of internuclear distances.
- Permutational invariance: polynomial basis invariant under group permutation operator.
- Invariant polynomial is unique product of primary and secondary generators.
- Computed using computational invariant theory (Magma)

- Fitting coefficients are solution of least square system of equations.
- Example: $(C_2H_5^+)$ 21 Primary inv. 1436 Secondary inv. (up to degree 7 of polynomial)
- Dimension : with symmetry 8717; without symmetry 1,184,040.
- Many successful application in Bowman Group.

Dipole Moment Surface

- Dipole moments computed using Molpro.
- DM is a vector quantity; cannot be expressed simply in terms of internuclear distances.
- Representation:

$$\mathbf{d} = \sum_i f_i(x) \mathbf{r}(i)$$

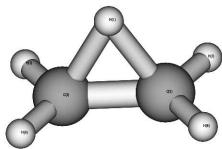
d: fitted dipole moment vector, *i*: nuclear indices, *x*: function of internuclear distance and **r**(*i*): vector position of the *i*-th nucleus.

- The functions f_i are polynomials
- Constrained so that **d** satisfies the required permutational invariance and translation
- Coefficients of the f_i determined by least squares fitting.

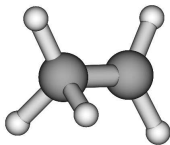
Properties of the PES

- Fit has a root mean square(rms) fitting error \approx milli-Hartree over all configurations (≈ 1.4 kcal/mol).
- For points 0.0 to 0.1 Hartree (0 to 62.7 kcal/mol) above the global minimum, rms error $<$ milli-Hartree.
- Normal mode frequencies: Excellent agreement.
- Used of anharmonic ro-vibrational calculations.
- PES routinely generated for various molecules.

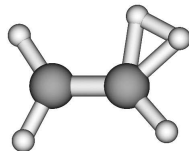
Global minimum and other stationary points on the PES .



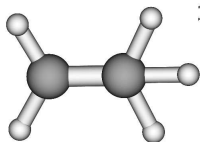
(a) 0.0



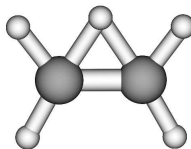
(b) 7.144



(c) 43.555



(d) 53.342



(e) 60.291

Energy relative to global minimum (kcal/mol).

Harmonic frequencies (cm⁻¹) for the nonclassical bridged structure of C₂H₅⁺

Mode Number	Symmetry	PES (1)	MP4(SDTQ) (2)	Quapp and Heidrich (3)	Trinquier (4)
1	B ₂	753.4	733.72	763	425
2	B ₁	817.5	827.60	865	902
3	A ₂	1094.9	1081.36	1113	1136
4	B ₁	1108.7	1104.49	1156	1223
5	A ₁	1132.7	1138.32	1177	1233
6	A ₂	1260.2	1251.80	1291	1336
7	B ₂	1276.3	1283.77	1347	1384
8	A ₁	1347.9	1351.07	1398	1428
9	B ₂	1450.5	1476.49	1527	1568
10	A ₁	1574.0	1569.44	1626	1696
11	A ₁	2163.8	2175.33	2273	2316
12	B ₂	3126.7	3136.51	3234	3283
13	A ₁	3136.2	3139.82	3237	3290
14	A ₂	3250.7	3243.94	3351	3397
15	B ₁	3257.7	3261.09	3366	3413

1. Normal mode frequencies calculated from fitted PES.

2. Frequencies calculated at the MP4(SDTQ) level of theory with aug-cc-pVTZ basis.

3. Frequencies calculated using a 6-31G** basis set at the MP2 level of theory. (2002)

4. Frequencies calculated using a DZP basis set at the SCF level of theory. (1992)

Harmonic frequencies (cm⁻¹) and zero point vibrational energy (kcal/mol) for nonclassical bridged structure.

Mode Number	Mode Description	Symmetry	PES(1)	RCCSD(T)(2)	Lee and Schaefer(3)	Lindh <i>et al.</i> (4)
1	HCCH asymm	<i>a</i> ₂	534	572	587	617
2	HCCH asymm	<i>b</i> ₂	594	580	513	696
3	HCCH symm opb	<i>b</i> ₁	775	751	757	770
4	HCCH symm ipb	<i>a</i> ₁	928	908	923	917
5	CHC ipb	<i>b</i> ₂	1258	1255	1279	1315
6	CC stretch	<i>a</i> ₁	1929	1932	2000	1939
7	CHC stretch	<i>a</i> ₁	2358	2352	2471	2385
8	HCCH asymm stretch	<i>b</i> ₂	3245	3265	3339	3304
9	HCCH symm stretch	<i>a</i> ₁	3355	3370	3443	3403

(1) Normal mode frequencies calculated from fitted PES.

(2) Frequencies calculated at the RCCSD(T) level of theory with aug-cc-pVTZ basis.

(3) Frequencies calculated using a DZ+P basis set at the SCF/CISD level of theory.

(4) Frequencies calculated using a TZ2Pf basis set at the MP2 level of theory.

C₂H₃⁺ Multimode Results

- Vibrational frequencies computed using Multimode (J=0).

Diffusion Monte Carlo (DMC) ZPE is $7389.3 \text{ cm}^{-1} \pm 2.1 \text{ cm}^{-1}$.

State	5MR	5MR-4MR	5MR-3MR
(000000000) Zero-Point Energy	7367.0	-35.5	-4.8
(010000000) <i>b</i> ₂	535.0	2.4	-11.1
(100000000) <i>a</i> ₂	562.1	3.3	-11.3
(001000000) <i>b</i> ₁	767.8	3.4	-10.3
(000100000) <i>a</i> ₁	859.3	4.3	-20.3
(020000000) <i>a</i> ₁	1067.1	2.3	-27.1
(000010000) <i>b</i> ₂	1134.2	2.3	-4.9
(000200000) <i>a</i> ₁	1874.7	4.7	-55.9
⋮	⋮	⋮	⋮
(010000100) <i>b</i> ₂	2923.6	1.1	-9.0
(000011000) <i>b</i> ₂	3042.4	1.4	9.0
(001000100) <i>b</i> ₁	3104.6	3.4	-11.1
(000000010) <i>b</i> ₂	3119.6 (3142.2)(1)	0.3	22.8
(000100100) <i>a</i> ₁	3201.8	6.5	1.7
(000000001) <i>a</i> ₁	3219.2	-4.2	6.7
(000010100) <i>b</i> ₂	3442.9	4.2	-3.1
(010000010) <i>a</i> ₁	3656.6	1.6	30.9

(1) Experimental data taken from Oka and co-workers

Methane: Multimode calculations

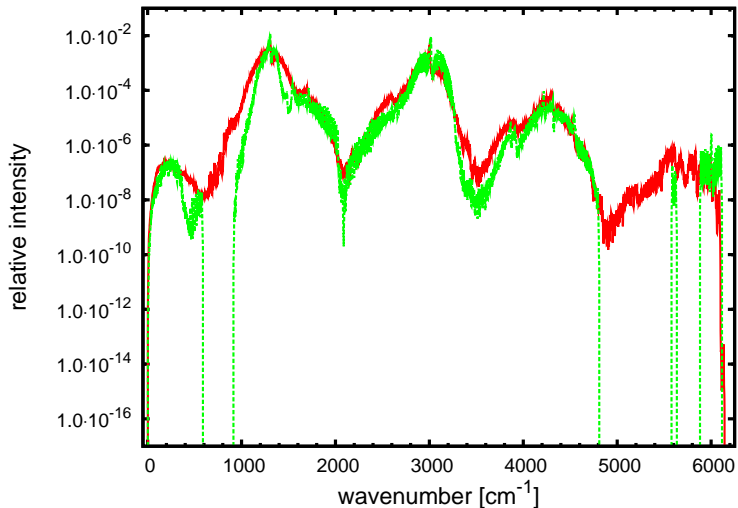
- Methane most widely studied molecule.
- Selected for first test calculations.
- Ro-vibrational states computed up to 6200 cm^{-1} ($J=34$).
- Calculations performed by Robert Warmbier.

$J, n_1 n_2 n_3 n_4$	exact	adiabatic	experiment(1)
1, 0000	10.40	9.92	10.48
1, 0001	1309.31, 1323.59	1318.45	1311.4, 1326.2
1, 0100	1531.96, 1532.09	1531.48	1544.0
1, 1000	2911.89	2911.27	-
1, 0010	3012.70, 3014.42	3012.87	3028.8, 3030.5
5, 0000	156.04	133.62	-
5, 0001	1433.79, 1484.90	1441.52	-
5, 0100	1678.69, 1679.91	1658.10	-
5, 1000	3057.82	3033.09	-
5, 0010	3155.29, 3161.13	3132.38	-
10, 0000	571.13	495.82	-
10, 0001	1820.51, 1912.41	1799.74	-
10, 0100	2096.97, 2100.87	2029.37	-
10, 1000	3473.43	3389.52	-
10, 0010	3565.92, 3576.88	3482.56	-

(1) References in: S. Carter and J.M. Bowman, J. Phys. Chem. A, 104 (11), 2355-2361, 2000.

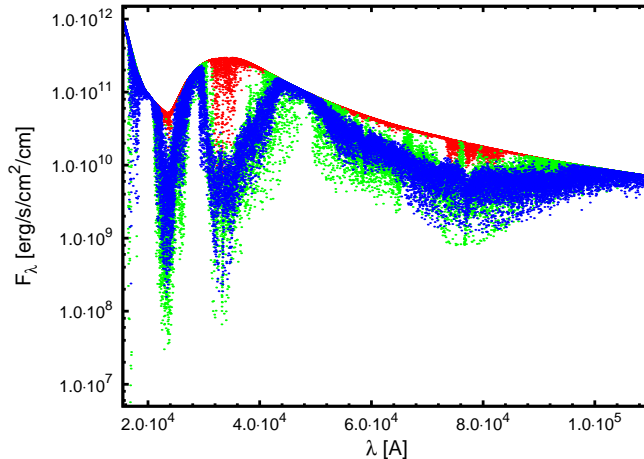
Spectroscopic data for CH4

Qualitative agreement with HITRAN database (green).



Radiation Transport

Radiation transport in model system using present dataset.
Comparison of radiative flux distribution with HITRAN(green) and GEISA(red)



Conclusion

- Method of calculation of spectroscopic data from astrophysical applications.
- Einstein coefficients and line energies calculated for Methane.
- Good agreement with existing database HITRAN (spontaneous emission coeff.).
- Fill-in missing regions.
- More elaborate tests for radiation transport (work in progress).
- Spectroscopic data can be generated for other molecules!!
- Acknowledgement :US Department of Energy, Office of Science (#DE-FG02-07ER54914).

Thank You